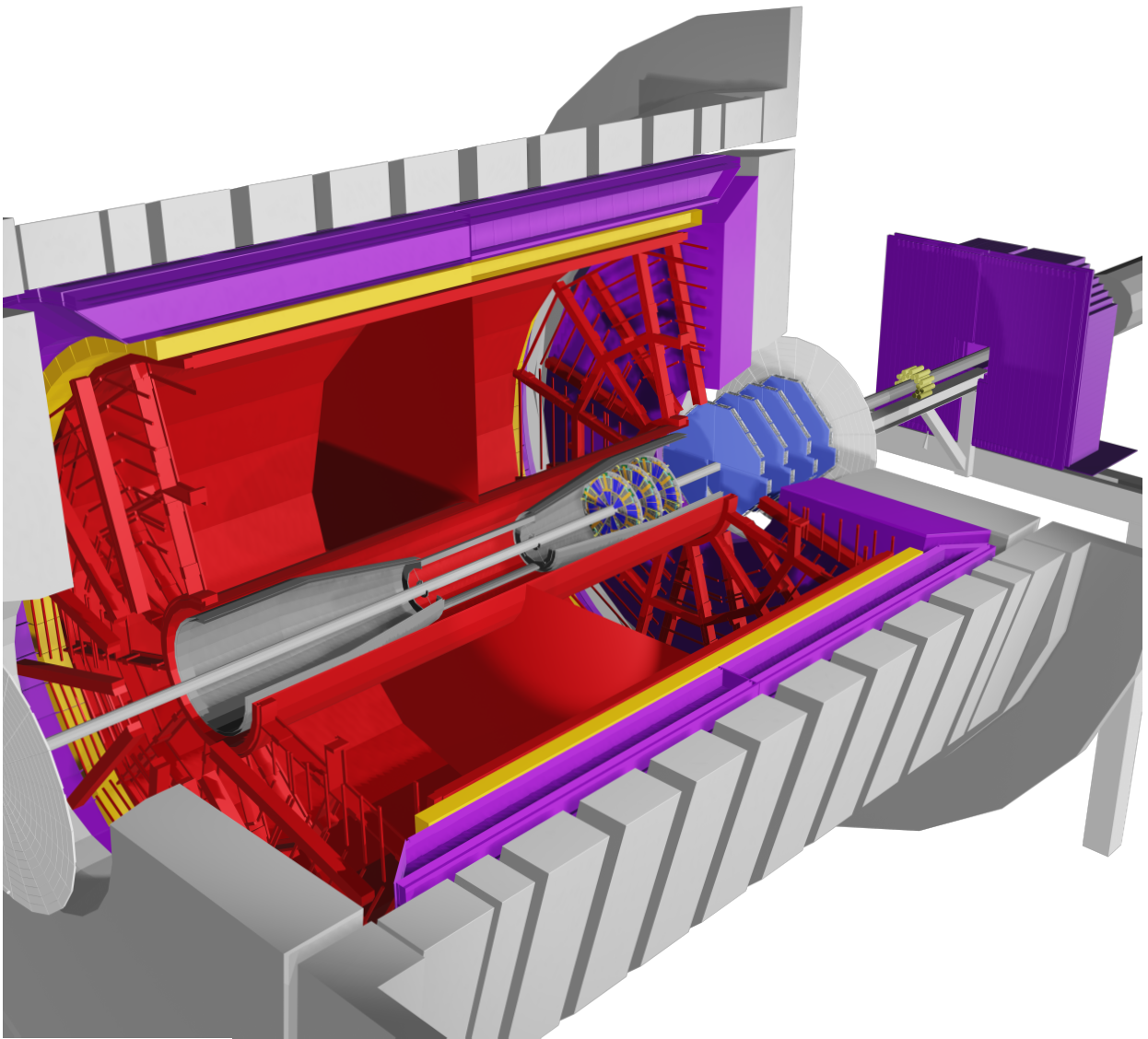


The STAR Beam Use Request for Run-21, Run-22 and data taking in 2023-25

The STAR Collaboration



1 Executive Summary

2 This Beam Use Request outlines the physics programs that compels the STAR collaboration
3 to request data taking during the years 2021-2025.

4 STAR's **highest scientific priorities** for Run-21 and Run-22 are to complete the NSAC-
5 endorsed second phase of the Beam Energy Scan (BES-II) program, and initiate the "must-
6 do" Cold QCD forward physics program enabled by the newly completed suite of forward
7 detectors via the collection of transversely polarized p+p data at 510 GeV. From 2023-25
8 we will use a combination of soft and hard probes to explore the microstructure of the QGP
9 and continue the forward physics program via the collection of high statistics Au+Au, p+Au
10 and p+p data at $\sqrt{s_{NN}} = 200$ GeV.

11 The BES-II program has so far been very successful. As shown in Table 1, we have
12 recorded collisions at $\sqrt{s_{NN}} = 9.2-27$ GeV in collider mode, and $\sqrt{s_{NN}} = 3-7.7$ GeV in fixed
13 target (FXT) mode. We expect to complete data collection at $\sqrt{s_{NN}} = 9.2$ GeV by the end
14 of Run-20b. In Run-21, as shown in Table 2, our number one priority is to complete the
15 BES-II by recording 100 M good events at $\sqrt{s_{NN}} = 7.7$ GeV.

Table 1: Summary of all BES-II and FXT Au+Au beam energies, equivalent chemical potential, event statistics, run times, and date collected.

Beam Energy (GeV/nucleon)	$\sqrt{s_{NN}}$ (GeV)	μ_B (MeV)	Run Time	Number Events Requested (Recorded)	Date Collected
13.5	27	156	24 days	(560 M)	Run-18
9.8	19.6	206	36 days	400 M (582 M)	Run-19
7.3	14.6	262	60 days	300 M (324 M)	Run-19
5.75	11.5	316	54 days	230 M (235 M)	Run-20
4.59	9.2	373	in progress	160 M ¹	Run20+20b
31.2	7.7 (FXT)	420	0.5+1.1 days	100 M (50 M+112 M)	Run-19+20
9.8	4.5 (FXT)	589	0.9 days	100 M (108 M)	Run-20
7.3	3.9 (FXT)	633	1.1 days	100 M (117 M)	Run-20
19.5	6.2 (FXT)	487	1.4 days	100 M (118 M)	Run-20
13.5	5.2 (FXT)	541	1.0 day	100 M (103 M)	Run-20
5.75	3.5 (FXT)	666	0.9 days	100 M (116 M)	Run-20
4.59	3.2 (FXT)	699	2.0 days	100 M (200 M)	Run-19
3.85	3.0 (FXT)	721	4.6 days	100 M (259 M)	Run-18
3.85	7.7	420	11-20 weeks	100 M	Run-21 ²

¹ Run-20b is still in progress at the time of submission of this BUR, we expect to reach our goals

² Data not yet collected, Run-21 forms part of this year's BUR.

16 Based on guidance from the Collider-Accelerator Department (C-AD) and past experience
17 we expect that the bulk of Run-21 will be devoted to Au+Au collisions at $\sqrt{s_{NN}} = 7.7$ GeV,

Table 2: Proposed Run-21 assuming 24 -28 cryo-weeks, including an initial one week of cool-down, one week for CeC ,and a one week set-up time for each collider energy and 0.5 days for each FXT energy.

Single-Beam Energy (GeV/nucleon)	$\sqrt{s_{NN}}$ (GeV)	Run Time	Species	Events (MinBias)	Priority
3.85	7.7	11-20 weeks	Au+Au	100 M	1
3.85	3 (FXT)	3 days	Au+Au	300 M	2
44.5	9.2 (FXT)	0.5 days	Au+Au	50 M	2
70	11.5 (FXT)	0.5 days	Au+Au	50 M	2
100	13.7 (FXT)	0.5 days	Au+Au	50 M	2
3.85	3 (FXT)	3 weeks	Au+Au	2 B	3
8.35	17.1	2.5 weeks	Au+Au	250 M	3
100	200	1 week	O+O	400 M 200 M (central)	3

18 the lowest collider energy of the program. Collection of these events is our highest priority.
19 However, if we assume optimistic, but not overly so, rates and up-times, and 28 cryo-weeks,
20 we project that the opportunity to collect of other exciting datasets will arise.

21 The *second highest priority* for Run-21 identified by the STAR collaboration is four short
22 FXT runs; the collection of 300 M good events at $\sqrt{s_{NN}} = 3$ GeV and 50 M good events at
23 each of three higher beam energies ($\sqrt{s_{NN}} = 9.2, 11.5,$ and 13.7 GeV). 300 M events at 3
24 GeV with the enhanced iTPC and eTOF coverage gives access to the proton higher moments,
25 precision ϕ , hypernuclei and dilepton measurements. The higher $\sqrt{s_{NN}}$ FXT data combined
26 with the collider data at the same energy will provide full proton rapidity coverage allowing
27 us to probe in detail the mechanisms of stopping at play in heavy-ion collisions. We estimate
28 the total run time required to collect all these datasets is 6 days.

29 The STAR collaboration also finds important scientific opportunities are presented by
30 the collection of our *third highest priority* datasets:

- 31 • 2 B good events at $\sqrt{s_{NN}} = 3$ GeV in FXT mode. These enhanced statistics make
32 possible the measurements of mid-rapidity proton 5-th/6-th order moments/cumulants,
33 the system size dependence of ϕ meson production and a potential discovery of the
34 double- Λ hypernuclei.
- 35 • A sixth collider beam energy at $\sqrt{s_{NN}} = 17.1$ GeV. These data will provide for a finer
36 scan in a range where the energy dependence of the net-proton kurtosis and neutron
37 density fluctuations appear to undergo a sudden change.
- 38 • O+O data at $\sqrt{s_{NN}} = 200$ GeV, in the context of understanding the early-time condi-
39 tions of small systems. These data would allow for a direct comparison with a similarly
40 proposed higher-energy O+O run at the LHC, and further motivate the case for a small

41 system scan complementary to ongoing efforts by the NA61/SHINE collaboration at
 42 SPS energies, and other proposed light-ion species at the LHC.

43 The sequence with which we collect these datasets is currently somewhat fluid; we do
 44 not want to take partial datasets. We expect to refine the ordering of our goals as Run-21
 45 progresses. Collection of these data during future RHIC running periods is also of interest
 46 to the collaboration.

47 For Run-22, as shown in Table 3, we propose **a dedicated 20 cryo-week transversely**
 48 **polarized p+p run at $\sqrt{s} = 510$ GeV**. This run will take full advantage of STAR’s new
 49 forward detection capabilities, consisting of a Forward Calorimeter System (FCS) and a
 50 Forward Tracking System (FTS) located between $2.5 < \eta < 4$, and further capitalizes on the
 51 recent BES-II detector upgrades.

52 These data will enable STAR to explore, with unprecedented precision, forward jet
 53 physics that probe both the high-x (largely valence quark) and low-x (primarily gluon)
 54 partonic regimes.

Table 3: Proposed Run-22 assuming 20 cryo-weeks, including an initial one week of cool-down and a two weeks set-up time.

\sqrt{s} (GeV)	Species	Polarization	Run Time	Sampled Luminosity	Priority
510	p+p	Transverse	16 weeks	400 pb ⁻¹	1

55 Looking further out, the STAR collaboration has determined that there is a compelling
 56 scientific program enabled by the first opportunity to capitalize on the combination of the
 57 BES-II and Forward Upgrades in the data collected from Au+Au, p+Au, and p+p collisions
 58 at $\sqrt{s_{NN}} = 200$ GeV as outlined in Table 4.

59 Significantly increased luminosities, the extended acceptance at mid-rapidity due to the
 60 iTPC, improved event plane and triggering capabilities of the EPD, and the ability to probe
 61 the previously inaccessible forward region are all exploited in our Hot QCD program, that
 62 informs on the microstructure of the QGP, and our Cold QCD program that will utilize
 63 transverse polarization that sets the stage for related future measurements at the EIC.

Table 4: Proposed Run-23 - Run-25 assuming 24 (28) cryo-weeks of running every year, and 6 weeks set-up time to switch species in 2024. Sampled luminosities assume a "take all" triggers.

$\sqrt{s_{\text{NN}}}$ (GeV)	Species	Number Events/ Sampled Luminosity	Date
200	Au+Au	10B / 38 nb ⁻¹	2023
200	p+p	235 pb ⁻¹	2024
200	p+Au	1.3 pb ⁻¹	2024
200	Au+Au	10B / 52 nb ⁻¹	2025

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1 Highlights from the STAR Program

1.1 Highlights from the Heavy Ion Program

1.1.1 Heavy-Ion Jet Measurements

Jet quenching, the modification of parton showers due to interactions in the QGP, is manifest in several distinct ways: energy transport to large angles, observable via jet energy loss and large-angle energy recovery; multiple-soft and single-hard coherent scatterings off of plasma constituents, observable via jet deflection or acoplanarity; and the modification of jet substructure. This broad spectrum of phenomena provides unique and incisive probes of the microscopic structure of the QGP. It also provides a robust experimental program, in which different observables with different systematic sensitivity probe the same underlying physics, providing stronger constraints on theoretical models of jet quenching than single measurements. STAR has a comprehensive jet quenching program which covers the full spectrum of these phenomena, using hadrons, direct photons, and reconstructed jets as probes.

STAR has led the development of essential analysis techniques for the challenging task of measuring reconstructed jet observables in heavy-ion collisions at RHIC. These include a data-driven Mixed-Event technique to measure uncorrelated jet background for semi-inclusive observables [1], enabling unbiased jet measurements over a broad phase space in heavy-ion collisions, notably low jet p_T (~ 10 GeV/ c) and large jet resolution parameter ($R \sim 0.5$); and sub-jet observables that are robust to the underlying event and yet sensitive to the jet splitting kinematics, applying them as a tool to access the resolution scale in jet-medium interactions [2].

Jet quenching measurements have traditionally utilized p+p collisions to provide an unmodified reference, and p +A collisions to measure initial state effects that may mask signals of quenching in the final state. More recently it has become evident that small systems themselves exhibit QGP-like flow signatures for event selection corresponding to high Event Activity (EA), and an urgent question in the field is whether evidence can likewise be found for jet quenching in such systems. The STAR jet quenching program therefore includes measurements in (unpolarized) p+p and p +Au collisions, as well as Au+Au collisions.

In this section we present recent highlights of the STAR jet quenching program. The STAR papers published in this area in the past year can be found in Refs. [3–7].

Inclusive and semi-inclusive jet yield suppression: Inclusive jet yield suppression is a hallmark of jet quenching in heavy-ion collisions. STAR has recently reported the first measurement at RHIC of inclusive charged-particle jet distributions in central and peripheral Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV [7], together with measurements of their yield suppression, R_{AA} (normalized by the yield in p+p collisions calculated by PYTHIA tuned to other STAR data [8]) and R_{CP} . Figure 1 shows the extracted R_{CP} compared to a similar measurement by ALICE, and to charged-hadron R_{CP} measured at both colliders. A striking similarity is seen between the two inclusive jet measurements, and between the two

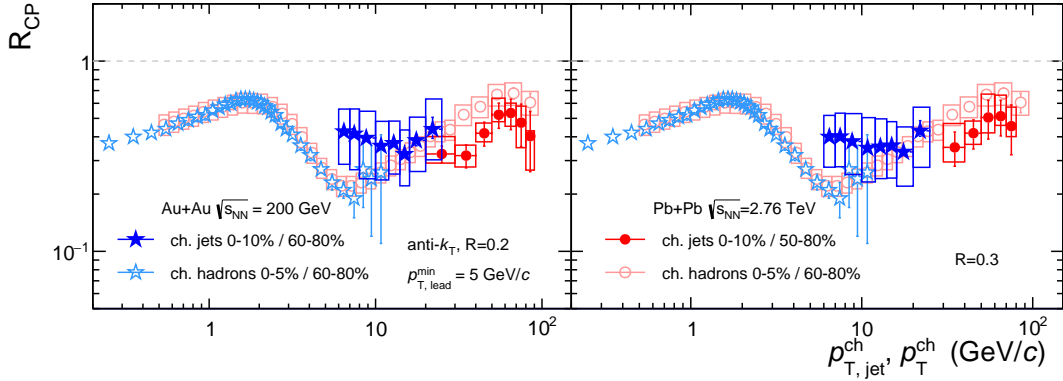


Figure 1: Measurement of R_{CP} as a function of $p_{T,jet}^{ch}$ for charged-particle jets (anti- k_T , $R = 0.2$ and 0.3) measured by STAR (blue points) [7], compared to charged-jet R_{CP} at the LHC and to inclusive hadron R_{CP} at RHIC and the LHC. Note the different centrality selections.

153 inclusive hadron measurements. The p_T -dependence of R_{CP} is stronger for hadrons in the
 154 region of overlap. While there remain differences in centrality selection between the datasets,
 155 this is the most direct comparison to date of reconstructed jet measurements at RHIC and
 156 the LHC.

157 This paper also reported the ratio of jet yields in central and peripheral Au+Au collisions
 158 for $R = 0.2$ and 0.4 , which is a probe of jet shape and its in-medium modification. Consis-
 159 tency to theoretical calculations is found within uncertainties. However, the calculations
 160 exhibit significant spread in the jet shape ratio, presenting an opportunity for more precise
 161 measurements to discriminate between them. A measurement of the inclusive jet yield in
 162 Au+Au collisions including both charged and neutral particle constituents using the much
 163 larger data set recorded in 2014, corresponding to 9.9 nb^{-1} [9] is underway. STAR also has
 164 full jet measurements in p+p collisions for use for the R_{AA} normalization.

165 A recent STAR measurement, likewise using the 9.9 nb^{-1} 2014 dataset, extends the semi-
 166 inclusive measurement of charged jets (anti- k_T , $R = 0.2$ and 0.5) recoiling from a high- E_T
 167 photon trigger to photon triggers in the range $15 < E_T^{trig} < 20 \text{ GeV}$ [10]. Currently, the
 168 recoil jet yield suppression for 0–15% Au+Au collisions (I_{AA}) is determined by comparison
 169 to the yield in p+p collisions calculated using PYTHIA-6 (STAR tune [8]) and PYTHIA-8.
 170 Significant yield suppression in central Au+Au collisions is observed for $R = 0.2$, with less
 171 suppression for $R = 0.5$. Theoretical calculations predict a stronger dependence of I_{AA} on
 172 $p_{T,jet}^{ch}$ for $R = 0.5$ than observed. A measurement of this observable in p+p collisions is in
 173 progress, to provide a data reference rather than PYTHIA calculations for I_{AA} .

174 Jet yield suppression is an indirect measurement of energy loss, because it convolutes
 175 out-of-cone energy loss with the shape of the jet spectrum – a fixed energy loss generates
 176 greater suppression for a steeper spectrum. Since the jet spectrum shape depends strongly on
 177 the choice of observable (inclusive, semi-inclusive) and collision energy, direct comparison of
 178 different jet quenching measurements requires this effect to be taken into account. Figure 2

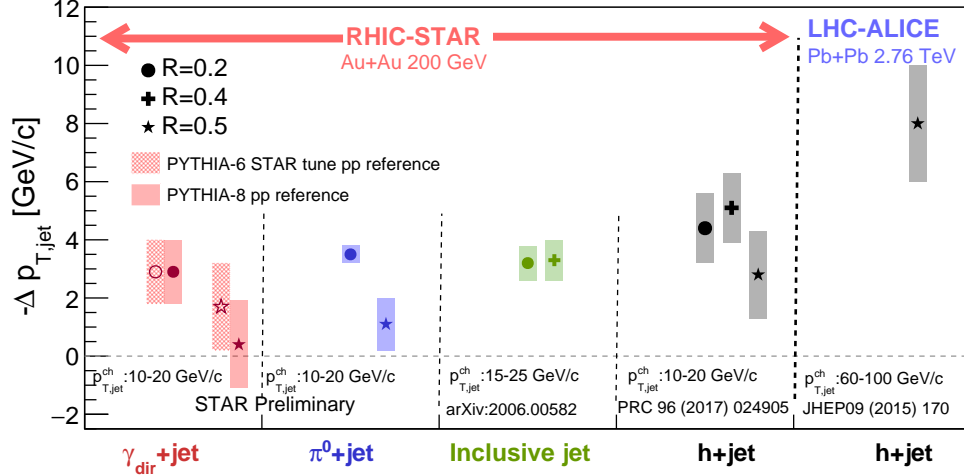


Figure 2: Out-of-cone jet energy loss derived from jet yield suppression measurements in A+A collisions (see text) for $\gamma_{\text{dir}}+\text{jet}$, $\pi^0+\text{jet}$, inclusive jet, and h+jet measurements at RHIC, and h+jet measurements at the LHC [9, 10]. Note the different $p_{T,\text{jet}}$ ranges.

179 shows the $p_{T,\text{jet}}$ shift needed between jet spectra measured in a reference system (p+p or
 180 peripheral A+A collisions) and in central A+A collisions, for several jet yield suppression
 181 measurements at RHIC and the LHC [9, 10]. The absolute magnitude of medium-induced
 182 jet energy loss is similar for several different observables at RHIC, and is smaller than the
 183 LHC measurement. Note that the $p_{T,\text{jet}}^{\text{ch}}$ range is significantly higher for the LHC h+jet
 184 measurement, so that the *relative* energy loss is smaller than at RHIC.

185 This is a first look at comparing medium-induced out-of-cone radiation at RHIC and the
 186 LHC. Clearly, as the measured $p_{T,\text{jet}}$ range at RHIC moves up and that at the LHC moves
 187 down in upcoming measurements, more precise comparisons can be made. Nevertheless,
 188 Fig. 2 already provides significant constraints on jet quenching calculations that seek to
 189 model RHIC and LHC measurements in a unified way.

190 **Jet-structure modifications:** The Fragmentation Function (FF), normalized per jet,
 191 provides information of the longitudinal momentum fraction ($z = p_{T,\text{trk}} \cos(\Delta r)/p_{T,\text{jet}}$) of
 192 charged particles projected along the jet axis. While FF have been measured previously at
 193 the LHC [11, 12], STAR has utilized the semi-inclusive approach to measure the FF of charged
 194 jets for the first time at RHIC [13]. The Mixed-Event approach developed in [1] is extended
 195 for the FF measurement, and utilized for the correction of uncorrelated jet contributions.
 196 The fully corrected FF are shown in Fig. 3 for jets of varying $p_{T,\text{jet}}^{\text{ch}}$ for mid-peripheral 40-
 197 60% collisions compared to PYTHIA-8 [14] predictions shown by the dashed curved. The
 198 FF shape in data is reproduced by PYTHIA-8 in these peripheral collisions. Measurements
 199 are ongoing to extend to central collisions where one expects a larger path length for the
 200 recoil jet and enhanced medium effects.

201 Another observable of the jet transverse profile is the differential jet shape, measured in

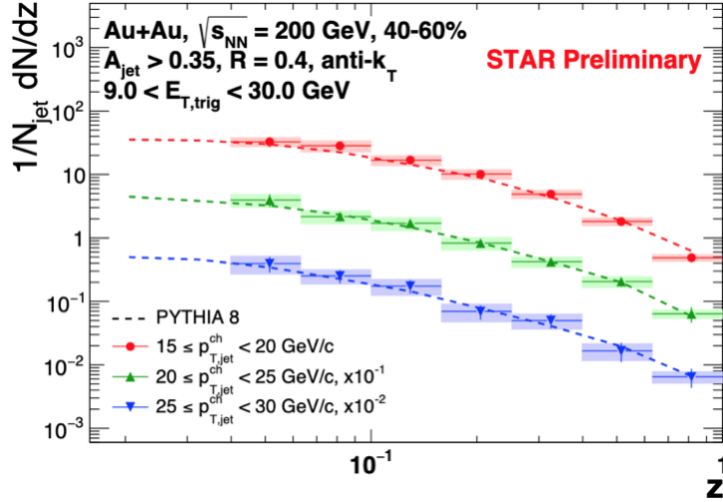


Figure 3: Fragmentation functions for recoil charged-particle jets of varying $p_{\text{jet},T}^{\text{ch}}$ with trigger $9.0 < E_T < 30.0$ GeV in 40–60% peripheral events compared to PYTHIA-8 simulations in the dashed curves.

202 radial annuli around the jet axis ($\rho(\Delta r)$). Utilizing the hard-core jet selection [15] which
 203 provides a pure sample of hard-scattered jets with a high constituent threshold, the fully
 204 corrected ρ as a function of Δr (distance between the constituent tracks and the jet axis)
 205 of leading jet with $20 < p_{T,\text{jet}} < 40$ GeV/ c for central (0–10%) and mid-central (20–50%)
 206 events are calculated. To probe possible in-medium modification of the jet structure and
 207 its dependence on the path length in medium, this observable is also differentially measured
 208 based on the jet’s orientation with respect to the event plane for 20–50% mid-central colli-
 209 sions, as shown in Fig 4. High- $p_{T,\text{trk}}$ particles are found closer to the jet core, whilst softer
 210 constituents are more evenly distributed around the jet. In comparing the soft particle pro-
 211 duction for in-plane vs. out-of-plane jets one finds subtle hints of path-length dependence.

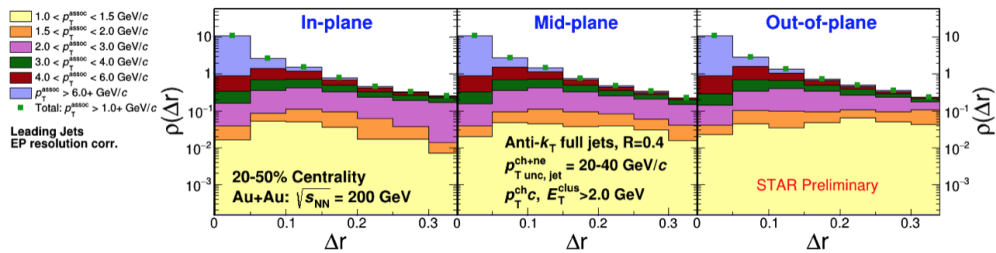


Figure 4: Differential measurement of the leading jet ($20 < p_{T,\text{jet}} < 40$ GeV/ c) shapes in 20–50% central Au+Au collisions shown for different jet azimuthal angles with respect to the event-plane angle. The $p_{T,\text{trk}}$ -dependence of the associated tracks are shown in the different stacked histograms. Results are corrected for event-plane resolution effects.

213 **Jets in p+Au collisions:** STAR has searched for jet-medium interactions in p+Au colli-
 214 sions by looking at potential modifications of semi-inclusive charged-particle jet yields and
 215 jet substructure observables such as the jet mass and SoftDrop groomed jet mass. p+Au col-
 216 lisions are classified as low or high event-activity (EA) according to the particle multiplicity
 217 in the Au-going direction as measured by the BBC-East detectors.

218 The charged-particle jet spectra, normalized per HT trigger (uncorrected for detector
 219 effects) are shown on the left of Fig. 5 where the open (full) markers correspond to low (high)
 EA. The different colored markers represent the azimuthal separation between the trigger and

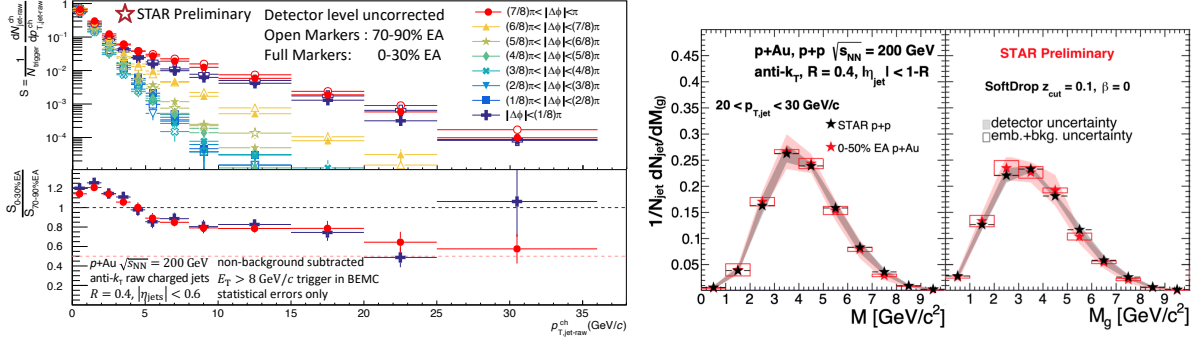


Figure 5: Left: Semi-inclusive charged jet spectra in p+Au collisions for high and low event-activity (EA) events, the ratio is shown in the bottom panel. Right: Fully corrected (groomed) jet-mass distributions in p+Au with high EA and p+p collisions.

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221 the recoil jet. We see for jet with $p_{T,jet-raw} > 10$ GeV/c, a significant suppression in high to
 222 low EA events for both the trigger-side and recoil-side spectra. These suppression ratios are
 223 qualitatively different from jet suppression in Au+Au collisions, where the recoil jets traverse
 224 more QGP on average and are suppressed compared to the trigger-side. In investigating if
 225 this suppression is a result of modification of jet structure, STAR also measured the fully
 226 corrected jet-mass and groomed jet-mass distributions, normalized per jet, on the right of
 227 Fig. 5. The distribution in high EA p+Au collisions is comparable to that in p+p collisions
 228 within the systematic uncertainties, and this leads to the conclusion that CNM effects do not
 229 significantly affect the jet substructure. The jet mass measurements in p+Au will be followed
 230 in a more differential fashion by studying finer EA classes and rapidity selections which can
 231 isolate jets originating from the Au vs p side. Both of these measurements from STAR point
 232 to lack of jet modification from nuclear effects and also to a more fundamental selection bias
 233 when identifying classes of high vs low activity events in asymmetric collisions.

234 1.1.2 Bulk Correlations

235 Over the past years, the STAR collaboration has performed a series of correlation measure-
 236 ments directed towards a comprehensive understanding of the QCD phase diagram and the
 237 bulk properties of the QGP phase. Here we highlight the most recent STAR results on bulk
 238 correlations, which are expected to shed light on the QCD phase diagram as well as on the
 239 transport properties of the QGP.

240 **Net-proton number fluctuations and the QCD critical point:** One of the main
 241 goals of the STAR Beam Energy Scan (BES) program is to search for possible signatures
 242 of the QCD critical point (CP) by scanning the temperature (T) and the baryonic chemical
 243 potential (μ_B) plane by varying the collision energy. When the system produced in the heavy
 244 ion collisions approaches the CP, the correlation length diverges. Higher order cumulants
 245 of conserved net-particle multiplicity distributions are sensitive to such correlation lengths
 246 as the divergence of correlation length leads to enhanced fluctuations in the net-particle
 multiplicity distributions.

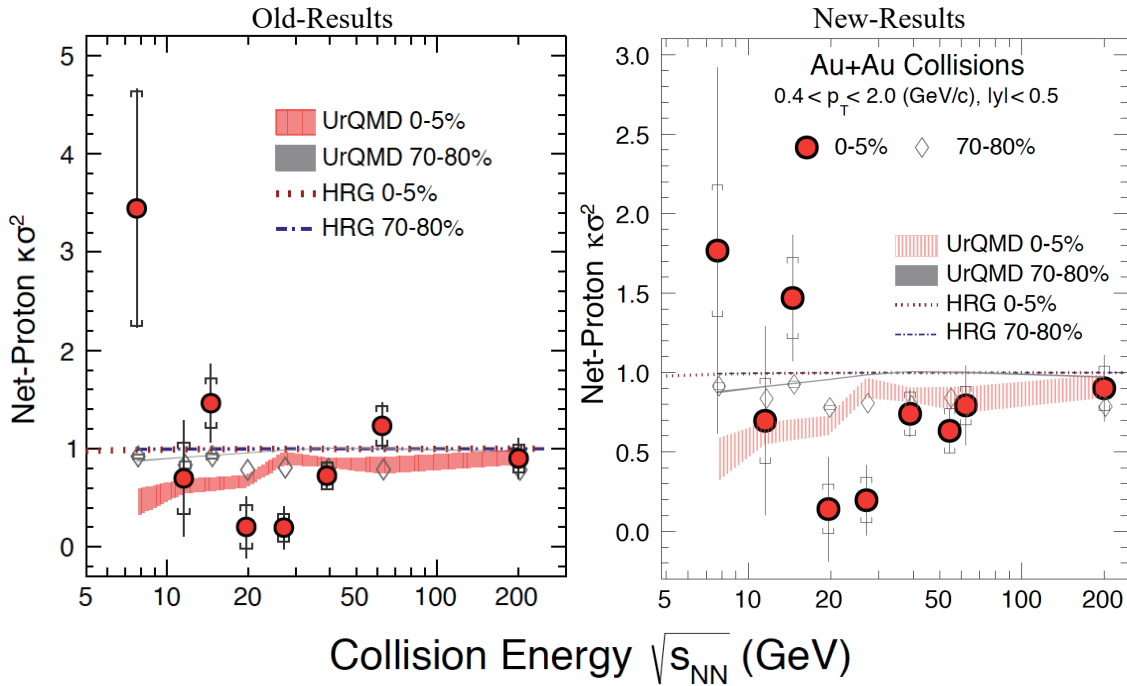


Figure 6: $\kappa\sigma^2$ as a function of collision energy for net-proton distributions measured in central (0-5%) and peripheral (70-80%) Au+Au collisions within $0.4 < p_T$ (GeV/c) < 2.0 and $|y| < 0.5$. The error bars and caps show statistical and systematic uncertainties, respectively. The dashed and dash-dotted lines correspond to results from a hadron resonance gas (HRG) model. The shaded bands are the results of a transport model calculation (UrQMD). The model calculations utilize the experimental acceptance and incorporate conservation laws for strong interactions, but do not include the dynamics of phase transition or critical point. **The new results are obtained after removing the spoiled events, the largest changes are seen in central Au+Au collisions at 7.7 and 62.4 GeV.**

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The ratios of the cumulants of identified net-particle multiplicity distributions, such as net-protons, have been predicted to be ideal observables sensitive to the onset of the QCD

250 phase transition and the location of the CP. A non-monotonic variation of these ratio of
 251 cumulants, such as C_4/C_2 ($=\kappa\sigma^2$), as a function of collision energy has been proposed to be
 252 an experimental signature of the CP. Taking the ratios of cumulants has advantages as it
 253 cancels the volume fluctuations to first order. Further, these ratios of cumulants are related
 254 to the ratio of baryon-number susceptibilities at a given T and μ_B . Near the critical point,
 255 QCD-based calculations predict the net-baryon number distributions to be non-Gaussian
 256 and susceptibilities to diverge, causing these ratios to have non-monotonic variation as a
 257 function of collision energy. However, the finite-size and finite-time effects in heavy-ion
 258 collisions limit the growth of correlation length, and hence it could restrict the values of $\kappa\sigma^2$
 259 from its divergence as a function of collision energy.

260 Figure 6 shows the collision energy variation of net-proton $\kappa\sigma^2$ for central and peripheral
 261 Au+Au collisions within the acceptance of $0.4 < p_T < 2.0$ GeV and $|y| < 0.5$. In central
 262 collisions, a non-monotonic variation with beam energy is observed for $\kappa\sigma^2$ with a significance
 263 of 3.0σ . In contrast, monotonic behavior with beam energy is observed for the statistical
 264 hadron gas (HRG) model, and for a nuclear transport UrQMD model without a critical
 265 point, and experimentally in peripheral collisions.

266 High statistics data from the ongoing BES-II program can provide precision measure-
 267 ments at higher μ_B region in the QCD phase diagram. In addition, due to the iTPC [16]
 268 and eTOF [17] upgrades, a differential measurement in $|y| < 1.5$ and $p_T > 0.15$ GeV/c will be
 269 explored. The study of acceptance dependence of net-proton $\kappa\sigma^2$ and other cumulants ra-
 270 tios are important to understand critical fluctuation. Furthermore, the forward Event-Plan
 271 Detector (EPDs) [18] can also be used to determine the centrality selection in heavy-ion
 272 collisions for this measurement.

273 **Global polarization measurements at 27 GeV:** In heavy-ion collisions, many theoret-
 274 ical models propose that the large angular momentum in the collisions of two nuclei [19–21]
 275 can be transferred to the microscopic constituent of the created matter. Consequently, the
 276 spin of the produced quarks and gluons might be polarized along the direction of the global
 277 angular momentum due to spin-orbit coupling. The direction of the global angular momen-
 278 tum is perpendicular to the reaction plane, as defined by the incoming beam and the impact
 279 parameter vector. This direction can be determined from directed flow measurements of the
 280 spectators. STAR observed significant non-zero polarization of hyperons [20] with increasing
 281 strength with decreasing collision energy (from 200 to 7.7 GeV).

282 We recently report more differential measurements using our newly installed EPDs in
 283 Au+Au collisions at 27 GeV as functions of the hyperon’s transverse momentum, and pseudo-
 284 rapidity. In Fig. 7 left panel we observe that the polarization does not show a strong
 285 dependence on p_T , albeit large uncertainties. There are several expectations on the p_T
 286 dependence on the polarization. If global polarization is generated by the vorticity of the
 287 initial state that does not have a strong p_T dependence then the result is compatible with
 288 expectations. Alternatively, at lower p_T , due to the smearing effect caused by scattering at
 289 later stages of the collisions, we might expect a decrease of the polarization. In addition,
 290 one might expect a decrease in the polarization at higher p_T due to the expected larger

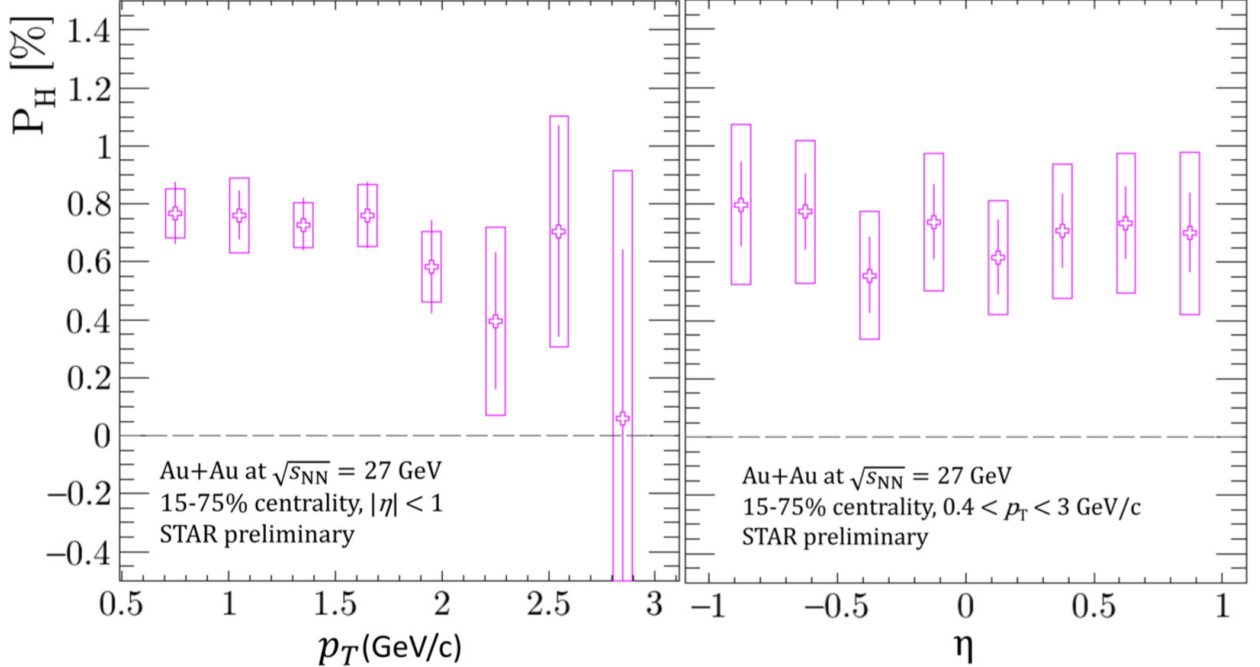


Figure 7: The global polarization measurements as a function of p_T and η in 15-75% central Au+Au collisions at $\sqrt{s_{NN}} = 27$ GeV. The figure is taken from Ref [22].

291 contribution from jet fragmentation. Fig 7 right panel shows the pseudo-rapidity dependence
 292 of the polarization measurement, no η -dependence of the polarization is observed within
 293 uncertainties. The vorticity is expected to decrease at larger rapidity, but might also have a
 294 local minimum at $\eta=0$ due to complex shear flow structure [21, 23, 24] however, this might
 295 be difficult to observe within STAR's acceptance. This preliminary observation of no p_T or η
 296 dependence of the polarization is consistent with our previous measurements at 200 GeV [19].
 297 STAR plans to perform the same measurement with an extended pseudo-rapidity coverage
 298 using the iTPC detector upgrade and with higher statistics BES-II data set enabling higher
 299 a precision result.

300 **Global spin alignment of K^{*0} and ϕ :** Unlike the self-analyzing (anti) Λ , the polarization
 301 of vector mesons such as $\phi(1020)$ and $K^{*0}(892)$ cannot be directly measured since vector
 302 mesons mainly decay through the strong interaction in which parity is conserved. The spin
 303 alignment of vector mesons can be given by a 3×3 spin density matrix with unit trace [25].
 304 The spin density matrix diagonal elements ρ_{nn} , $n=0,1$ and -1 , represent the probabilities
 305 for the spin component along the quantization axes. When there is no spin alignment this
 306 means that all three spin states (ρ_{nn}) have equal probability to be occupied meaning $\rho_{nn} =$
 307 $1/3$. Out of the three diagonal elements, only the $n=0$ case is independent of the other two.
 308 Consequently, it is intriguing to experimentally investigate the ρ_{00} of vector mesons.

309 Figure 8 shows the centrality dependence of ρ_{00} for both vector meson species for Au
 310 + Au collisions at 200 GeV. The ϕ -meson results are presented for transverse momentum

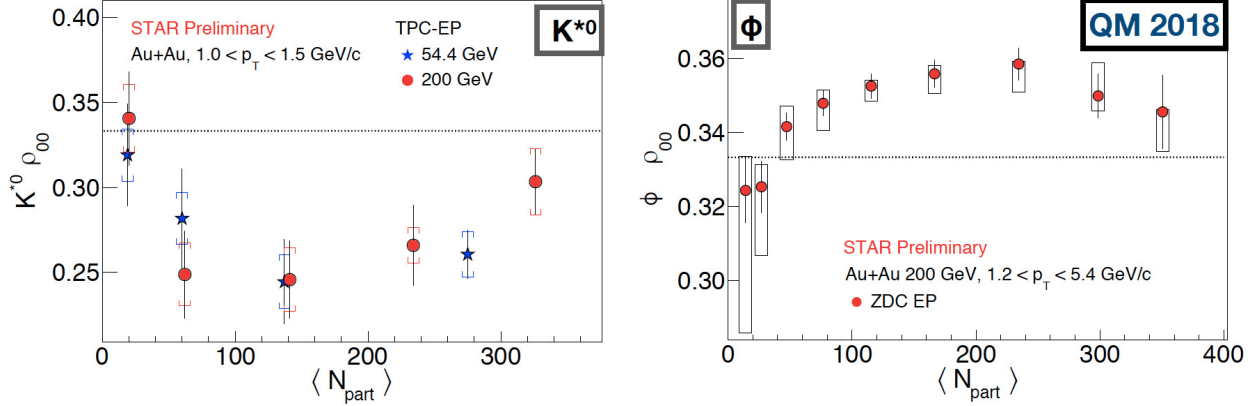


Figure 8: The spin alignment ρ_{00} measurements of vector mesons K^{*0} and ϕ as a function of N_{part} for the indicated p_T range of the Au+Au collisions at 200 and 54.4 GeV. The figure is taken from Ref [26].

311 $1.2 < p_T < 5.4$ GeV/c, and ρ_{00} for this species is significantly above 1/3 for mid-central
 312 collisions, indicating finite global spin alignment. The K^{*0} -meson results are presented for
 313 transverse momentum $1.0 < p_T < 1.5$ GeV/c, and the magnitude of ρ_{00} for this particle
 314 species is observed to be significantly less than 1/3 for mid-central collisions.

315 The distinction between the global spin alignment for K^{*0} and ϕ may be assigned to
 316 different in-medium interactions due to the difference in the lifetime (ϕ -meson is 10 times
 317 larger than K^{*0} -mesons), and/or a different response to the vector meson field. These global
 318 spin alignment results are expected to shed light on the possible vector meson fields [27, 28].
 319 Such investigations are extremely important since vector meson fields are a crucial part of
 320 the nuclear force that binds nucleons to atomic nuclei and are also central in describing
 321 properties of nuclear structure and nuclear matter.

322 **Nuclear deformation measurements:** Deformation is a fundamental property of atomic
 323 nuclei that reflects the correlated nature of the dynamics of nucleons within the quantum
 324 many-body system. The majority of atomic nuclei possess an intrinsic deformation, most of
 325 which is an axial quadrupole, or ellipsoidal, deformation.

326 Prior relativistic heavy-ion collision measurements from STAR reported strong signatures
 327 of nuclear deformation using detailed comparisons between Au+Au collisions and U+U col-
 328 lisions [29]. These measurements suggest that U+U collisions being much more deformed
 329 in their ground state. Consequently, we can say that these detailed comparisons between
 330 Au+Au and U+U collisions enabled us to examine the geometry of the colliding nuclei.

331 The study of mean transverse momentum dependence of the elliptic and triangular flow
 332 harmonics in Au+Au and U+U collisions are recently proposed by theory calculations [31]
 333 that are more sensitive to the deformation of the colliding nuclei.

334 Figure 9 shows the scaled mean p_T dependence of v_2 and v_3 for the central Au+Au and
 335 U+U collisions. STAR preliminary data show a clear positive correlation for v_2 and v_3 in
 336 Au+Au collisions that is in agreement with the v_3 from U+U collisions. In contrast, a
 337 nontrivial negative correlation is observed in v_2 as a function of scaled mean p_T in U+U

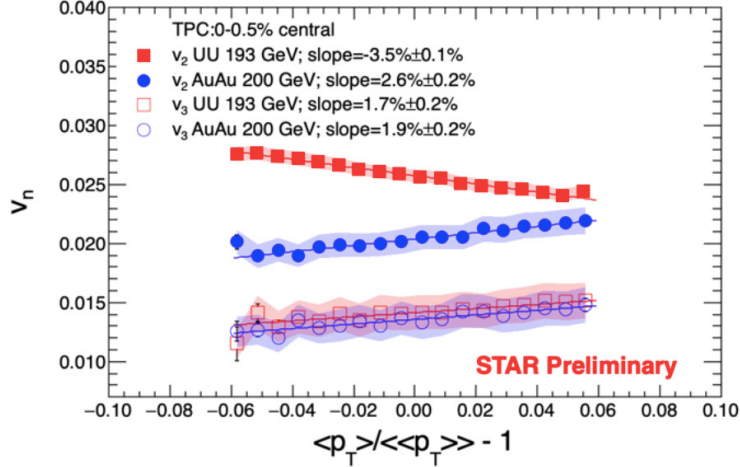


Figure 9: The scaled mean p_T dependence of the elliptic and triangular flow harmonics for 0-0.5% central Au+Au and U+U collisions. The figure is taken from Ref [30].

338 collisions. Also these preliminary results are consistent with the theoretical expectation for
 339 a deformed U nuclei [31].

340 **Flow correlations and fluctuations measurements:** Flow harmonics (v_n) calculated
 341 from the Fourier expansion of the particle azimuthal distributions are commonly employed
 342 to quantify the azimuthal anisotropy of particle emission relative to the collision symmetry
 343 planes. While the lower-order Fourier coefficients (v_2 and v_3) are more directly related to the
 344 corresponding eccentricities of the initial state, the higher-order flow harmonics ($v_{n>3}$) can
 345 be induced by a non-linear (mode-coupled) response to its lower-order harmonics and also
 346 with a linear response to the same-order anisotropy. These higher-order flow harmonics and
 347 their linear and mode-coupled contributions can be used to constrain the initial conditions
 348 and the transport properties of the medium in the theoretical calculations.

349 The v_2 and v_3 harmonics are sensitive to the respective influence of the initial-state
 350 eccentricity and the final-state viscous attenuation, which have proven difficult to disentangle.
 351 The mode-coupled coefficients show characteristically different dependencies on the viscous
 352 attenuation and the initial-state eccentricity [32]. Therefore, they can be used in conjunction
 353 with measurements for the v_2 and v_3 harmonics to leverage additional unique constraints for
 354 initial-state models, as well as reliable extraction of transport coefficient.

355 Figure 10 shows the mode-coupled response coefficients, $\chi_{4,22}$ and $\chi_{5,23}$, with a weak
 356 centrality dependence, akin to the patterns observed for similar measurements at the LHC
 357 for Pb+Pb collisions at 2.76 TeV [34]. These patterns suggest that the mode-coupled response
 358 coefficients are dominated by initial-state eccentricity couplings which is known from models
 359 to have a weak dependence on beam energy. The correlations of the event plane angles, $\rho_{4,22}$
 360 and $\rho_{5,23}$ show a strong centrality dependence that agrees well with the LHC measurements
 361 for Pb+Pb collisions at 2.76 TeV. The predictions from viscous hydrodynamic models [35,
 362 36] give a good qualitative description of the mode-coupled response coefficients and the

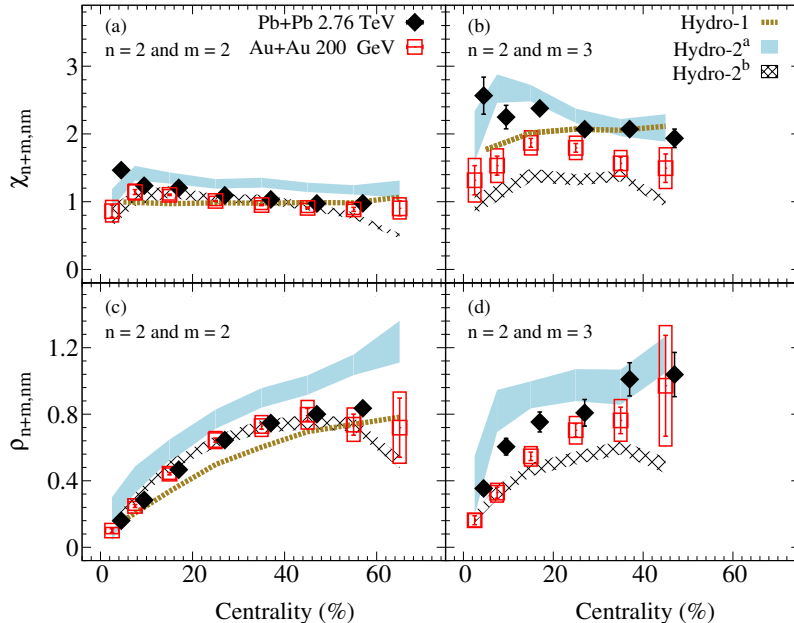


Figure 10: Results as a function of centrality for Au+Au collisions at 200 GeV [33]. Panels (a) and (b) shows the mode-coupled response coefficients, and panels (c) and (d) show the correlations of event plane angles. The closed-symbols represents similar LHC measurements [34]. The shaded bands indicate hydrodynamic model predictions Hydro-1 [35], Hydro-2^a and Hydro-2^b [36].

363 correlation of event plane angles.

364 **Small system measurements:** The comparisons of theoretical models to the flow har-
 365 monics, v_n , continue to be an essential avenue to evaluate the transport properties of partonic
 366 matter produced in large to moderate-sized collision systems [37–39]. For the small collision-
 367 systems formed in p/d/³He+Au and p+Pb collisions, collective flow might not develop due
 368 to the presence of large gradients in the energy-momentum tensor that could trigger non-
 369 hydrodynamic modes [40, 41]. Certainly, the most important question that divided our field
 370 is whether an alternative initial-state-driven mechanism [42] dominates over hydrodynamic
 371 expansion for these collision systems.

372 Current measurements for p/d/³He+Au collisions, which supplement earlier measure-
 373 ments at both RHIC [43] and the LHC [44] aim to address the respective influence of collision-
 374 system size and its subnucleonic fluctuations, and viscous attenuation on the measured v_n .

375 Figure 11 shows the $v_2(p_T)$ and $v_3(p_T)$ values for p/d/³He+Au collisions at 200 GeV
 376 before and after non-flow subtraction, compared for all three subtraction techniques. The
 377 presented results show non-flow contributions which are system-dependent, but the non-
 378 flow subtracted v_2 (top panels) and v_3 (bottom panels) are method-independent within the
 379 uncertainties.

380 These STAR measurements with non-flow subtracted show that for the comparable
 381 charged-hadron multiplicity (N_{ch}) events v_2 and v_3 , values are independent of collision sys-

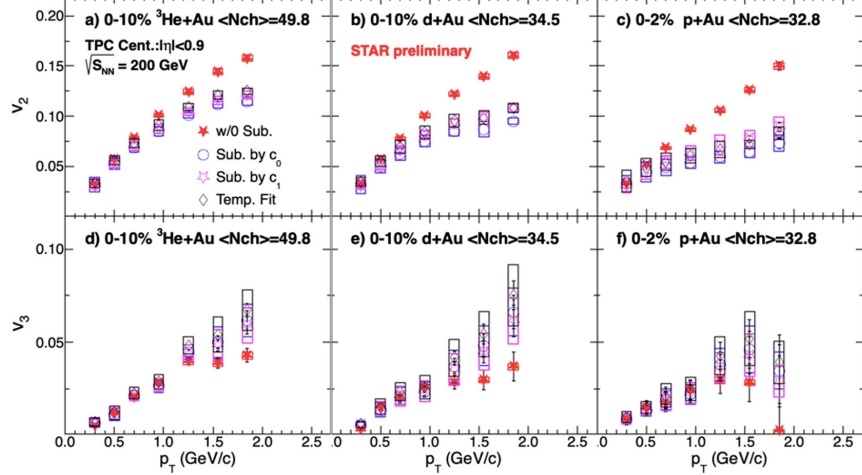


Figure 11: Comparison of the $v_{2,3}(p_T)$ values for p/d/ ^3He +Au collisions at 200 GeV, before and after non-flow subtraction. The figure is taken from Ref [45].

382 tem. These observations are compatible with the significant influence of the subnucleonic
 383 fluctuations-driven eccentricities, $\epsilon_{2,3}$, in a system whose size is primarily determined by
 384 N_{ch} . However, they are incompatible with the notion of shape engineering in p/d/ ^3He +Au
 385 collisions.

386 1.1.3 Light Flavor Spectra and Ultra-peripheral Collisions

387 The Light Flavor Spectra and Ultra-peripheral Collisions (LFSUPC) physics working group
 388 is responsible for the measurements of calibrated production yields and spectra in inclusive
 389 ion-ion collisions, ultra-peripheral collisions, and exclusive p+p collisions.

390 In ion-ion collisions, analysis efforts can be grouped based on the methodology and physics
 391 issues. The general categories include light charge hadrons (π , K , p), strange hadrons (ϕ ,
 392 Λ , Ξ , Ω), light nuclei (d, t, ^3He , ^4He), and hyper-nuclei ($^3_\Lambda H$ and $^4_\Lambda H$). Examples of recent
 393 results from light nuclei and hyper-nuclei are shown in other sections (see sections 2.2.1 and
 394 2.2.2). Here some recent results will be presented from the light charged hadron and strange
 395 hadron analyses.

396 **Light hadron production:** Light charged hadron spectra and yields are measured using
 397 particle identification through dE/dx in the TPC, $1/\beta$ in the time-of-flight detectors and
 398 careful study of the acceptance and efficiency of the detectors. These studies are particularly
 399 useful in defining the basic thermal properties (T and μ_B) of the system. Previous studies
 400 of the light charged hadrons from BES-I measured the spectra and yields at midrapidity.
 401 The newest results now include rapidity dependence which allows for a better understanding
 402 of baryon stopping, which is key to the dependence of μ_B with $\sqrt{s_{NN}}$. New preliminary π ,
 403 K, and p transverse mass spectra are shown as a function of rapidity in Fig. 12 for Au+Au
 404 collisions at $\sqrt{s_{NN}} = 27$ GeV. Additional pre-preliminary results have been produced from
 405 fast offline pre-calibration quality assurance productions from the other BES-II collider and

406 fixed-target energies.

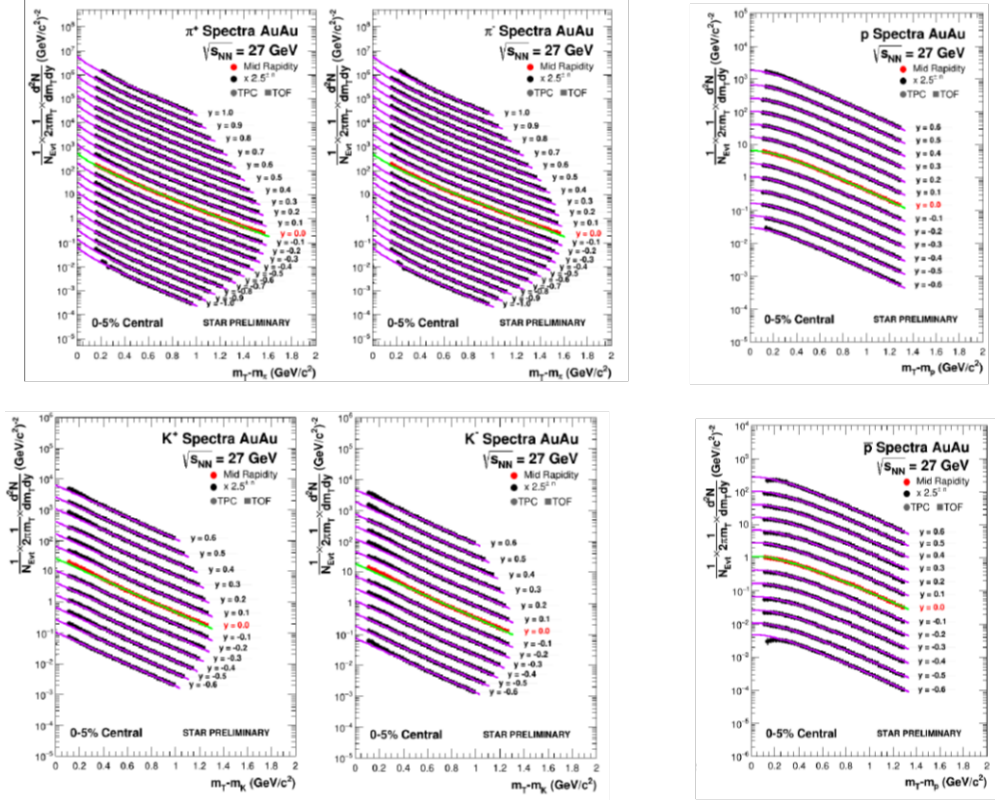


Figure 12: Transverse mass spectra for pions, Kaons, and protons from Au+Au collisions at $\sqrt{s_{NN}} = 27$ GeV as a function of rapidity.

407 **Strange hadron production:** Strange hadron spectra and yields are measured by deter-
 408 mining the invariant mass from the charged daughters from weak decays of neutral strange
 409 hadrons. These studies define the role of the strange quark in the thermodynamic evolution
 410 of the system. STAR has recently implemented a new V^0 finding routine called *KFparticle*
 411 which increases the sensitivity of our strange hadron studies. The highlights of recent
 412 measurements have come from the newest fixed-target data. The fixed-target energy range
 413 covers the production threshold energies for Ξ^- (3.247 GeV), Ω^- (4.09 GeV), Ξ^+ , Ξ^- (4.52
 414 GeV), and Ω^+ , Ω^- (5.22 GeV). Figure 13 shows the invariant mass plots for measurements
 415 of Λ 's, Ξ 's, and Ω 's for fixed-target Au+Au at $\sqrt{s_{NN}}$ of 3.0 and 7.2 GeV. Additional pre-
 416 preliminary measurements have been made at other collider and fixed-target energies. In
 417 addition, studies of the production of the ϕ meson have been made at 3.0 and 7.2 GeV.

418 **exclusive production:** Central exclusive production is measured in p+p collisions using
 419 the very forward roman pot detectors to identify the the two colliding protons and the TPC
 420 to measure the products. Figure 14 shows the invariant mass of pion pairs in exclusive p+p
 421 events at 200 GeV. These are the first measurements at this energy and show significant

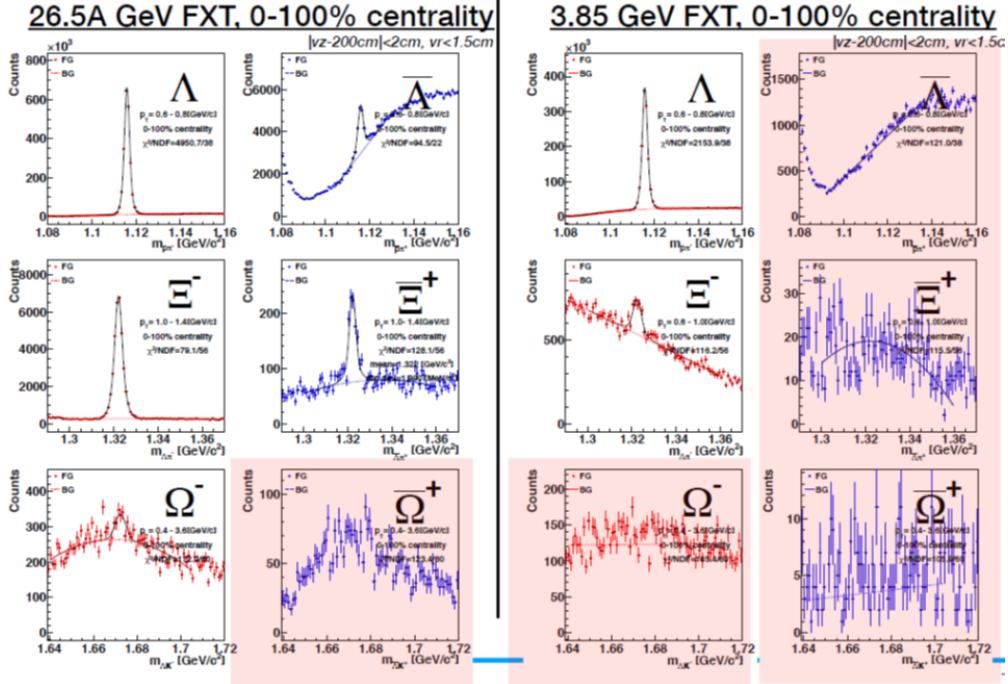


Figure 13: Invariant mass plots for measurements of Λ 's, Ξ 's, and Ω 's for fixed-target Au+Au collisions at $\sqrt{s_{NN}}$ of 3.0 and 7.2 GeV (single beam energies of 3.85 and 26.5 GeV respectively).

422 peaks in the invariant mass spectra that were not predicted by the models. Similar results
 423 are available for kaon and proton pairs.

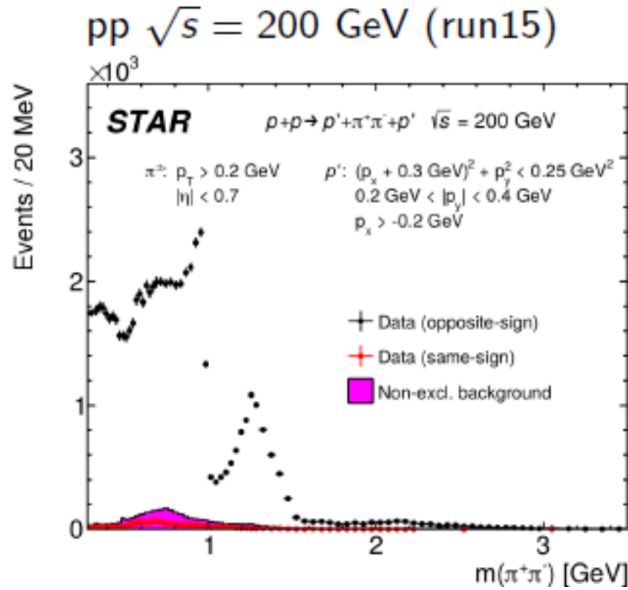


Figure 14: Invariant mass spectra for pion pairs from exclusive p+p events at 200 GeV.

424 **Electromagnetic probes:** Electromagnetic radiation from high-energy heavy-ion collisions provides rich information about the properties of the produced medium. Dileptons directly probe the in-medium electromagnetic correlator of hadronic currents [46, 47]. Dynamical information on in-medium spectral functions encodes not only changes in degrees of freedom, chiral symmetry restoration [48–50], transport properties of medium like the electrical conductivity [51, 52], but also the life time and average temperature of the interacting fireball [53], and the emission history and origin of the radiation [54–57].

431 STAR reported measurement of thermal dilepton radiation ranging from $\sqrt{s_{NN}} = 200$ GeV down to 19.6 GeV [58–61]. A significant excess in the low-mass region when compared to the known hadronic sources has been observed. It was shown that the predictions of hadronic manybody theory for a melting ρ meson, coupled with QGP emission utilizing a modern lattice QCD-based equation of state [51, 62], yield a quantitative description of dilepton spectra in heavy-ion collisions [58, 61]. This is demonstrated in Fig. 15 (left panel). Moreover, it has been shown that the integrated low-mass excess radiation provided a direct measure of the total fireball lifetime [60]. Secondary vertex rejection employing information provided by the Heavy Flavor Tracker installed for Run 2014 and 2016 will enable unique temperature measurements of the QGP.

441 The low-mass line shape will provide a critical test of the ρ -melting scenario (which is consistent with expectations of chiral symmetry restoration) at vanishing baryon chemical potential. A precision measurement at top RHIC energy will provide additional constraints that can be directly tested against the lattice QCD predictions and will be put in focus in runs 2023 - 2025 (see section 2.4.3).

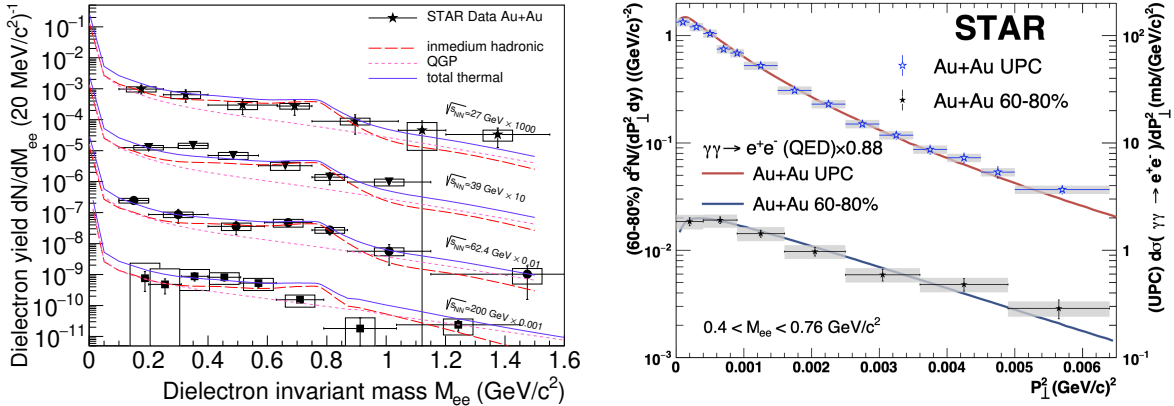


Figure 15: Left: Acceptance-corrected dilepton excess mass spectra, normalized by dN_{ch}/dy , for Au+Au collisions at $\sqrt{s_{NN}} = 27, 39, 62.4, 200$ GeV. Right: Comparison of the P_T^2 distribution in 60 – 80% central Au+Au collisions with that in UPCs [63].

446 Dileptons generated by the intense electromagnetic fields accompanying the relativistic heavy nuclei at large impact parameters [?], in ultra-peripheral collisions (UPC) where there is no nuclear overlap has recently become experimentally accessible, offering several opportunities. According to the equivalent photon approximation (EPA), the electromagnetic field generated by an ultra-relativistic nucleus can be viewed as a spectrum of quasi-real pho-

451 tons coherently emitted by the entire nucleus and the dilepton production process can be
 452 represented as $\gamma + \gamma \rightarrow l^+ + l^-$. Recently, the STAR and ATLAS collaborations made mea-
 453 surements of dileptons at small impact parameters with nuclear overlap, and found that the
 454 electromagnetic production of dileptons can also occur in hadronic collisions. Furthermore,
 455 a significant P_\perp broadening as shown in Fig. 15 (right panel) effect for lepton pairs produced
 456 by the two photon scattering process has been observed in hadronic collisions compared to
 457 measurements in UPC and to EPA calculations. Precision measurements will provide an im-
 458 portant constraints for quantitative theoretical analyses of magnitude and duration of initial
 459 magnetic fields. It was perceived that photons participating in such collisions are quasi-real
 460 with transverse-momentum $k_t \simeq 1/R$ (30 MeV/c) reflecting the virtuality and uncertainty
 461 principle of their origin. This led to the implementation in many EPA models that the initial
 462 transverse momentum of the dilepton pairs does not depend on impact parameter and the
 463 transverse space coordinates where the pair are created are randomly distributed due to the
 464 same principles. Our new measurements of centrality dependence and azimuthal distribu-
 465 tions have shown that the photons behave like real photons in all observables and the renewed
 466 models and theories have demonstrated that correction to the real photon approximation is
 467 suppressed at the order of $1/\gamma^2$ even to the pair's transverse momentum distribution. The
 468 discovery of the Breit-Wheeler process and the utilization of linearly polarized photons in
 469 UPC are conceptually and experimentally highly nontrivial. With future high statistics data
 470 with larger TPC acceptance in UPC, we can explore the phase space of photon collisions
 471 in transverse momentum, rapidity and momentum-space-spin correlations in extreme QED
 472 field [64, 65] (see section 2.4.3).

473 1.1.4 Heavy-Flavor

474 The production of heavy-flavor (HF) quarks proceeds predominately via the hard scatterings
 475 of partons in $p(A)+p(A)$ collisions. This fact gives rise of the utility of heavy-flavor hadron
 476 measurements in heavy-ion experiments since they are produced independently of the QCD
 477 medium and probe it's properties by scattering with the medium constituents. Topics of
 478 medium-induced parton energy loss, QGP transport properties, hadronization mechanisms,
 479 and quarkonia melting are some of the pivotal studies that have emerged within the HF
 480 category. Besides the highlights discussed in detail below, the following measurements have
 481 been recently published: First Measurement of Λ_c Baryon Production in Au+Au Collisions
 482 at $\sqrt{s_{NN}}=200$ GeV [66]; Measurement of inclusive J/ψ suppression in Au+Au collisions at
 483 $\sqrt{s_{NN}}=200$ GeV through the dimuon channel at STAR [67]; First Observation of the Directed
 484 Flow of D^0 and \bar{D}^0 in Au+Au Collisions at $\sqrt{s_{NN}}=200$ GeV [68].

485 Recent measurements of D^0 [69] and D^+ (shown at the Hard Probes 2020 conference)
 486 meson yields as a function of transverse momentum (p_T) in heavy-ion collisions show a
 487 significantly suppressed spectrum with respect to $p + p$ collisions. The two mechanisms that
 488 predominately produce suppressed meson distributions are parton energy loss in the QGP
 489 and different hadronization schemes. The latter is nicely illustrated via the measurement of
 490 the Λ_c^+/D^0 yield ratio [66], which is significantly enhanced with respect to the expectation
 491 in $p + p$ collisions and is attributed to baryon production via coalescence hadronization.

492 Recently, STAR has measured the ratio of D_s^+/D^0 yields in heavy-ion collisions, which
 493 is important as it probes charm hadronization and strangeness enhancement mechanisms.
 494 Utilizing the excellent pointing resolution resolution of the Heavy Flavor Tracker (HFT), the
 495 D_s is measured in 2014 and 2016 data and is topologically reconstructed using a multi-variate
 496 analysis (MVA) optimized for signal selection. The final results of D_s^+ yield with respect
 497 to D^0 were reported in the 2019 Quark Matter conference, and are shown in Figure 16.
 498 The ratios are split into 0-10% (blue points) and 10-40% (red points) centrality regions
 499 and are consistent within experimental uncertainties. Shown in the left panel are the ratios
 500 averaged over $p + p/e + p/e + e$ collisions, PYTHIA, and a model calculation (TAMU)
 501 including coalescence hadronization for 10-40% centrality. Shown in the right panel are model
 502 calculations including sequential (solid lines) and simultaneous (dashed lines) coalescence for
 503 both Au+Au collisions at RHIC and Pb+Pb collisions at LHC energies. Data from ALICE
 504 are also shown in the right panel, and are consistent within uncertainties with STAR data.
 505 The data show the D_s yield is significantly enhanced in Au+Au collisions with respect to that
 506 of elementary $p + p/e + p/e + e$ collisions. Comparing the STAR data to model calculations,
 507 the model of sequential coalescence is able to better capture the trends in data.

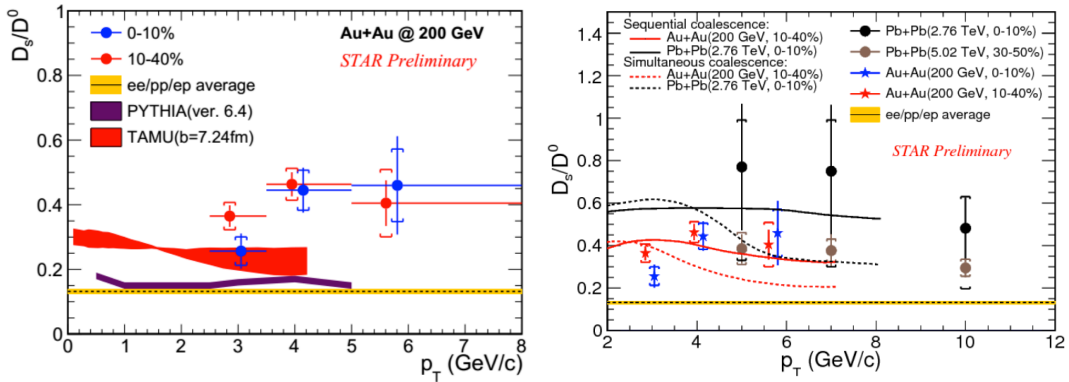


Figure 16: Left: STAR measurement of the D_s/D^0 ratio in 0-10% (blue circles) and 10-40% (red circles) centrality Au+Au collisions as a function of $D_s p_T$. The yellow shaded band shows the average from $p + p/e + p/e + e$ collisions. The purple and red shaded bands show the expectations from PYTHIA and TAMU model calculations, respectively. Right: STAR D_s/D^0 ratio measurements compared to model calculations including sequential (solid lines) and simultaneous coalescence (dotted lines) hadronization, and data from the ALICE measurements (blue and brown circles) and the respective model calculations at LHC energies.

508 The mass dependence of parton energy loss has been probed in heavy-ion collisions with
 509 the measurements of light- and heavy-flavor hadron nuclear modification factors (R_{AA}).
 510 At high p_T , where mass effects are predicted to significantly modify the quark energy loss
 511 from gluon radiation in the QGP, the values of light-flavor and charm hadron R_{AA} are
 512 measured to be degenerate, and can be explained by mechanisms that are not related to
 513 parton energy loss (e.g, see [70]). In that respect systematic comparisons of both bottom
 514 and charm hadron nuclear modification factors are predicted to be a clean probe of the

515 mass dependence of parton energy loss by several model calculations [71–74]. However, from
516 an experimental point of view measuring bottom hadrons have been difficult at RHIC due
517 to the low bottom quark production cross-section, and have only been accessible via the
518 measurement of displaced electrons or charmed hadrons. STAR has now reported at the
519 2019 Quark Matter conference an updated measurement of single electrons from bottom
520 semileptonic decays utilizing both 2014 and 2016 data sets. The contribution of bottom-
521 and charm-decayed electrons, and backgrounds, are topologically separated using the three-
522 dimensional distance-of-closest approach (DCA) utilizing the HFT detector. In contrast
523 to previous measurements utilizing the transverse dimension DCA (DCA_{xy}), the 3D DCA
524 is able to separate charm- and bottom-decay electrons with greater significance since the
525 longitudinal and transverse DCA have similar resolution. The updated STAR measurement
526 also includes an improved electron identification selection, which is based off a projective
527 likelihood MVA. The improvement in the mis-identified hadron fraction is a factor of two
528 when compared to traditional cut-based particle identification. The results of bottom- and
529 charm-decayed electron R_{AA} are shown in Figure 17 in the top panel, and their ratio in the
530 bottom panel. A constant fit to the double ratio is used to quantify the enhanced $b \rightarrow e$
531 R_{AA} and is measured to be $1.92 \pm 0.25(\text{stat.}) \pm 0.21(\text{syst.})$. Shown in the bottom panel as the
532 hashed blue curve is a null hypothesis where we assume equal values of R_{AA} for charm and
533 bottom hadrons and then fold the distributions to the decay-electron, and subsequently take
534 a double ratio. Performing this exercise shows the effects from different production spectra,
535 fragmentation, and decay phase-space of charm and bottom hadrons, and it is clearly seen
536 these effects have a small impact on the double ratio. The double ratios of $b \rightarrow e$ to $c \rightarrow e$
537 R_{CP} are also measured and a similar constant fit as in the R_{AA} case is performed and found
538 to be $1.68 \pm 0.15(\text{stat.}) \pm 0.12(\text{syst.})$ and $1.38 \pm 0.08(\text{stat.}) \pm 0.03(\text{syst.})$ for the ratios of $R_{CP}(0-$
539 $20\%/40-80\%)$ and $R_{CP}(0-20\%/20-40\%)$, respectively. We additionally compare the data to
540 a modified Langevin transport model (DUKE) [72] which includes the mass dependence
541 of parton energy loss, and within uncertainties the data and model are consistent in both
542 the absolute R_{AA} data and the double ratios of R_{AA} and R_{CP} . Combining the agreement
543 between model and experiment, and the quality of the data, these observations represent,
544 for the first time, evidence of mass-ordering of parton energy loss in heavy-ion collisions.

545 Measurements of heavy-flavor flow are also essential to understanding the QGP properties
546 as particle flow and yield provide a test-bed for model calculations to simultaneously describe
547 the data. It has already been established by STAR [75] in 200 GeV Au+Au collisions that D^0
548 hadrons have a significant elliptic flow that is comparable to light-flavor hadrons after taking
549 into account particle mass and number of constituent quarks. Measurements of heavy-flavor
550 hadron flow at lower collision energies have been explored via the measurements of single
551 electron elliptic flow in the 2017 and 2018 data sets, and was reported at the 2020 Hard
552 Probes conference. Previous STAR measurements of heavy-flavor electron v_2 in 62.4 and 39
553 GeV Au+Au collisions [76] were statistically limited and within experimental uncertainties
554 consistent with zero. The data collected during Run-17 and Run-18 at $\sqrt{s_{NN}}$ 54.4 and 27
555 GeV Au+Au collisions, respectively, are more than an order of magnitude larger in statistics
556 and allow for a more precise measurement. The heavy-flavor decay electron v_2 is extracted

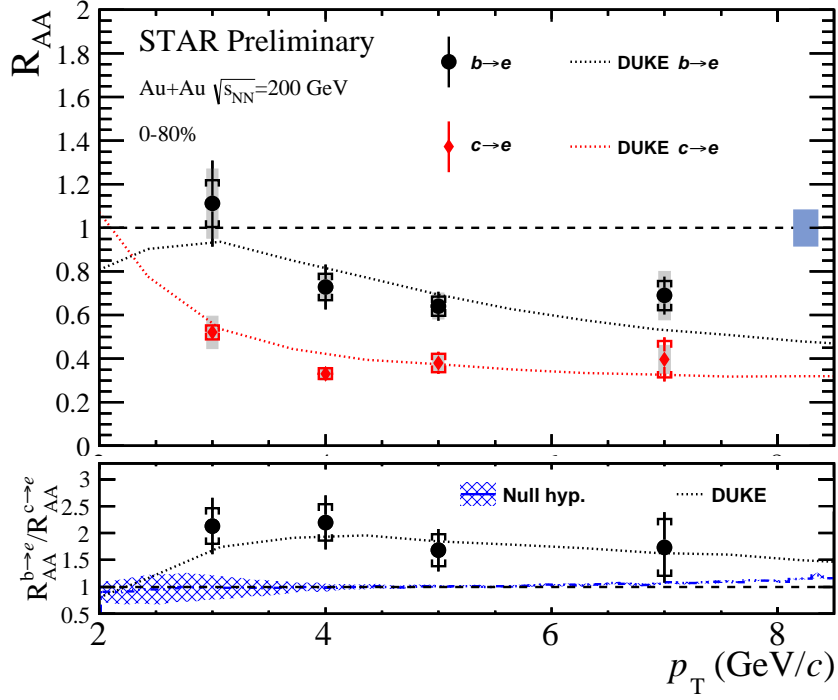


Figure 17: Top: Data for bottom- (blue stars) and charm-decay (red diamonds) electron R_{AA} as a function of electron p_T . The DUKE model calculation are shown as the respectively colored dotted lines. Bottom: The double ratio of bottom to charm R_{AA} , and the null hypothesis (explained in the text) shown as the blue shaded band and DUKE calculation as the dotted line.

557 from the inclusive electron v_2 by correcting for electron v_2 from hadron and photon decays.
 558 The data are shown in Figure 18 for both 54.4 and 27 GeV Au+Au collisions, and compared
 559 to previously published STAR data in 200 GeV Au+Au collisions [76]. The 54.4 GeV data
 560 show a significant v_2 that is comparable to 200 GeV Au+Au collisions, indicating heavy-
 561 flavor hadrons gain significant collective flow in the produced medium in 54.4 GeV Au+Au
 562 collisions. The data at 27 GeV indicate a hint of non-zero v_2 , but still have considerable
 563 uncertainties due to a lower signal-to-background ratio.

564 STAR has also recently reported at the 2019 Quark Matter conference charm-decayed
 565 electron v_1 and v_2 , and bottom-decay electron v_2 in 200 GeV Au+Au collisions utilizing
 566 the HFT to isolate charm- or bottom-decayed electrons, respectively. The measurement of
 567 charm-decayed electron v_1 and v_2 was compared to previous STAR data of D^0 mesons, and
 568 show consistency between the two measurements. In the former case the measured slope
 569 of the charm-decayed electron v_1 versus rapidity corroborated the recently measured large
 570 negative v_1 slope by STAR [68], and with an improved significance of 5σ . The bottom-decay
 571 electron v_2 was measured to have a non-zero v_2 with a significance of about 3.4σ , and is
 572 consistent in magnitude with expectations from the DUKE model [72]. This is the first
 573 significant measurement of bottom hadron v_2 at RHIC.

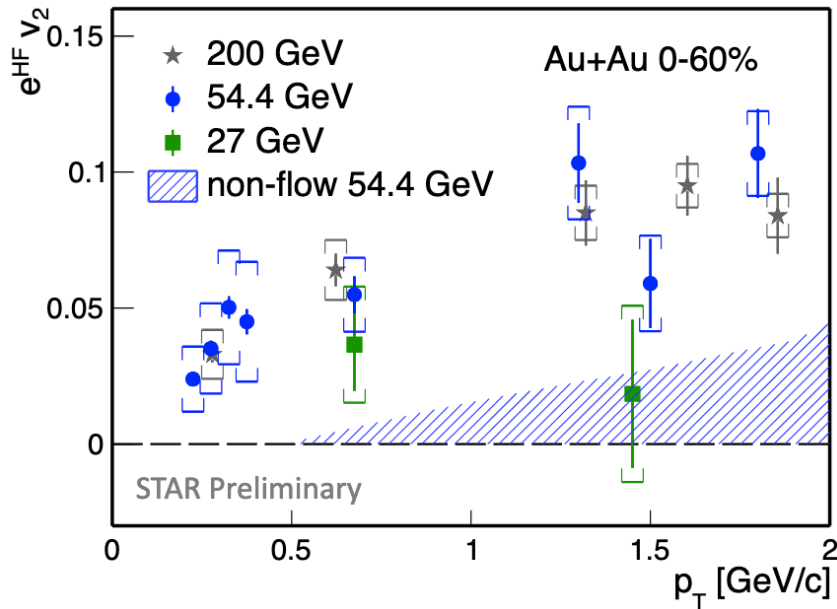


Figure 18: Heavy-flavor electron v_2 as a function of electron p_T in 54.4 (blue circles) and 27 (green squares) GeV Au+Au collisions. The STAR published data at 200 GeV are also shown as the gray stars. The shaded blue histogram shows the estimated non-flow contribution in the 54.4 GeV data.

574 Recent measurements of quarkonia have opened up new ways to probe production mech-
 575 anisms by measuring their distributions with jets. It has been observed that there is no
 576 simultaneous description of models to the J/ψ spectra and polarization data. It has also
 577 been shown that measurements of quarkonia fragmentation functions in jets can have good
 578 discriminating power between different models. Recently reported at the 2020 Hard Probes
 579 conference, was the measurement of J/ψ mesons within jets as a function of the J/ψ p_T
 580 fraction ($z = p_T(J/\psi)/p_T(jet)$). The results are shown in Figure 19 for a given set of jet
 581 reconstruction requirements, and unfolded to account for detector smearing. The depen-
 582 dence on cone size was also investigated and the data showed as the cone sized is increased
 583 the z distribution became more populated at lower values. Compared with the data in the
 584 same plot is the expectation of PYTHIA simulations (shown as the gray shaded histogram).
 585 The data show a clear discrepancy with PYTHIA, suggesting J/ψ mesons are not produced
 586 mostly in isolation. This new measurement is expected to provide valuable input to the
 587 theory community.

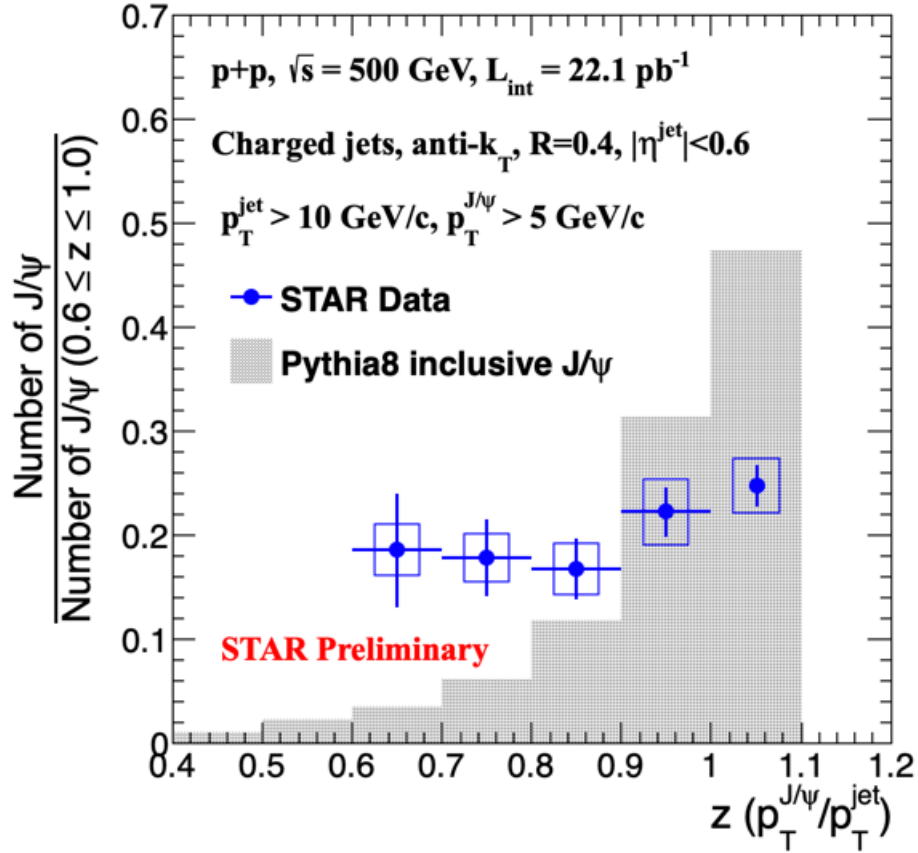


Figure 19: J/ψ momentum fraction in jets in 500 GeV $p+p$ data for a given set of jet reconstruction parameters listed. Also shown is the expectation from PYTHIA simulation as the gray histogram.

588 1.2 CME Search and Isobar Run

589 1.2.1 Introduction

590 Finding a conclusive experimental signature of the Chiral Magnetic Effect (CME) has be-
 591 come one of the major scientific goals of the heavy-ion physics program at the Relativistic
 592 Heavy Ion Collider (RHIC). The existence of CME will be a leap towards an understanding
 593 of the QCD vacuum, establishing a picture of the formation of deconfined medium where
 594 chiral symmetry is restored and will also provide unique evidence of the strongest known
 595 electromagnetic fields created in relativistic heavy-ion collisions [77, 78]. The impact of such
 596 a discovery goes beyond the community of heavy-ion collisions and will possibly be a mile-
 597 stone in physics. Also, as it turns out, the remaining few years of RHIC run and analysis
 598 of already collected data probably provides the last chance for dedicated CME searches in
 599 heavy-ion collisions in the foreseeable future. Over the past years significant efforts from
 600 STAR as well as other collaborations have been dedicated towards developing new methods
 601 and observables to isolate the possible CME-driven signal and non-CME background con-
 602 tributions in the measurements of charge separation across the reaction plane. Many clever
 603 ideas have been proposed and applied to existing data. The general consensus is that mea-
 604 surement from the isobar collisions, Ruthenium+Ruthenium (Ru+Ru) that has 10 – 18%
 605 higher B-field than Zirconium+Zirconium (Zr+Zr), provides the best solution to this prob-
 606 lem. During the time when this beam user request document is being written, the analysts
 607 from the STAR collaboration are about to start the final step of the (four-step) blind analysis
 608 of the isobar data that we discuss at length in the following section.

609 1.2.2 Modality of Isobar Running at RHIC

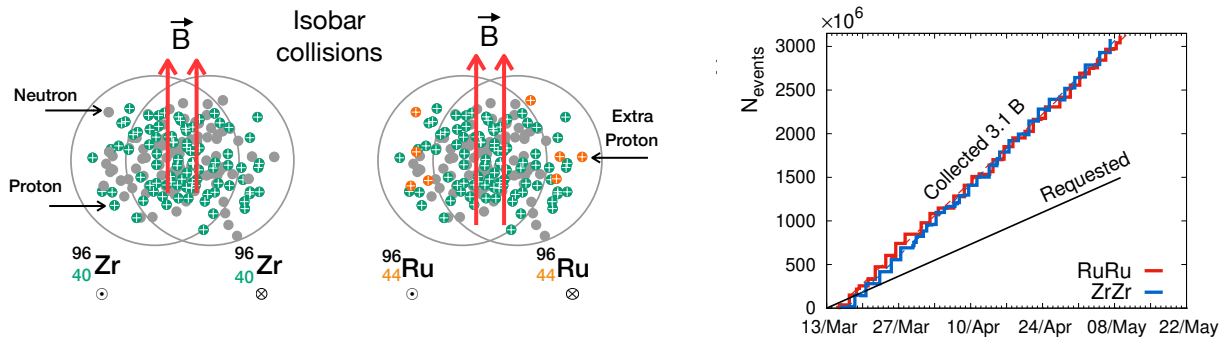


Figure 20: Left: Cartoon of the isobar collisions, about 10 – 18% stronger B-field is expected in Ru+Ru collisions as compared to Zr+Zr collisions due to four extra protons in each Ru nucleus. Right: Summary of the data collected for isobar collisions during Run 18 – almost a factor of two more events were collected than the request 1.5 Billion events over the course of 3.5 weeks.

610 The idea of colliding isobar, particularly Ru+Ru and Zr+Zr to make a decisive test of

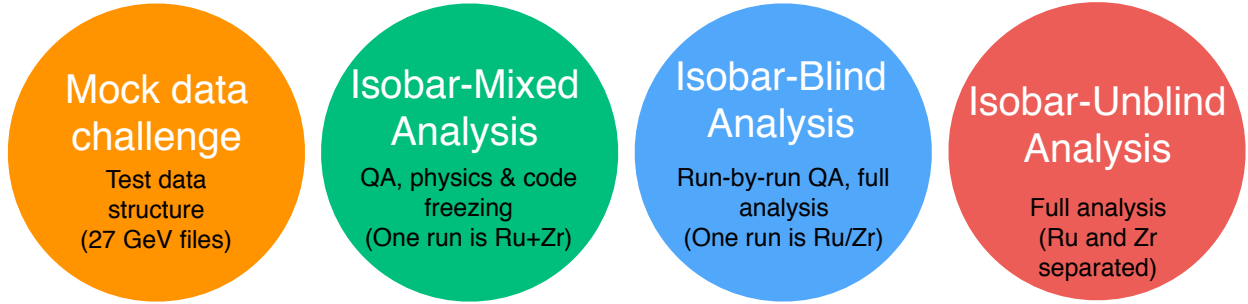


Figure 21: The steps of isobar blind analysis. This cartoon is based on the procedure for the blind analysis of isobar data that have been outlined in Ref [84]. During the time when this document is being written STAR collaboration has finished first three steps and about to begin the final step of isobar-unblind analysis (shown in red).

611 CME was proposed by Voloshin in Ref [79], the same paper which also proposed to use
 612 Uranium collisions to disentangle signal and background of CME. The possible difference in
 613 the signals relies on the 10–18% higher B-field in Ru+Ru compared to Zr+Zr [80] in contrast
 614 to about 4% difference in flow driven background [36]. Such estimates are sensitive to details
 615 of shapes, charge distribution and neutron skin thickness of the two isobar nuclei [80–82]. In
 616 the 2017-18 RHIC beam user request [83] STAR collaboration therefore proposed to collect
 617 data for two 3.5 week runs in the year 2018. The projection was based on the prospect of
 618 achieving five-sigma significance in a scenario where the measurement of $\Delta\gamma$ has 80% non-
 619 CME background. This however corresponds to the fact that the systematic uncertainty in
 620 the measurements has to be within a few percent and below the statistical uncertainty of the
 621 measurements. This started a large scale collaboration wide effort in synergy with the RHIC
 622 collider accelerator department to plan for the isobar running in the year 2018. Based on the
 623 studies of previous years of data from $Au + Au$ and $U + U$ collisions several major sources of
 624 systematics in the measurement of $\Delta\gamma$ were identified. The major sources include: run-to-run
 625 variation of detector response due to loss of acceptance, change in efficiency and variation in
 626 luminosity that affects the number of reconstructed tracks in the Time Projection Chamber.
 627 This eventually leads to uncorrectable systematic uncertainties in $\Delta\gamma$, the main observable
 628 to measure charge separation across event plane. In order to minimize such systematics the
 629 proposal were to: 1) switch species in RHIC between stores and, 2) keep long stores to level
 630 the luminosity aiming for specific rates in the coincidence measurements of beam fragments
 631 by the STAR zero-degree calorimeters. The aim was to maintain exact balance of run and
 632 detector conditions for the two species so that observations in the two systems are equally
 633 affected and can later on be largely eliminated in the ratios of observables.

634 1.2.3 Blinding of Data Sets and Preparation for Analyses

635 With the successful conclusion of the isobar run in the year 2018 STAR experiment collected
 636 more than 3 billion events for each isobar species. The next step was to develop the plans
 637 for a blind analysis, the main idea behind which is to eliminate possible unconscious biases.

638 A total of five institutional groups are expected to perform the analysis of the data. The
639 analysts from each group will focus on a specific aspect of the analysis described in the
640 previous section although in many cases there are substantial overlap in some analyses that
641 will help cross check the results. An important part of the blind analysis is the blinding of the
642 data. The details of the blinding of the data structure is decided by members of a blinding-
643 committee who are not part of the team of analysts and will work in close collaboration with
644 STAR experts who are part of the production team. The idea is to provide the analysts
645 the access to data in files where species-specific information are disguised or removed before
646 the final step of unblinding. A careful consideration is taken by the blinding-committee to
647 make sure only the essential information to do the analysis-specific quality assurance of the
648 data is available to the analysts. Some of the quality assurance, calibration and centrality
649 determination work that require species information are done only by STAR experts who are
650 not a part of the team of analysts. Above all, the main goal of the committee is to make sure
651 that under no circumstances physics analysts can access un-blinded data that can jeopardize
652 the blind analysis. For example, all the data sets are produced with pseudo-run-number that
653 cannot be used by the analysts to retrieve the exact species information.

654 1.2.4 Methods for the Isobar Blind Analyses

655 The detailed procedure for the blind analyses of isobar data have been outlined in Ref [84].
656 Figure.21 is a cartoon that summarizes the four steps and the main idea.

657 In the zeroth step shown in the extreme left of Fig.21 (by orange circle) is the mock
658 data challenge which is not exactly a step of the isobar data analysis but a crucial step
659 to familiarize the analysts with the technicalities of the data structures that have been
660 specifically designed for blind analysis.

661 The first step shown in Fig.21 (by green circle) as the "isobar-mixed analysis" or "mixed-
662 blind analysis" is truly the first step of blind analysis. This is also the most challenging
663 steps from the point of view of the analysts. In this step the analysts are provided with data
664 sample where each run comprise of events that are mixed samples from two species. In this
665 step the analysts perform the full quality assurance (QA) and physics analysis of the data,
666 document every details of steps of the procedure and freeze the codes. After the completion
667 of this step no changes to the analysis code is permissible. Also, no changes in the analysis
668 procedure is allowed. The only permissible change in the following step is to reject bad runs
669 or pile-up events. However, in order to avoid unconscious bias in analysis such rejection
670 cannot be done arbitrarily and an automated algorithm must be developed in this step and
671 the related codes have to be frozen. The stability of the automated QA algorithm is tested
672 with some of the existing data sets of Au+Au and U+U collisions.

673 The second step shown in Fig.21 (by blue circle) is referred to as the "isobar-blind anal-
674 ysis" or "unmixed-blind analysis". From this step on-wards the analysts are allowed to run
675 their previously frozen codes. The main purpose of this step is to perform run-by-run QA
676 of the data sample. For this the analysts are provided with files each of which contain data
677 from a single species that is either Ru or Zr. However, there are two conditions: the files
678 contain limited number of events that cannot lead to any statistically significant result and

679 the species information is not revealed. Although a pseudo-run-number is used for each file,
 680 the time ordering is preserved with a unique mapping that is unknown to the analysts. It
 681 is important to maintain the time ordering to identify time-dependent changes in detectors
 682 and run conditions as a part of the run-by-run quality assurance. With this limited data
 683 sample the analysts need run the frozen automated algorithm to identify bad runs. A similar
 684 automated algorithm is also used for identifying and rejecting bad runs. After this step no
 685 more changes are allowed in terms of QA.

686 The final step of isobar blind analysis is shown by red circle in Fig.21 is referred to as
 687 "isobar-unblind" analysis. In this step the species information will be revealed and the physics
 688 results will be produced by the analysts using the previously frozen codes. The finding from
 689 this step will be directly be submitted for publication without any kind of alteration. If a
 690 mistake is found in the analysis code, the erroneous results will also accompany the corrected
 691 results.

692 1.2.5 Observables for Isobar Blind Analyses

693 Isobar blind analysis will specifically focus on these five observables. The following five
 694 sections describe these procedures in brief with comments on the outlook for isobar blind
 695 analysis: 1) Differential measurements of $\Delta\gamma$ to identify and quantify backgrounds, 2) mea-
 696 surement of higher order harmonics of γ -correlator, 3) exploiting the relative charge separa-
 697 tion across participant and spectator planes, 4) the use of R-observable to measure charge
 698 separation and 5) the use of signed balance function. The first three approaches are based on
 699 aforementioned three-particle correlator and the last two employ slightly different approaches
 700 to quantify charge separation.

701 **Differential measurements of $\Delta\gamma$ to identify and quantify background** *Invariant*
 702 *mass dependence of charge separation:* Differential measurements of $\Delta\gamma$ with invariant mass
 703 and relative pseudorapidity provide interesting prospects to identify and quantify the sources
 704 of flow and non-flow driven backgrounds. The idea to use invariant mass is simple and was
 705 first introduced in Ref [85]. Resonances are widely identified by observing structures in the
 706 invariant mass spectra of the decay daughters. Take a pair of opposite sign pions for example,
 707 a large fraction of them come from the neutral resonances that show up in the invariant
 708 mass spectrum of $m_{inv}(\pi^+ + \pi^-)$. If we restrict the analysis to pairs of pions, differential
 709 measurements of $\Delta\gamma$ with $m_{inv}(\pi^+ + \pi^-)$ should also show similar peak like structures if
 710 background from neutral resonances dominate the charge separation. Indeed similar peak
 711 structures are observed and a careful analysis is performed by STAR collaboration to extract
 712 the possible fraction of CME signals from measurements [86]. This analysis relies on the
 713 assumption that CME signals do not show peak like structures in $m_{inv}(\pi^+ + \pi^-)$ therefore
 714 calls for more theoretical inputs in this direction.

715 *Relative pseudorapidity dependence:* The relative pseudorapidity dependence of azimuthal
 716 correlations are widely studied to identify sources of long-range components that are domi-
 717 nated by early time dynamics as compared to late time correlations that are prevented by

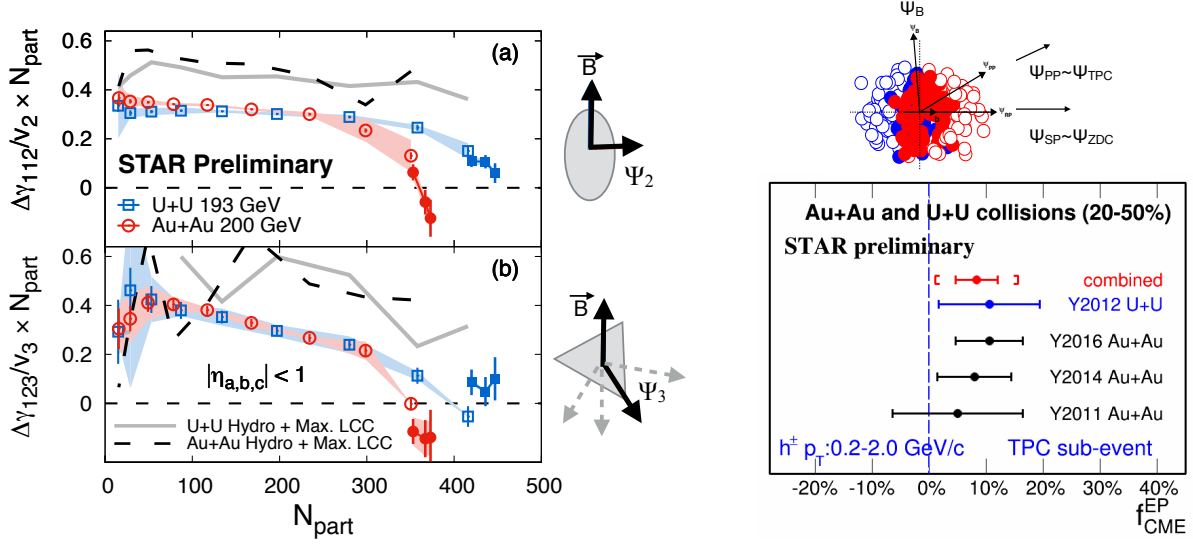


Figure 22: (Left) Measurement of charge separation along second and third order event planes in Au+Au and U+U collisions. (Right) Fraction of possible CME signal in the measurement of $\Delta\gamma$ with respect to spectator and participant planes [88].

causality to appear as short-range correlations. The same can be extended to charge dependent correlations that provides the impetus to explore the dependence of $\Delta\gamma$ on the pseudorapidity gap between the charge carrying particles $\Delta\eta_{ab} = |\eta_a - \eta_b|$ in $\langle \cos(\phi_a^\alpha + \phi_b^\beta - 2\Psi_{RP}) \rangle$. Such measurements have been performed in STAR with Au+Au and U+U data. It turns out that the possible sources of short range correlations due to photon conversion of $e^+ - e^-$, HBT and Coulomb effects can be identified and described as Gaussian peaks at small $\Delta\eta_{ab}$, the width and magnitude of which strongly depend on centrality and system size. Going to more peripheral centrality bins it becomes harder and harder to identify such components as they overlap with sources of di-jets fragmentation that dominate both same-sign and opposite sign correlations. An effort to decompose different components of $\Delta\gamma$ via study of $\Delta\eta_{ab}$ can be challenging although a clear sign of different sources of correlations are visible in change of shape of individual same-sign and opposite sign measurements of γ -correlator [87].

In any case, these differential measurements of $\Delta\gamma$ in isobar collisions provide the prospect to extract the $m_{inv}(\pi^+ + \pi^-)$ and $\Delta\eta$ dependence of CME signals that will provide much deeper insights on the origin of the effect.

Mixed harmonics measurements with second and third order event planes

In order to proceed in this section it is better to rewrite the conventional γ -correlator by a more general notation as $\gamma_{112} = \langle \cos(\phi_a^\alpha + \phi_b^\beta - 2\Psi_2) \rangle$. The idea is to measure charge separation across the third harmonic event plane by constructing a new correlator $\Delta\gamma_{123} = \gamma_{123}(OS) - \gamma_{123}(SS)$, where $\gamma_{123} = \langle \cos(\phi_a^\alpha + 2\phi_b^\beta - 3\Psi_3) \rangle$ that was introduced by CMS collaboration in Ref [89]. Since the Ψ_3 plane is random and not correlated to B-field direction (see Fig.22), γ_{123} is purely driven by non-CME background, the contribution of

740 which should go as v_3/N . This is very useful to contrast signal and background sce-
741 nario by comparing the measurements in two isobaric collision systems. Since Ru+Ru has
742 larger B-field than Zr+Zr but have comparable background, the case for CME would be
743 as follows: $(\Delta\gamma_{112}/v_2)^{\text{Ru+Ru}}/(\Delta\gamma_{112}/v_2)^{\text{Zr+Zr}} > 1$ and $(\Delta\gamma_{112}/v_2)^{\text{Ru+Ru}}/(\Delta\gamma_{112}/v_2)^{\text{Zr+Zr}} >$
744 $(\Delta\gamma_{123}/v_3)^{\text{Ru+Ru}}/(\Delta\gamma_{123}/v_3)^{\text{Zr+Zr}}$. Fig.22 (left) shows the measurement of these observables
745 in U+U and Au+Au collisions. Within the uncertainties of the measurements no significant
746 difference in the trend of $\Delta\gamma_{112}/v_2$ and $\Delta\gamma_{123}/v_3$ is observed for the two collision systems
747 except for the very central events. Predictions from hydrodynamic model calculations with
748 maximum possible strength of local charge conservation [36] is shown on the same plot.
749 Overall observation indicates background dominate the measurements and a similar analysis
750 of the isobar data is highly anticipated.

751 **Charge separation along participant and spectator planes** This analysis makes use
752 of the fact that B-field driven signal is more correlated to spectator plane in contrast to
753 flow driven background which is maximum along the participant planes. The idea was first
754 introduced in Ref [90] and later on followed up in Ref [91]. It requires measurement of
755 $\Delta\gamma$ with respect to the plane of produced particles, a proxy for participant plane as well
756 as with respect to the plane of spectators. In STAR the two can be done by using Ψ_2
757 from TPC and Ψ_1 from ZDC respectively. The approach is based on three main assump-
758 tions: 1) measured $\Delta\gamma$ has contribution from signal and background that can be expressed
759 as $\Delta\gamma = \Delta\gamma^{\text{bkg}} + \Delta\gamma^{\text{sig}}$, 2) the background contribution to $\Delta\gamma$ should follow the scaling
760 $\Delta\gamma^{\text{bkg}}(\text{TPC})/\Delta\gamma^{\text{bkg}}(\text{ZDC}) = v_2(\text{TPC})/v_2(\text{ZDC})$ and, 3) the signal contribution to $\Delta\gamma$ should
761 follow the scaling $\Delta\gamma^{\text{sig}}(\text{TPC})/\Delta\gamma^{\text{sig}}(\text{ZDC}) = v_2(\text{ZDC})/v_2(\text{TPC})$. The first two have been
762 known to be working assumptions, widely used for a long time and can be used to test the
763 case of CME [91] if $(\Delta\gamma/v_2)(\text{ZDC})/(\Delta\gamma/v_2)(\text{TPC}) > 1$. The validity of the last one was
764 studied and demonstrated in Ref [90]. Using all three equations one can extract [88] the
765 fraction of possible CME signal $f_{\text{CME}} = \Delta\gamma^{\text{sig}}/\Delta\gamma$ in a fully data-driven way as shown in
766 Fig.22(right). This analysis will be done with the isobar data and the case for CME will be
767 $f_{\text{CME}}^{\text{Ru+Ru}} > f_{\text{CME}}^{\text{Zr+Zr}} > 0$.

768 **Alternate measure: The novel R-observable** The R -observable is actually a distri-
769 bution, introduced in Ref [94], and defined as the ratio of two distribution functions of
770 the quantity ΔS parallel and perpendicular to B-field direction defined as $R_{\Psi_m}(\Delta S) =$
771 $C_{\Psi_m}(\Delta S)/C_{\Psi_m}^\perp(\Delta S)$. Here ΔS measures the difference in the dipole moment of the positive
772 and negative charge in an event (see Ref [94] for details). The shape of $R_{\Psi_2}(\Delta S)$ will be
773 sensitive to CME as well as non-CME background whereas $R_{\Psi_3}(\Delta S)$ is purely driven by
774 non-CME background and serves as a baseline. Model calculations have established several
775 unique features of this observable: 1) presence of CME signal will lead to a concave shape of
776 the $R_{\Psi_2}(\Delta S)$, 2) increasing strength of CME signal will increase the concavity of $R_{\Psi_2}(\Delta S)$,
777 3) in presence of CME, the concavity of $R_{\Psi_2}(\Delta S)$ will be larger than that of $R_{\Psi_3}(\Delta S)$. The
778 measurement of R_{Ψ_m} is shown in Fig.23. The quantity $\Delta S''$ shown is a slight variant of (ΔS)
779 that incorporates correction for particle number fluctuations and event plane resolution. The

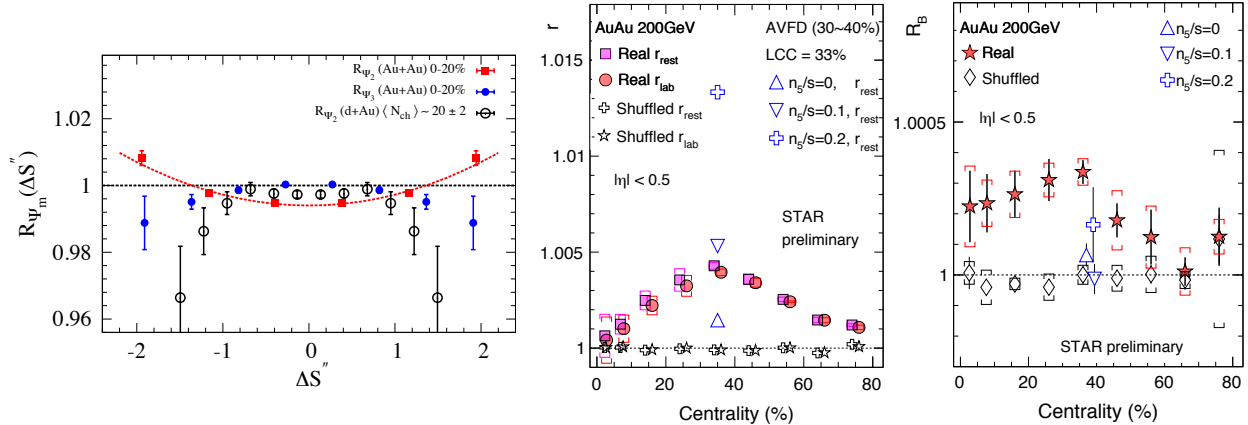


Figure 23: (Left) The R-observable shown for different collision systems, concave shape is consistent with CME expectation [92]. (Right) The two main quantities r and R_B derived from signed balance function, deviation from unity is consistent with CME expectation [93].

780 observation of Fig.23 indicates more concave shape for R_{Ψ_2} compared to R_{Ψ_3} in Au+Au
 781 whereas flat or convex shapes for $p/d + Au$ indicating that the measurements are consistent
 782 with expectations of CME [92]. For isobar collisions the case of CME will be confirmed if:
 783 1) a concave shape is observed for the ratio of the observables $R_{\Psi_2}(\Delta S)^{Ru+Ru}/R_{\Psi_2}(\Delta S)^{Zr+Zr}$
 784 and 2) the concavity should be weaker for $R_{\Psi_3}(\Delta S)^{Ru+Ru}/R_{\Psi_3}(\Delta S)^{Zr+Zr}$.

785 **Alternate measure: The signed Balance function** A very recently proposed observ-
 786 able to search for CME is the signed balance function (SBF) [95]. The idea is to account
 787 for the ordering of the momentum of charged pairs measured by the width of SBF that is
 788 expected to be different for out-of-plane as compared to in-plane measurement captured in
 789 the ratio r_{lab} . In addition, one can also account for the boost due to collective expansion
 790 of the system that forces all pairs to move in the same direction and measure the ratio in
 791 pair rest frame r_{rest} . In presence of CME the individual ratios as well as the double ratio
 792 $R_B = r_{rest}/r_{lab}$ is expected to be greater than unity. The preliminary measurements shown
 793 in Fig.23 (right) from STAR in Au+Au 200 GeV seems to be consistent with CME expect-
 794 ation. This observable will be studied with the isobar data in STAR but not as a part of
 795 the blind analysis and the CME expectation will be: 1) $r(Ru + Ru) > r(Zr + Zr)$, and 2)
 796 $R_B(Ru + Ru) > R_B(Zr + Zr)$.

797 **Prospect of CME search beyond isobar-era** It is important to discuss the strategy
 798 for CME search beyond the isobar-era. It is true that such strategy needs to be planned
 799 based on the outcome of the isobar program. We would like to get started by considering
 800 two possible scenarios at the top RHIC energy: 1) isobar program results in a significance
 801 of 3σ and below, 2) isobar program results in a significance of 3σ and above.

802 In the first scenario one can infer from the projection plot of fig.24(left) that the upper
 803 limit of the fraction of CME signal should be less than or equal to 8%. Question is under such

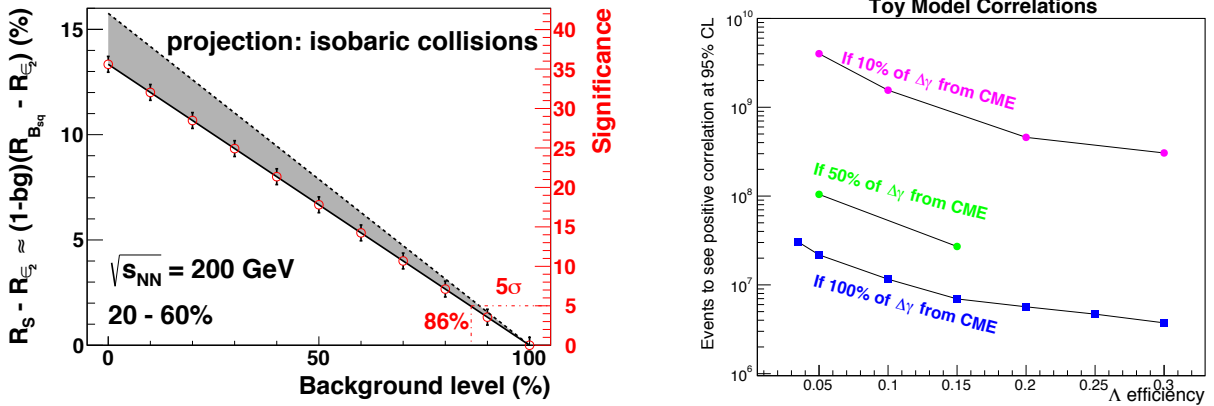


Figure 24: (Left) Projection plot taken from a previous beam user request document [83] indicating the anticipated significance in the measurement of charge separation as a function of the CME signal fraction. (Right) Estimation of the number of events required to see positive correlation between net Λ helicity with out-of-plane charge separation sensitive to local parity violation at 95% confidence level, plotted against the efficiency of $\Lambda(\bar{\Lambda})$ reconstruction (see [96] for details).

804 a scenario can STAR perform a follow up measurement to achieve a decisive 5σ significance
805 and establish a conclusive evidence of CME? It turns out such a measurement is possible
806 even with a single Au+Au 200 GeV data set during the year 2023 running of STAR concu-
807 rrently with sPHENIX. Current analysis of aforementioned CME signal in Au+Au 200 GeV
808 extraction using elliptic flow and charge separation with respect to spectator and participant
809 planes yields 4% statistical uncertainty with 2.4 B events ($2 - 3\sigma$ significance). In order to
810 get 5σ significance with the same analysis one needs to have statistical uncertainty of order
811 1.6% which would require about $(4/1.6)^2 \times 2.4 = 15$ Billion events. Therefore, as per the
812 previous estimates of anticipated 10 Billion events that can be collected by STAR during
813 the 2023 year, one can achieve about 4σ significance on the upper limit of possible CME
814 signal fraction in the measurement of charge separation. This estimate does not account
815 for two important facts that can lead to higher significance and a decisive measurement of
816 CME. The first one is that the magnitude of projected B-field on reaction plane is higher
817 in Au+Au collisions as compared to isobar collisions. The second one is that iTPC upgrade
818 enhances the charge particle multiplicity by 50% and therefore triplets ($\sim dN/d\eta^3$) (pairs
819 $\sim dN/d\eta^2$) statistics by a factor of 3.4 (2.3). So the final conclusion is that even if isobar
820 program results in a 3σ measurement running STAR in the year 2023 will result in a $> 4\sigma$
821 measurement. This conclusion assumes that the systematic uncertainty can be controlled to
822 be smaller than the statistical uncertainty, i.e. below 2%.

823 For the second scenario ($> 3\sigma$ measurement from isobar program) we will also be able
824 to establish an upper limit of the fraction of CME signal. For example, in fig.24(left) we see
825 that 5σ significance will establish 13% CME signal and a discovery of the CME phenomenon
826 in heavy ion collisions. The impact of such a discovery will set a milestone in physics.
827 Running STAR in the year 2023 concurrently with sPHENIX would be essential to perform

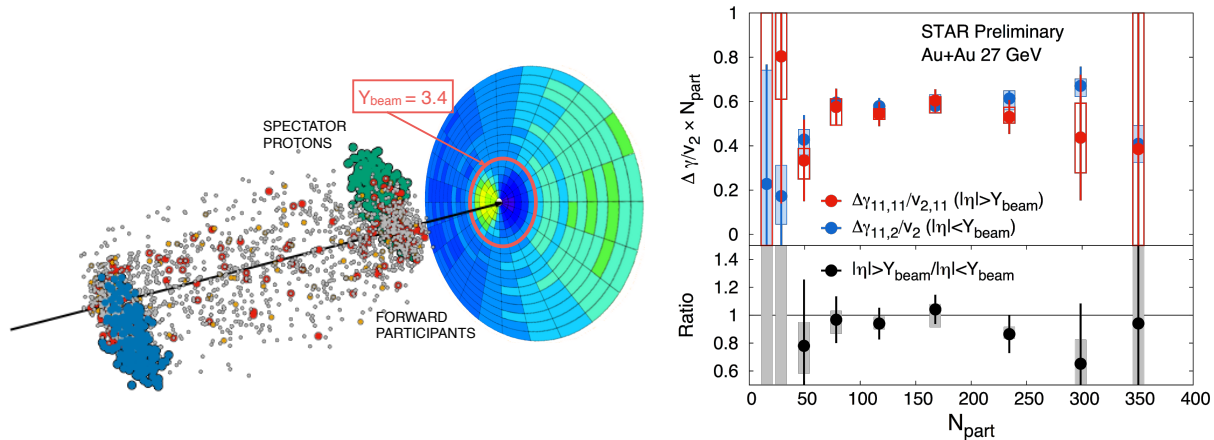


Figure 25: Prospect of precision CME search with the BES-II data. (Left) Single simulated UrQMD event and EPD detector acceptance that covers beam rapidity and detects both forward participants and spectators in 27 GeV Au+Au collisions that have large directed flow which changes sign at $\eta = Y_{\text{beam}} = 3.4$. (Right) γ -correlators scaled by v_2 across different event-planes and double ratio of spectators/participant event planes which should be unity for no-CME scenario.

828 dedicated precision measurements to further investigate and characterize the phenomena. In
 829 this context STAR collaboration has stated a new analysis to understand the origin of parity
 830 violation in hot QCD by measuring the correlation of net- Λ helicity with charge separation
 831 across reaction plane [96]. The difference between the number of positive and negative
 832 helicity $\Lambda(\bar{\Lambda})$ $N_L^\Lambda - N_R(\Lambda)$ should be associated with net-chirality, i.e. the difference between
 833 right and left handed quarks, in a given event. Since net chirality in the event also drives out-
 834 of-plane charge separation (a_1) in the presence of B-field, one expects a correlation between
 835 a_1 and $N_L^\Lambda - N_R(\Lambda)$ as a results of local parity violation. Currently available data sets do
 836 not allow us to perform a significant measurement for this observable. Using a toy model
 837 simulation, shown in fig.24(right), we estimate the number of event required to see non-
 838 zero correlations between a_1 and $N_L^\Lambda - N_R(\Lambda)$ at the 95% confidence level as a function of
 839 the efficiency of $\Lambda(\bar{\Lambda})$ reconstruction. Different curves correspond to different magnitudes
 840 of CME fraction in the measurement of γ -correlator. With about 10 B Au+Au 200 GeV
 841 events in run 2023 it will be possible to perform a significant measurement to study this
 842 phenomenon.

843 Regardless of the outcome of the measurements with the isobar program, that will be
 844 performed at the top RHIC energy, one question will remain. What happens at lower collision
 845 energy? In this context a new idea has emerged. The newly installed event-plane detector
 846 (EPD) upgrade provides a new capability at STAR towards CME search at lower collision
 847 energy and for the Beam Energy Scan phase-II program [18]. The idea is simple, at lower
 848 energies EPD acceptance ($2.1 < |\eta| < 5.1$) falls in the region of beam rapidity (Y_{beam}) and
 849 can measure the plane of strong directed flow (Ψ_1) of spectator protons, beam fragments
 850 and stopped protons, therefore strongly correlated to the B-field direction (See fig25). The
 851 next step is to measure $\Delta\gamma$ with respect to Ψ_1 and compare it with the measurement of

852 $\Delta\gamma$ along Ψ_2 planes from outer regions of EPD and TPC at mid-rapidity that are weakly
 853 correlated to the B-field directions. A test of CME scenario will be to see if large difference
 854 is observed in the measurements. First preliminary measurements from STAR as shown in
 855 Fig 25 is dominated by uncertainty but seems to show a lot of prospects for the CME search
 856 at lower energies. With higher statistics data from the BES-II program (7.7-19.6 GeV) and
 857 STAR fixed target run more precise measurement is possible.

858 1.3 Highlights from the Spin and Cold QCD Program

859 *Introduction*

860
 861 The goal of the STAR Cold QCD program is to probe the spin and flavor structure
 862 of the proton and understand the role of spin in Quantum Chromodynamics, exploiting
 863 the unique capability of RHIC to provide longitudinally and transversely polarized p+p
 864 and p+A collisions at multiple energies. Measurements with longitudinal beam polariza-
 865 tions have given new insights into the helicity structure of the proton, while measurements
 866 with transverse polarizations have provided new ways to probe polarized parton distribution
 867 functions (PDFs) in the collinear and transverse momentum dependent frameworks. Addi-
 868 tionally, the cross-section measurements in the unpolarized p+p collisions provide valuable
 869 information to constrain collinear and transverse momentum dependent unpolarized PDFs.
 870 This program is complemented by studies of polarized p+p elastic scattering and central
 871 exclusive production, in which a far-forward proton is detected intact.

Table 5: Summary of polarized p+p and p+A running periods at RHIC since 2009, including center-of-mass energy, STAR’s integrated luminosity and the average beam polarization for blue (B) and yellow (Y) beams from the H-jet polarimeter.

Year	System	\sqrt{s} (GeV)	Recorded Lumi. (pb^{-1})	Polarization	B/Y $\langle P \rangle$ (%)
2009	<i>pp</i>	200	25	Longitudinal	55/55
2009	<i>pp</i>	500	10	Longitudinal	39/39
2011	<i>pp</i>	500	12	Longitudinal	48/48
2011	<i>pp</i>	500	25	Transverse	48/48
2012	<i>pp</i>	200	22	Transverse	61/56
2012	<i>pp</i>	510	82	Longitudinal	50/53
2013	<i>pp</i>	510	300	Longitudinal	51/52
2015	<i>pp</i>	200	52	Transverse	53/57
2015	<i>pp</i>	200	52	Longitudinal	53/57
2015	<i>pAu</i>	200	0.45	Transverse	60/-
2015	<i>pAl</i>	200	1	Transverse	54/-
2017	<i>pp</i>	510	320	Transverse	55/55

872 Since 2009, RHIC STAR has completed several highly successful polarized p+p runs

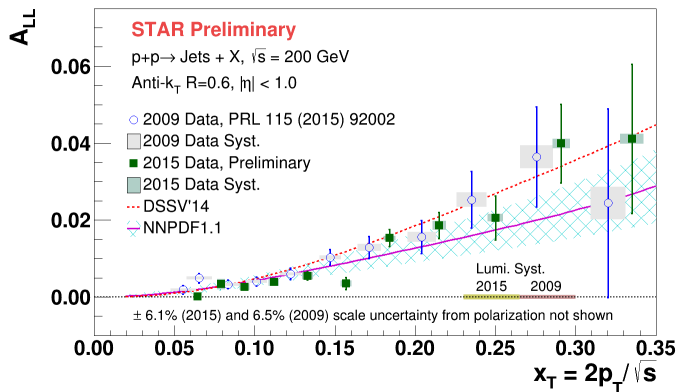


Figure 26: A_{LL} for inclusive jets with $|\eta| < 1.0$ versus x_T . The filled points show 2015 preliminary results [97], whereas the open points show the 2009 data of Ref. [99]. The bars show the size of the statistical uncertainties, whereas the boxes indicate the size of systematic uncertainties. The curves show the expected A_{LL} values for the DSSV14 [101] and NNPDF-pol1.1 [102] parton distributions.

873 both at $\sqrt{s} = 200$ GeV and $\sqrt{s} = 500/510$ GeV. Moreover, p+Au and p+Al datasets with
 874 a transversely polarized proton beam have been recorded in 2015 at $\sqrt{s} = 200$ to address
 875 important physics problems, including the ridge phenomenon and the possible onset of gluon
 876 saturation effects. Table 5 summarizes the STAR sampled luminosity and the luminosity
 877 averaged beam polarization as measured by the hydrogen jet (H-jet) polarimeter.

878

879 *Longitudinal program*

880

881 Since last year's PAC meeting, the STAR spin and cold QCD physics working group has
 882 released preliminary results focused on double-spin asymmetries A_{LL} of inclusive jet [97] and
 883 dijet [98] production in longitudinally polarized p+p collisions at a center-of-mass energy
 884 $\sqrt{s} = 200$ GeV based on the 2015 data set. These analyses are aimed at providing additional
 885 constraints to the gluon helicity distribution $\Delta G(x, Q^2)$, especially for the medium gluon
 886 momentum fractions in the range from $x \simeq 0.05$ to $x \simeq 0.5$. Figures 26 and 27 show the
 887 preliminary results of inclusive jet A_{LL} together with the 2009 data results of Ref. [99] and
 888 dijet A_{LL} together with the 2009 results from [100] respectively. Expected A_{LL} values for
 889 the DSSV14 [101] and NNPDF-pol1.1 [102] parton distributions are also presented. The
 890 results are in good agreement with previous measurements at $\sqrt{s} = 200$ GeV and with the
 891 theoretical evaluations of prior world data. They have better precision and thus provide
 892 further evidence that $\Delta G(x, Q^2)$ is positive for $x > 0.05$.

893 The results for the inclusive jet and dijet A_{LL} based on the 2012 $\sqrt{s} = 510$ GeV longi-
 894 tudinally polarized p+p data, which enabled exploration of $\Delta G(x, Q^2)$ down to $x \simeq 0.015$,
 895 were discussed in the previous PAC report and have since been published in Physical Review
 896 D [103].

897

898 *Transverse program*

899

900 There have been several new STAR preliminary results on transverse spin physics released
 901 since the last PAC meeting. Highlights include the Collins asymmetry for charged pions
 902 inside a jet [104] and the dijet Sivers asymmetry [105] at $\sqrt{s} = 200$ GeV p+p collisions.

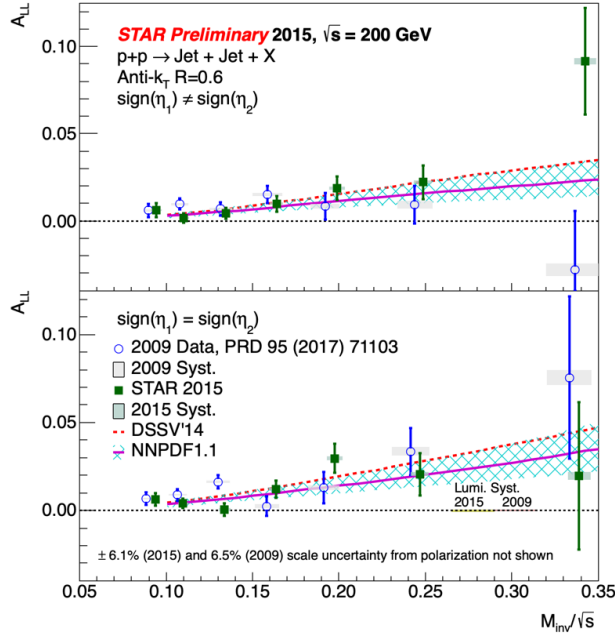


Figure 27: A_{LL} as a function of the parton-level invariant mass for dijets with the opposite-sign (top) and same-sign (bottom) event topologies. The filled points show 2015 preliminary results [98], whereas the open points show the 2009 data of Ref. [100]. The bars show the size of the statistical uncertainties, whereas the boxes indicate the size of systematic uncertainties. The curves show the expected A_{LL} values for the DSSV14 [101] and NNPDF-pol1.1 [102] parton distributions.

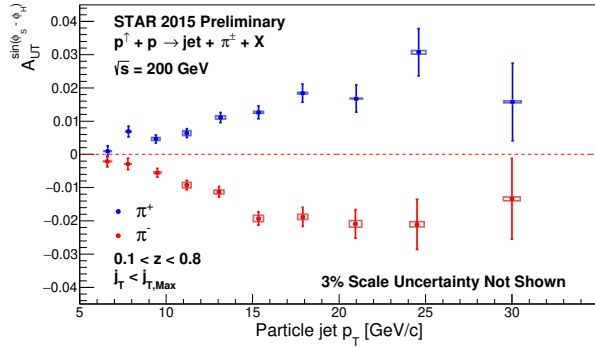


Figure 28: Preliminary results for the Collins asymmetry plotted for identified π^+ (blue) and π^- (red) particles as a function of jet p_T for jets that scatter forward to polarized beam ($x_F > 0$) [104]. The full range of both z and j_T are integrated over.

903 Moreover, the final publications of the transverse single spin asymmetries (TSSA) for neutral
 904 pions produced at forward rapidity in $\sqrt{s} = 200$ GeV for p+p, p+Au and p+Al [106], and
 905 500 GeV p+p [107] collisions are in God Parent Committees.

906 The Collins asymmetry in p+p collisions combines the collinear quark transversity in the
 907 proton with the transverse momentum dependent Collins fragmentation function [108–110],
 908 thus provides a cleaner probe of the Collins fragmentation functions than that in semi-
 909 inclusive deep inelastic scattering (SIDIS), and enables tests of evolution, universality and
 910 factorization breaking in the TMD formalism. Figure 28 shows the preliminary Collins
 911 asymmetries for charged pions inside jets that scatter forward ($x_F > 0$) to the polarized
 912 beam from 2015. The measured asymmetries are consistent with previous measurements
 913 from 2012 [111], but have 30% smaller statistical uncertainty.

914 The Sivers effect describes the correlation of the parton transverse momentum with the
 915 transverse spin of the nucleon. Figure 29 shows the first observation of non-zero Sivers
 916 asymmetries in dijet production of transversely polarized proton collisions using the STAR

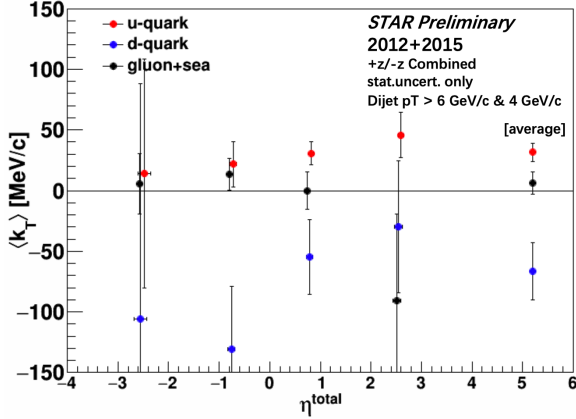


Figure 29: Preliminary results for the spin-dependent k_T values for u , d and $gluon + sea$ from the dijet Sivers measurement as a function of the sum of dijet pseudorapidities $\eta_1 + \eta_2 \sim \ln(\frac{x_1}{x_2})$ [105].

917 2012+2015 polarized p+p data. Compared to the previous 2006 result [112], fully recon-
 918 structed jets are analyzed with 33 times more statistics. Charge-tagging methods are em-
 919 ployed in order to separate the u and d quark signals. With detailed simulation, the individual
 920 parton spin-dependent $\langle k_T \rangle$ are extracted for u , d and $gluon + sea$ quarks, and indicates that
 921 $\langle k_T^u \rangle \approx 32$ MeV/c, $\langle k_T^d \rangle \approx -67$ MeV/c and $\langle k_T^{g+sea} \rangle \approx 0$ MeV/c.

922 The transverse single spin asymmetry (TSSA) for forward neutral pions produced in
 923 polarized proton collisions with protons (p+p), with aluminium nuclei (p+Al) and with gold
 924 nuclei (p+Au) at $\sqrt{s} = 200$ GeV from FMS are also measured using the data taken in
 925 2015. The preliminary results for (p+p) and (p+Au) have been released [106] and the final
 926 publication is soon to be submitted to Physical Review D. Measured asymmetries presented
 927 in Fig. 30 are found to rise with transverse momentum at $x_F < 0.5$, while they flatten or
 928 fall at larger x_F . The results are consistent with a weak nuclear A dependence. Moreover, a
 929 further observation is that TSSA is significantly larger for isolated π^0 s than for non-isolated
 930 π^0 s, which are accompanied by additional jet-like fragments.

931 The transverse single-spin asymmetry of neutral pions at $\sqrt{s} = 200$ GeV and 500 GeV
 932 from FMS data are compared in Fig. 31. The 200 GeV data are from 2015, while the 500
 933 GeV data are from 2011. The theoretical calculations presented in the plot are based on
 934 the Transverse Momentum Dependent (TMD) and collinear twist-3 functions from a recent
 935 global analysis [113], which also includes previous forward π^0 and charged hadron TSSA data
 936 from RHIC in the fit. The theoretical calculation differs from our measurement and only
 937 provides a reasonable description of the non-isolated π^0 in the low- x_F region. A continu-
 938 ous increase of the TSSA with Feynman- x indicates the independence on the center-of-mass
 939 energy. Pions with no nearby particles, which may not arise from conventional jet fragmen-
 940 tation, tend to have a higher TSSA than non-isolated pions, which suggests that a different
 941 mechanism other than the Sivers or Collins effects is required to explain these results.

942 *Unpolarized Results*

943
 944
 945 The azimuthal correlation of forward di-pions produced in p+p and p+Au collisions pro-
 946 vides an essential tool to access the underlying gluon dynamics in the nonlinear evolution

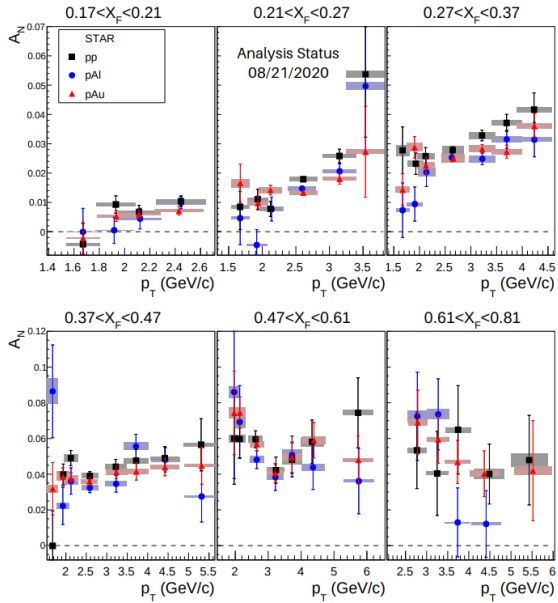


Figure 30: Transverse single spin asymmetry for forward π_0 production as a function of transverse momentum for six Feynman x_T regions. Results for three collisions systems are shown, black squares for p+p, blue circles for p+Al and red triangles for p+Au collisions. The statistical uncertainties are shown with vertical error bars and the filled boxes indicate the horizontal and vertical systematic uncertainties. Analysis status on 08/21/2020. Preliminary results available in Ref. [106].

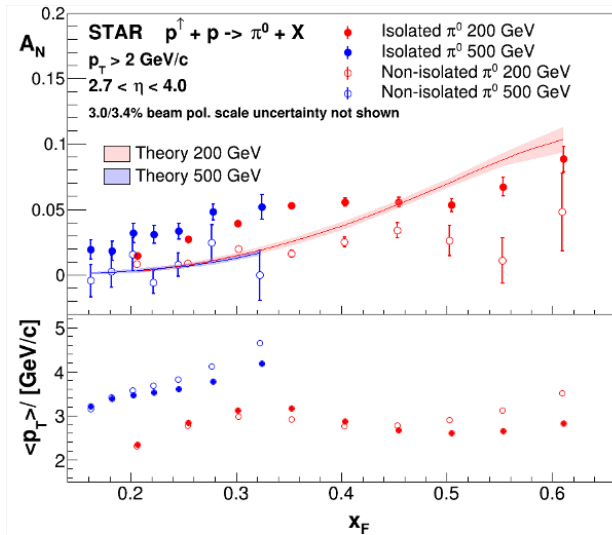


Figure 31: Preliminary results for the transverse single-spin asymmetry as function of Feynman- x for the isolated and non-isolated π^0 in transversely polarized proton-proton collisions at 200 and 500 GeV [107]. Theory curves based on a recent global fit [113] are also shown. The average transverse momentum of the π^0 for each x_F bin is shown in the lower panel.

947 region. π^0 measured in the FMS in the pseudorapidity region $2.5 < \eta < 4.0$ probe low
 948 momentum fraction partons down to $x \approx 0.001$ at $\sqrt{s} = 200$ GeV, which are dominated
 949 by gluons. 2015 p+Au collisions have a unique opportunity to study this phenomenon with
 950 much higher luminosities and smaller background than 2008 d+Au [114]. Figure 32 shows
 951 the status of di-pion correlation measurement from Run15 p+p and p+Au collisions. The
 952 away-side peak is suppressed in high activity p+Au collisions compared with p+p. This
 953 effect is more significant when the more central part of the nucleus is probed (with higher
 954 multiplicity as indicated by BBCE). Further analysis to characterize the p_T dependence and
 955 compare with theoretical expectations is ongoing.

956

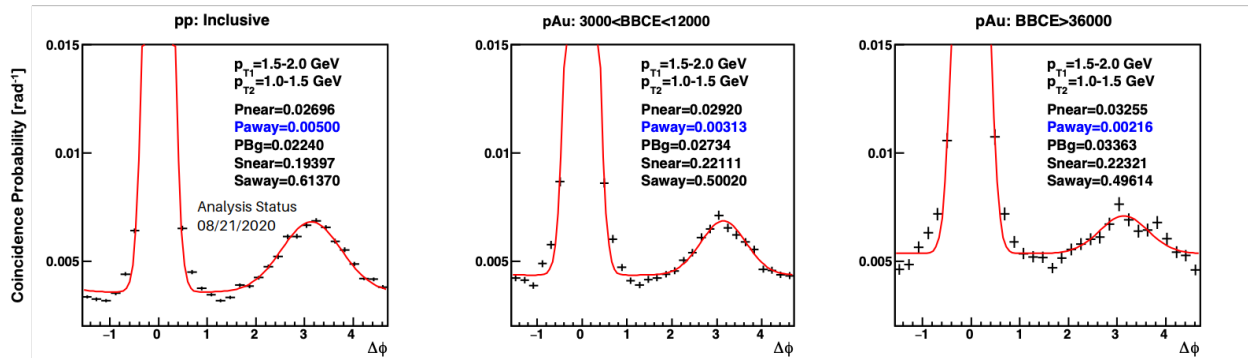


Figure 32: Coincidence probability as a function of azimuthal angle difference between two forward neutral pions in p+p, peripheral and central p+Au collisions. Analysis status on 08/21/2020.

957 The STAR measurement of the unpolarized cross-section ratio of the W^+ and W^- bosons
 958 from the STAR 2011 to 2013 data at $\sqrt{s} = 500/510$ GeV has released preliminary results [115]
 959 and is soon to be submitted to Physical Review D. Figure 33 shows the ratio plotted as a
 960 function of lepton pseudorapidity. This unique measurement is sensitive to the unpolarized
 961 \bar{d}/\bar{u} quark distribution and will provide insights into unpolarized light quark distributions
 962 $\bar{d}(x)$ and $\bar{u}(x)$ at $x > 0.05$. The measurement at STAR is complementary to the Drell-Yan
 963 results from NuSea [116] and SeaQuest [117], covering the overlapping x region of about
 964 $0.1 - 0.3$ at higher $Q^2 = M_W^2$.

965 Differential cross sections of Z^0 -boson production as function of transverse momentum
 966 are valuable input to global fits of TMD parton distribution functions, and STAR kinematics
 967 ($0.1 < x < 0.3$) are complementary to LHC and Tevatron data. Figure 34 shows prelimi-
 968 nary results from 2011-2013 data with an integrated luminosity of 350 pb^{-1} [118]. Data on
 969 disk from 2017 comprise about the same luminosity, and preliminary results are expected
 970 soon along with transverse single-spin asymmetries. While the measurement of Z^0 -bosons is
 971 an experimentally very clean observable, it requires a good understanding of the calorime-
 972 ter performance. These will inform the background studies of the measurements of Sivers
 973 asymmetries for W -bosons, which are also expected very soon.

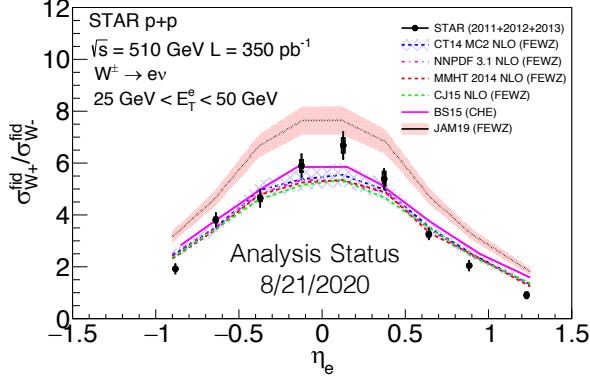


Figure 33: W^+ and W^- cross-section ratio as a function of lepton pseudorapidity for the combined 2011, 2012, and 2013 datasets. The central values correspond to the mean value of η_e distribution for that bin. The error bars represent the statistical uncertainty, whereas the rectangular boxes represent the systematic uncertainty for the respective data point. These measurements are compared to various theory frameworks, which use several different PDF inputs. Analysis status on 08/21/2020. Preliminary results available in Ref. [115].

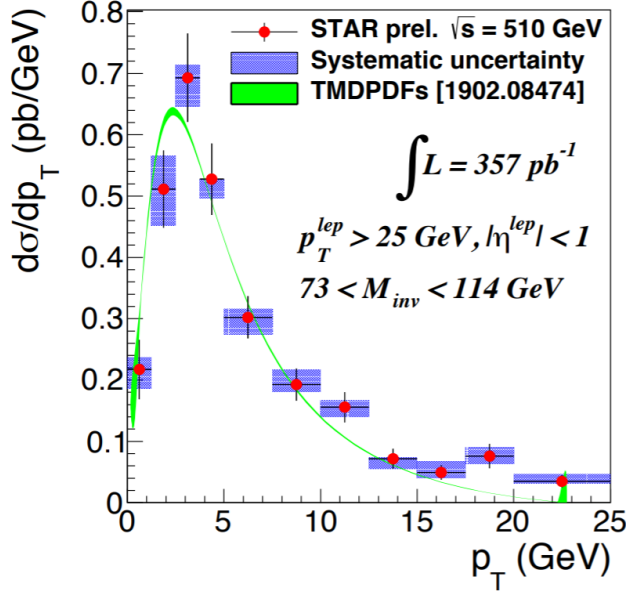


Figure 34: Preliminary results for the differential cross-section of Z^0 -bosons as function of transverse momentum p_T [118] and comparison with theory predictions based on calculations developed in [119]. Results are based on data from 2011-2013.

1.4 Run-20 Performance

In this section, we will review the BES-II collider and fixed-target performance to date. Careful study of these performance metrics will be used to make projections about the required time to complete the 7.7 GeV collider system in run 21.

The BES-II collider program performance is over-viewed in table 6. The 27 GeV system, which was run in 2018, was not officially part of the BES-II physics program, however it is close enough in energy to help provide some performance evaluation. The most important lines in the table from the point of view of performance evaluation are the *good event rate*, which is a measure of the useful luminosity, and the *data hours per day*. In general, we had seen improvements over the luminosities recorded in 2010/2011 of a factor of three to four. For the 27 GeV system, which was run in 2018 we saw the good event rate rise to 620 Hz, which implies a luminosity increase of factor of 3.3 over the 2011 performance.

The 19.6 GeV system was completed in 2019. For this system, the good event rate rose from 100 Hz in 2011 to 400 Hz in 2019 for a factor of 4.0 increase. We should note that it took 5.1 calendar weeks to complete the energy, however during the running period for the 19.6 GeV system, the facility was dedicating two twelve hour shifts per week to LEReC development. Correcting the 5.1 calendar weeks by 6/7 means that 4.4 beam weeks were used to complete this energy which should be compared to the 4.5 weeks which was requested in the STAR BUR for 2019. Historically, it has been shown that the luminosities scale with γ^2 above injection energy (9.8 GeV) and with γ^3 below injection energy. Scaling the 27 GeV performance would have predicted a good event rate of 330 Hz at 19.6 GeV. One should also note that the number of events recorded exceeded the required number significantly. Overall, the performance for the 19.6 GeV system significantly exceeded expectations.

The 14.6 GeV system was completed in 2019. This energy had been run previously in 2014, however the STAR good event rate was unusually low at that time so comparing the 23 Hz rate from 2014 to the 170 Hz rate in 2019 is not a good metric for performance. It required 8.6 calendar weeks to complete the required number of events, however during the running of the 14.6 GeV system, 40% of the beam time was used for LEReC development. Scaling the 8.6 calendar weeks by 60% yields effectively 5.1 weeks of beam time which favorably compares with the 5.5 weeks estimated in the BUR for 2019. Using the performance at 19.6 GeV (good event rate of 400 Hz), and the γ^3 scaling, we would have expected a good event rate at 14.6 GeV of 165 Hz. This compares well the the 170 Hz rate which was achieved for this energy. Performance at this energy slightly exceeded expectations.

The 11.5 GeV system was completed in 2020. The good event rate rose from 30 Hz in 2010 to 80 Hz in 2020 for a factor of 2.67 increase. It took 8.9 calendar weeks achieve the required event statistics. In the BUR for 2020 a range from 7.5 (optimistic) to 10 (pessimistic) was proposed. The actual time required fell in the middle of the expected range. Scaling the good event rate from 14.6 GeV by γ^3 predicted that the good event rate for 11.5 GeV would be 83 Hz, which compares favorably with the 80 Hz actually achieved for the run. Overall, performance at this energy met expectations.

The 9.2 GeV system will be completed in 2020 and is far enough along that we can project to completion. This energy was not run during BES-I, so there is not a historical

1016 comparison. Scaling the good event rate from 11.5 GeV by γ^3 predicts that we should have
1017 seen a good event rate of 38 Hz. Prior to the shutdown of the laboratory in March, we had
1018 achieved an average good event rate of 38 Hz and we had been averaging 16 hours of data
1019 taking per day, however the challenges of running in the summer have reduced the average
1020 number of hours of data taking to 13 and the average good event rate to 33. It is projected
1021 that this system will take 14.0 weeks to achieve the required event statistics, which is at the
1022 high end of the range that was included in the BUR for 2020.

1023 Quality assurance studies of the BES-II and FXT data indicate that roughly 98% of the
1024 data recorded will ultimately be used in physics analyses. The quality assurance takes place
1025 on multiple levels. At the time of data acquisition, online performance plots are reviewed
1026 as each run starts by the shift crew member. There are two levels of online plots; the first
1027 use the raw detector specific data to overview the performance of all systems; the second
1028 level does event-by-events tracking and vertex reconstruction using the High Level Trigger
1029 (HLT) computer farm to generate event level performance plots and to tag the good Au+Au
1030 collision events. The next level of quality assurance uses a FastOffline production of a small
1031 percentage of all recorded events. The plots generated by this review take place on a daily
1032 basis a provide the opportunity for corrections to any issues that might arise. A third level
1033 of quality assurance takes place in a weekly QA meeting which reviews the overlap between
1034 the events flagged as good from the HLT system, a significant fraction of which are recorded
1035 and available to preliminary offline physics analysis, and the events identified as good using
1036 the FastOffline processed data, which utilizes a more sophisticated tracking algorithm. The
1037 overlap of good events has been at the 98% level for all BES-II collider and FXT systems.
1038 The QA meeting also reviews preliminary physics working group quality assurance analysis
1039 of the FastOffline data sets. The final level of quality assurance comes from preliminary
1040 physics analyses using the FastOffline and the ExpressStream data sets. This multi-level to
1041 quality assurance guarantees that the data will meet the needs of the physics analyses for
1042 the BES-II science program.

1043 The relevant data sets recorded in 2018 have been fully calibrated and produced. These
1044 data sets are the 27 GeV collider system and the 3.0 and 7.2 GeV FXT data sets. Preliminary
1045 results for all of the key physics analyses have been performed and highlights of these new
1046 results are reviewed in the previous section of this document. The 2019 data sets have
1047 required extensive calibrations of the new detector systems, the iTPC and the eTOF. The
1048 large volume of cosmic ray data that were recorded have been used to do the fine spatial
1049 alignment of the new iTPC modules. New methods needed to be developed to calibrate
1050 the precise start timing for each event, which is needed to get the correct reconstruction of
1051 the z -location of the hits in the TPC to account for the long bunches used to maximize the
1052 luminosity for BES-II and the FXT programs. The complete set of calibrations for the first
1053 collider energy from 2019, the 19.6 GeV system, have been completed and a test production
1054 using these calibrations has been generated and are undergoing.

1055 As several of the physics opportunities discussed in the following sub-sections utilize fixed-
1056 target systems, it is best to review the performance in this mode of operation. An overview of
1057 the performance for all fixed-target energies is shown in table 7. The first fixed-target physics

Table 6: Achieved and projected experiment performance criteria for the BES-II collider program.

Collision Energy (GeV)	7.7	9.2	11.5	14.6	17.1	19.6	27
Performance in BES-I	2010	NA	2010	2014	NA	2011	2011
Good Events (M)	4.3	NA	11.7	12.6	NA	36	70
Days running	19	NA	10	21	NA	9	8
Data Hours per day	11	NA	12	10	NA	9	10
Fill Length (min)	10	NA	20	60	NA	30	60
Good Event Rate (Hz)	7	NA	30	23	NA	100	190
Max DAQ Rate (Hz)	80	NA	140	1000	NA	500	1200
Performance in BES-II (achieved)	2021	2020	2020	2019	2021	2019	2018
Required Number of Events	100	160	230	300	250	400	NA
Achieved Number of Events	2.9	160	235	324	TBD	582	560
fill length (min)	20-45	45	25	45	50	60	120
Good Event Rate (Hz)	16-24	33	80	170	265	400	620
Max DAQ rate (Hz)	400	700	550	800	1300	1800	2200
Data Hours per day	12-15	13	13	9	15	10	9
Projected number of weeks	11-20	8.5-14	7.6-10	5.5	2.5	4.5	NA
weeks to reach goals	TBD	14.0	8.9	8.6	TBD	5.1	4.0

1058 run was in 2018 using a 3.85 GeV beam. A total of three and half days was spent on this
 1059 system; first developing the conduct of operation and then recording a robust data sample.
 1060 The lowest energy beam was selected for this first run in 2018 because at that time the
 1061 iTPC and eTOF upgrades were not yet available; the lowest beam energy means the lowest
 1062 center-of-mass boost, which meant that we could still complete the physics program even
 1063 without the detector upgrades. Additionally in 2018, fixed-target data were recorded with
 1064 a single beam energy of 26.5 GeV. Obviously, at such a high energy the detector upgrades
 1065 would be essential for the mid-rapidity physics program. However, the 26.5 GeV beam
 1066 was not requested by STAR; this beam was being using by the Coherent Electron Cooling
 1067 program, and STAR was simply taking these data parasitically. This parasitic data taking
 1068 gave us further opportunities to refine the fixed-target conduct of operations, which gave us
 1069 confidence going forward that we could average 100 M good events per day in fixed-target
 1070 mode. This is limited by the STAR data acquisition system and not by RHIC.

1071 In 2019, eTOF detector upgrade system suffered damage at the start of the 14.6 GeV
 1072 collider system. This meant that it would be unavailable for any fixed-target energies taken
 1073 that year. It was felt that the physics program could still be achieved using the 4.59 GeV
 1074 beam, but that for all higher energies the loss of the eTOF system would compromise the
 1075 physics, so only modest samples at 5.75 and 31.2 GeV were taken.

1076 The eTOF detector was repaired for 2020, and relatively early in the run it was decided
 1077 to spend one week cycling through the seven remaining fixed-target energies. Roughly one

1078 day was spent at each energy. The conclusion from this series of fixed target energies is
 1079 that the collider and the experiment can quickly and efficiently set up and run fixed-target
 1080 systems. STAR can efficiently trigger on good fixed-target events with roughly 80-90% of
 1081 triggers passing the HLTgood test. The operators monitor the STAR event rate to keep the
 1082 current on target at a level to keep the STAR DAQ system running at full capacity and
 1083 minimizing the pile-up of multiple collisions in the target. Stores last for many hours (8-24
 1084 hours) and refill and realignment are fast and efficient.

1085 Preliminary physics results from the 3.0 and 7.2 GeV data sets recorded in 2018 are
 1086 available and highlights have been shown the the previous sections. Internal preliminary
 1087 physics analyses of the ExpressStream and FastOffline data sets have been preformed and
 1088 these confirm the quality of the data taken.

Table 7: Event statistics (in millions) needed in the fixed-target part of the BES-II program for various observables, and the total number of events acquired (those events taken in 2018 did not include the iTPC or eTOF detectors; those taken in 2019 did not include the eTOF).

$\sqrt{s_{NN}}$ (GeV)	3.0	3.2	3.5	3.9	4.5	5.2	6.2	7.2	7.7
Beam Energy	3.85	4.59	5.75	7.3	9.8	13.5	19.5	26.5	31.2
μ_B (MeV)	721	699	666	633	589	541	487	443	420
Rapidity y_{CM}	1.06	1.13	1.25	1.37	1.52	1.68	1.87	2.02	2.10
Observables									
Elliptic Flow	300	150	80	40	20	40	60	70	80
CME	70	60	50	50	50	70	80	90	100
Directed Flow	20	30	35	45	50	60	70	80	90
Femtoscopy	60	50	40	50	65	70	80	90	100
Kurtosis	36	50	75	125	200	400	950	NA	NA
Strange hadrons	300	100	60	40	25	30	50	75	100
Hypertritons	200	100	80	50	50	60	70	85	100
Event Totals									
Good events (2018)	258							158	
Good events (2019)	3.7	200	53						50
Good events (2020)			116	117	108	103	118	TBD	112

2 Proposed Program - Hot QCD in Run-21, 23, and 25

2.1 Beam Request for Run-21

2.1.1 Completion of the BES-II Program

The highest priority for Run-21 is the completion of the proposed BES-II program. At this time, the only system that remains to be taken is the 7.7 GeV collider data set. This energy is extremely important for several reasons. First, theoretical calculations suggest that the highest baryon density is achieved in collisions at this energy; second, several of the BES-I experimental signatures which have been put forth to be sensitive to the presence of deconfined matter either lose significance or are no longer present at this energy; third, the BES-I data showed enhanced fluctuations at this energy; finally, this energy provides the best acceptance overlap with the fixed-target program. Although the 7.7 GeV collider data set is extremely important from the point of view of the science, it is also technically the most challenging data set. The technical challenge of achieving a viable collision rate at this energy was the motivation to develop the Low Energy RHIC electron Cooling (LEReC) and is the reason that this energy has been left to the final year of the program.

The specific physics goals (are required statistics) include: measurement of the elliptic flow of the phi meson for which the the constituent quark scaling was suggested to break down in the lowest energy BES-I data (80 M events required); measurement of the correlators associated with the charge separation induced by the chiral magnetic effect which were seen to collapse at the lowest BES-I energies (50 M events required); differential measurements of the directed flow of protons which was seen to show evidence of a softening of the equation of state in the lowest BES-I data (20 M events required); Azimuthal femtoscopy measurements of protons to study the tilt angle of the source (35 M events required); measurement of the net-proton kurtosis which showed significant enhanced fluctuations at 7.7 GeV in the BES-I data (70 M events required); measurements of the di-lepton invariant mass distributions to determine in the excess in the low mass region is proportional to the total baryon density (100 M events required); and the global lambda polarization to determine the magnetic field significance (50 M events required). These analyses are being pursued at all of the BES-II collider energies; for several of the physics measurements, the 7.7 GeV energy is expected to be either the most significant or the most challenging.

The 7.7 GeV collider system provides the essential bridge between the collider and fixed-target energy scans. Although in later sections we detail a request to acquire fixed-target data at higher overlap energies, there is the largest region of common coverage at this energy. This will provide critical cross checks between the different modes.

Although the 7.7 GeV collider system is the most technically challenging system of the suite of BES-II and FXT energies, one can use the performances which have already been achieved during the BES-II program to help develop projections for the 7.7 GeV collider energy. These BES-II performance trends have been detailed in 6. In 2010, STAR achieved a good event rate of 7 Hz; a factor of three improvement would result in a 21 Hz rate. Scaling the performance at 9.2 GeV by γ^3 would predict a good event rate of 19.3 Hz. We project the

1129 good event rate to fall between 16 and 24 Hz. We project the range of hours of data taking
 1130 per day to fall between 12 and 15. These numbers suggest a range in the expected number
 1131 of weeks to reach the goals from 11 to 20 weeks. We should note that CAD has provided
 1132 projections which suggest that it will take 28 weeks to reach the goals. Our projections are
 1133 more optimistic. Although we recognize that it is likely that running the 7.7 system will
 1134 require all the available beam time in 2021, the optimistic range of our predictions suggests
 1135 that we should prepare for success and we have therefore considered and prioritized other
 1136 programs which could be run in 2021 if time were to be available.

1137 **2.1.2 Au+Au Collisions in FXT Mode at $\sqrt{s_{NN}} = 3.0$ GeV - I: 300 million goal**

1138 QCD matter at high baryon chemical potential region contains a wealth of unexplored physics
 1139 and is one the central focus of current and future heavy-ion collision programs in few GeV
 1140 energy range around the world. RHIC has been able to deliver beams with the energy as
 1141 low as 3.85 GeV per nucleon. Utilizing the gold fixed target (FXT) installed in the STAR
 1142 experiment, we were able to record collision events at the center-of-mass-energy as low as
 1143 $\sqrt{s_{NN}} = 3.0$ GeV, which corresponds to baryon chemical potential of $\mu_B \sim 720$ MeV in
 1144 central collisions. STAR detector configuration (including the iTPC and eTOF) has the
 1145 full midrapidity coverage ($|y| < 0.5$) at this energy and enables us to carry a systematic
 1146 investigation of the dynamics of the QCD matter created in these collisions at $\sqrt{s_{NN}}$ from
 1147 3.0 up to 200 GeV.

1148 At such a high μ_B region and moderate temperatures, baryon dynamics become important
 1149 or even dominant in understanding the QCD matter properties. Strange quarks, due to their
 1150 heavier masses, play an important role in study the high net-baryon density QCD matter.
 1151 The combination of increased sensitivity of strange quarks with the existing high baryon
 1152 density in low energy heavy-ion collisions offers a unique condition to create various light
 1153 hypernuclei, which enables us to study *e.g.* the hyperon-nucleon ($Y-N$) interactions, which
 1154 have potential implications for the inner structure of compact stars in nuclear astrophysics.

1155 STAR has collected ~ 250 million FXT Au+Au events at $\sqrt{s_{NN}} = 3.0$ GeV in 2018,
 1156 however without the iTPC and eTOF. We propose to collect a minimum of 300 million
 1157 events with the extended phase-space coverage enabled by iTPC and eTOF for the following
 1158 measurements:

- 1159 • high moments of proton multiplicity distributions covering the same midrapidity ac-
 1160 ceptance $|y| < 0.5$, $0.4 < p_T < 2.0$ GeV/ c , comparable to that with the BES-I and BES-II
 1161 measurements in collider mode.
- 1162 • precision ϕ meson production at midrapidity to test the validity of Canonical Ensemble
 1163 (CE) for strangeness production at high baryon density region.
- 1164 • systematic measurements of lifetime, binding energy, production yield, collective flow
 1165 of light hypernuclei (${}^3_{\Lambda}\text{H}$, ${}^4_{\Lambda}\text{H}$, ${}^5_{\Lambda}\text{He}$ etc.). as well as towards a discovery of double- Λ
 1166 hypernuclei.

- measurement of low- and intermediate-mass dileptons to extract fireball lifetime, its average temperature and to access the microscopic properties of matter. This would be the first measurement of electromagnetic radiation at this energy which will guide the future high μ_B facilities at FAIR, NICA.

With additional beam time allowed, we would like to further collect up to 2 billion Au+Au FXT events at $\sqrt{s_{NN}} = 3.0$ GeV which will be elaborated in the next section.

One feature we would like to point out is that the single beam energy for FXT collisions at $\sqrt{s_{NN}} = 3.0$ GeV is 3.85 GeV per nucleon, the same beam energy to be used for colliding to collect the major 7.7 GeV collision dataset in year 2021. This leads to a negligible transition time for operation between $\sqrt{s_{NN}} = 7.7$ GeV collider mode and $\sqrt{s_{NN}} = 3.0$ GeV FXT mode.

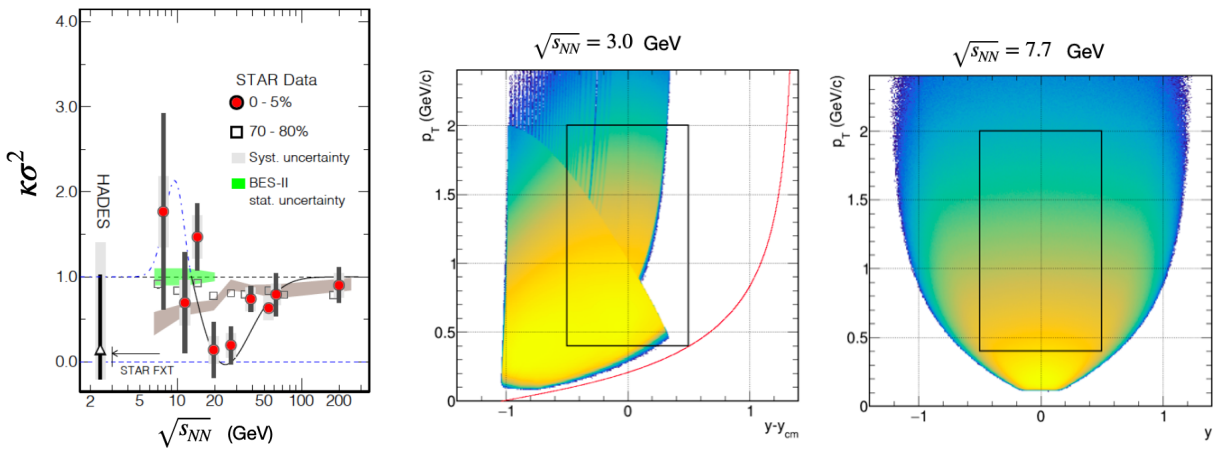


Figure 35: (Left) The net-proton $\kappa\sigma^2$ in most central (0-5%) and peripheral (70-80%) Au+Au collisions as a function of collision energy. (Middle/Right) Proton acceptance plot p_T vs. y in the center-of-mass frame at $\sqrt{s_{NN}} = 3.0$ GeV (FXT data from 2018) and 7.7 GeV (collider data from 2010), respectively. The red curve in the middle panel indicates the acceptance boundary with iTPC and eTOF.

High moments of proton multiplicity distributions A non-monotonic behavior of net-proton high moments $\kappa\sigma^2$ as a function of collision energy has been suggested to be an evidence of the existence of QCD critical point [120, 121]. Figure 35 (left panel) shows the final STAR measurement from the BES-I data as a function of energy exhibiting a suggestive non-monotonic behavior [122, 123]. A complete picture of the non-monotonic behavior requires measurements at collision energies below the lowest collider mode energy (7.7 GeV) by utilizing the FXT mode collisions. STAR detector configuration has the best midrapidity coverage for fixed target collisions at the lowest collision energy $\sqrt{s_{NN}} = 3.0$ GeV. Figure 35 middle and right panels show the proton acceptance with TPC and barrel TOF in 2018 FXT data at 3.0 GeV and 2010 collider data at 7.7 GeV, respectively. In the 2018 FXT data, to ensure $> 95\%$ purity of the proton sample, one needs to utilize the barrel TOF for high momentum particle identification. With this requirement, the proton acceptance in

1189 2018 covers full negative rapidity region ($-0.5 < y < 0$, $0.4 < p_T < 2.0$ GeV/c), while missing a
 1190 considerable acceptance in the positive rapidity region. A new run, with eTOF and iTPC,
 1191 would allow for phase space coverage comparable to the one in collider mode (indicated by
 1192 the box in the middle panel). The estimated acceptance boundary for protons is indicated
 1193 by the red line shown in Fig. 35 middle panel. We can therefore cover the full midrapidity
 1194 $|y| < 0.5$ region from $0.4 < p_T < 2.0$ GeV/c which will be the same as these measurements
 1195 conducted in collider mode data, shown in the right panel. This would allow to perform
 1196 a systematic scan of the net-proton high moments analysis within the same mid-rapidity
 1197 acceptance across the collision energy from 3.0 up to 200 GeV.

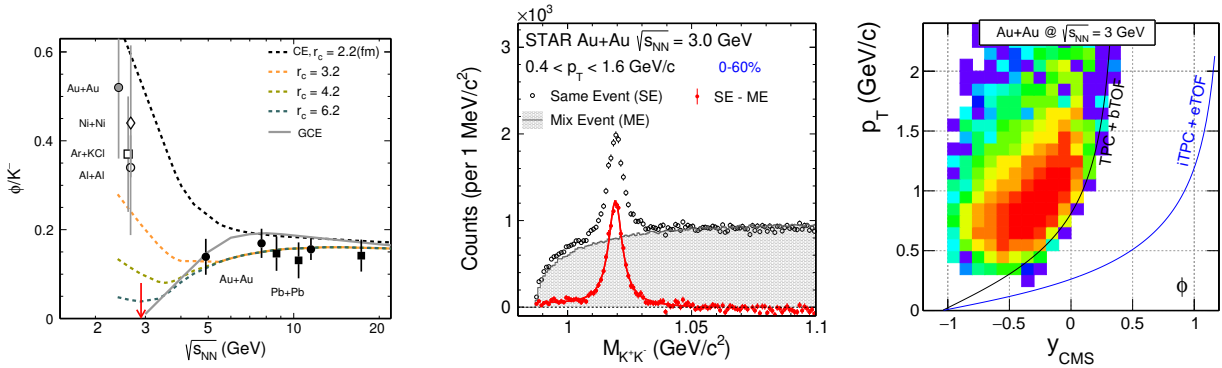


Figure 36: (Left) ϕ/K^- ratio as a function of collision energy from several heavy-ion experiments in comparison to thermal model calculations assuming strangeness following GCE and CE with different canonical radius. (Middle) Invariant mass distributions of K^+K^- pairs and the ϕ meson signal in 2018 FXT data at $\sqrt{s_{NN}}=3.0$ GeV. (Right) Reconstructed ϕ meson candidate phase space distributions using 2018 FXT data taken at $\sqrt{s_{NN}}=3.0$ GeV. The black line shows the boundary of combining the TPC and barrel TOF detector for kaon identification. The blue line indicates the anticipated boundary extended by iTPC and eTOF for kaon identification in the proposed 2021 FXT run at $\sqrt{s_{NN}}=3.0$ GeV.

1198 **ϕ meson production** Yields of strange hadron produced in relativistic heavy-ion col-
 1199 lisions from RHIC BES-I energies up to the LHC energy ($\sqrt{s_{NN}}=7.7-5500$ GeV) can be
 1200 well described by thermal model with Grand Canonical Ensemble (GCE) in which strange
 1201 quark number is conserved on average [124–127]. It has been argued that at low energy
 1202 heavy-ion collisions when the fireball created in these collisions becomes small enough the
 1203 GCE for strange quarks will break down. Strangeness needs to be conserved on the event-
 1204 by-event basis, therefore only Canonical Ensemble (CE) is applicable to strange hadron
 1205 production [125, 127]. Strange hadrons with finite strangeness number (e.g. K , Λ etc.) will
 1206 suffer from a suppression due to the strangeness number conservation, often characterized
 1207 by a canonical radius (r_c) for strange quark profile in comparison to the regular radius (r)
 1208 for light quarks [128, 129]. The ϕ meson is the lightest bound state of s and \bar{s} quarks with
 1209 zero net-strangeness number. Its production yield, on the contrary, will not suffer from the
 1210 canonical suppression. Therefore CE models predict the ϕ/K^- ratio will show an enhance-

1211 ment in very low energy heavy-ion collisions while GCE models calculate the ϕ/K^- ratio will
1212 gradually drop to zero at the ϕ production threshold in pp collisions. ($\sqrt{s_{\text{NN}}} = 2.89$ GeV).

1213 Experimentally, the measured ϕ/K^- values stay around 0.15 at $\sqrt{s_{\text{NN}}} > 5$ GeV up to
1214 the LHC energy. At collision energies below the ϕ production threshold in pp collisions,
1215 measurements from HADES and FOPI suggest an enhancement compared to those at high
1216 energies, consistent with the CE description for strange quarks at such low energies within
1217 appreciable uncertainties [130, 131]. High precision measurement of the ϕ/K^- at such low
1218 energies will be of great interest to systematically investigate the ϕ meson and strangeness
1219 production mechanism in heavy-ion collisions.

1220 We have performed such a measurement using the FXT data at $\sqrt{s_{\text{NN}}} = 3$ GeV taken
1221 in 2018. Fig. 36 middle panel shows the reconstructed K^+K^- invariant mass distributions
1222 in 0-60% centrality. The shaded histogram shows the K^+K^- pair distributions from the
1223 mixed-event technique while normalized at the mass region above the ϕ meson signal. The
1224 red data points show the mixed-event background subtracted distributions and the ϕ meson
1225 signal obtained in this data is about 60σ . The right panel shows the ϕ meson acceptance
1226 coverage in center-of-mass frame. Due to the small production yield of kaons, one needs to
1227 rely on clean particle identification using TOF detector to obtain a control background in
1228 the ϕ meson reconstruction. The black curve indicates the single track acceptance boundary
1229 from TPC and barrel TOF in 2018 year run. One can see the ϕ meson p_T acceptance at
1230 midrapidity is limited at ~ 0.6 - 0.8 GeV/ c . This covers roughly only 40% of the ϕ meson yield
1231 in the full p_T region, leading to a considerable amount of systematic uncertainty due to the
1232 p_T extrapolation. The blue curve in the same panel indicates the anticipated single track
1233 boundary with iTPC and eTOF. The p_T lower limit can be extended down to ~ 0.2 GeV/ c ,
1234 yielding a p_T coverage of $\sim 90\%$ of total dN/dy at midrapidity. This will greatly reduce the
1235 systematic uncertainty in the total ϕ meson yield measurement.

1236 We therefore request to take the FXT data at $\sqrt{s_{\text{NN}}} = 3$ GeV with iTPC and eTOF
1237 detectors in RHIC 2021 year run. A roughly similar amount of statistics (300 million) will
1238 allow us to perform the measurement of ϕ/K^- ratio with high precision both statistically
1239 and systematically.

1240 **Hypernuclei production and potential discovery for double- Λ hypernuclei** Hy-
1241 pernuclei are those nuclei with one or more nucleons replaced with hyperons (typically Λ s).
1242 The study of hypernuclei lifetime, binding energy and their production mechanism offer in-
1243 sights to the understanding of hyperon-nucleon (Y - N) interactions. The Y - N interactions
1244 could have significant implications to our understanding of the internal structure of compact
1245 stars in nuclear astrophysics.

1246 Heavy-ion collisions have shown great potential in studying the light hypernuclei proper-
1247 ties and their production mechanism. There have been unprecedented measurements from
1248 RHIC and LHC on both the lifetime and binding energy (anti-)hypertriton (${}^3_{\Lambda}\text{H}$ and ${}^3_{\Lambda}\bar{\text{H}}$). At
1249 low energy heavy-ion collisions, due to the high baryon density and high strangeness popu-
1250 lation, statistical hadronization thermal model predicts a significant enhancement of various
1251 light hypernuclei production yield, shown in Fig. 37 left panel [132]. The STAR FXT energy

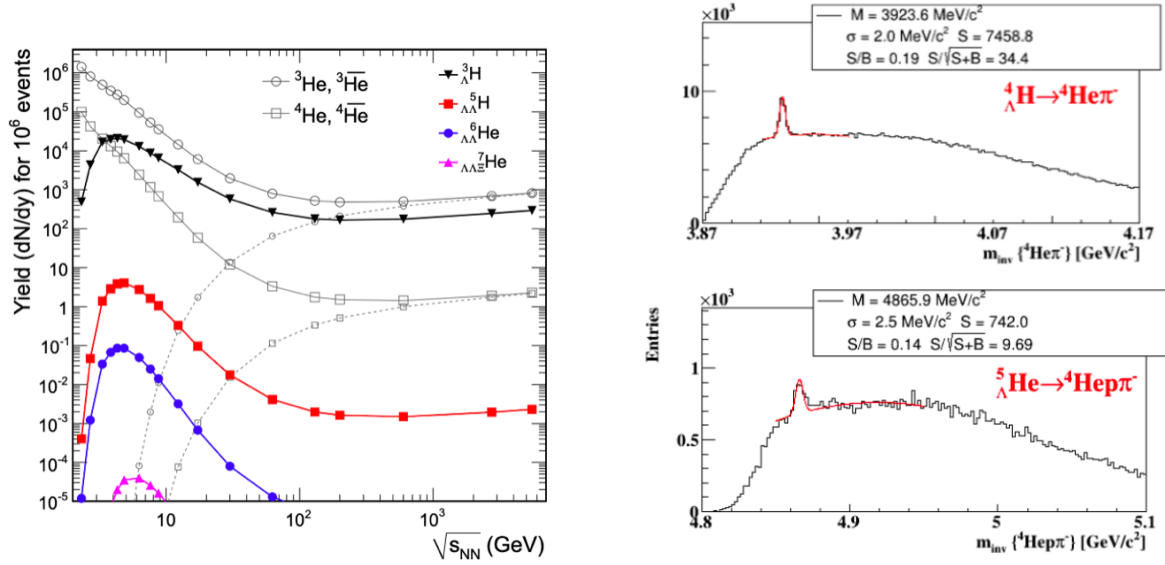


Figure 37: (Left) Thermal model predictions of various light nuclei and hypernuclei production yield at midrapidity in central heavy-ion collisions as a function of collision energy [132]. (Right) Invariant mass distribution of ${}^4\text{He}\pi^-$ (top) ${}^4\text{He}\pi^-$ (bottom) from 2018 FXT data at $\sqrt{s_{NN}}=3.0$ GeV. The ${}^4_{\Lambda}\text{H}$ and ${}^5_{\Lambda}\text{He}$ hypernuclei signal is clearly visible on top of background.

1252 region from $\sqrt{s_{NN}} = 3.0 - 7.7$ GeV sits nicely in the maximum mid-rapidity production yield
 1253 of various hypernuclei while STAR detector layout has the best midrapidity acceptance cov-
 1254 erage at 3.0 GeV. Figure 37 right panel shows the reconstructed ${}^4_{\Lambda}\text{H}$ and ${}^5_{\Lambda}\text{He}$ signal from the
 1255 2018 FXT dataset at $\sqrt{s_{NN}} = 3.0$ GeV. These are so far the most unprecedented statistics
 1256 on these light nuclei that will allow us to systematically investigate their lifetimes, binding
 1257 energies as well as their production yield and collective flow behavior in heavy-ion collisions.

1258 2.1.3 Au+Au Collisions in FXT Mode at $\sqrt{s_{NN}} = 9.2, 11.5, \text{ and } 13.7$ GeV

1259 The BES-II program aims to study the nature of QCD matter by varying the temperature
 1260 and baryon chemical potential. High baryon chemical potentials are achieved by ‘stopping’
 1261 the baryons which made up the two colliding nuclei. To better understand the development
 1262 of the baryon chemical potential and its profile through the interaction region, it is necessary
 1263 to study the rapidity density distribution of the protons across a broad range in rapidity.
 1264 It is important that the rapidity range covered includes the peak of the participants which
 1265 have been accelerated during the collision process. For all collider energies available at RHIC
 1266 (7.7 GeV and above), the peak of the rapidity distribution of the stopped protons is outside
 1267 or at the edge of the acceptance of the STAR TPC (which only extends 0.6 units beyond
 1268 mid-rapidity with particle identification via dE/dx , this is extended to 1.0 units of rapidity
 1269 using eTOF particle ID); for $\sqrt{s_{NN}} = 9.2, 11.5, \text{ and } 13.7$ GeV, the shifted 0.9, 1.0, and 1.1
 1270 units away from mid-rapidity respectively. However, in fixed-target mode the STAR detector
 1271 is excellent for studies of stopping as the acceptance extends 1.7 units from target rapidity

1272 (see figure 38 toward mid-rapidity; for $\sqrt{s_{\text{NN}}} = 9.2, 11.5, \text{ and } 13.7$ GeV, $y_{\text{CM}} = 2.28, 2.50,$
 1273 and 2.68 respectively. Combining collider and fixed-target measurements at each energy
 1274 will provide full coverage from target rapidity to center-of-mass rapidity. The stopping of
 1275 the incident protons is the key to changing the baryon chemical potential in the interaction
 1276 region and the changing baryon chemical potential is the key to mapping out the phase
 1277 diagram of QCD matter.

1278 Complete rapidity density distribution for identified particles will provide important con-
 1279 straints for models. It has been noted by Shen [?] that the high rapidity tails of the dN/dy
 1280 distributions are very important and that high rapidity data are rare. In the energy range
 1281 from $\sqrt{s_{\text{NN}}} = 5.0$ to 200 GeV, the only available proton rapidity density distribution mea-
 1282 surements are from NA49 at 8.77 and 17.3 GeV [133] and from BRAHMS at 62.4 and 200
 1283 GeV [134]. Shen used these data to constrain his 3-D models of the collisions to better un-
 1284 derstand the elliptic flow measurements in heavy-ion collisions. In the BES-II energy range,
 1285 around 10 GeV, these models can set strong constraints on the dependence of Quark-Gluon
 1286 Plasma shear viscosity on temperature and net baryon chemical, however, in order to do so,
 1287 it is necessary to have knowledge of the rapidity distributions of net-protons and produced
 1288 particles.

1289 It has been proposed that the trend of the rapidity shift of the stopped protons with
 1290 collision energy will provide a signature of the softening of the equation of state at the phase
 1291 transition [135]. Specifically, the model which has a two phase equation of state shows that
 1292 the increase in the rapidity shift with collision energy stalls in the $\sqrt{s_{\text{NN}}} = 8$ to 12 GeV
 1293 range.

1294 We proposed to extend the studies to proton stopping through the BES-II energy regime.
 1295 Specifically we propose to add three more energies to the high end of the FXT energy range.
 1296 These energies are chosen to provide three more overlap energies with the collider program.
 1297 Single beam energies of 44.5, 70, and 100 GeV will provide interactions at $\sqrt{s_{\text{NN}}} = 9.2,$
 1298 11.5, and 13.7 GeV (the top energy is not quite an overlap energy with the 14.6 GeV
 1299 collider system). Combining the midrapidity coverage from the collider mode and the target
 1300 rapidity coverage from the fixed-target mode will provide full rapidity coverage for inclusive
 1301 observables. Since the focus for program will be inclusive observables, 50 M events will be
 1302 sufficient at each energy. We propose that at each of these three energies, twelve hours be
 1303 spent on beam development and twelve hours be spent taking data.

1304 2.2 Further Opportunities in Run-21

1305 2.2.1 Au+Au Collisions in FXT Mode at $\sqrt{s_{\text{NN}}} = 3.0$ GeV - II: 2 Billion Goal

1306 In the previous section, we have discussed the great physics interests for low energy heavy-ion
 1307 collisions utilizing the FXT setup at the STAR experiment. We have made our arguments
 1308 of taking a minimum of 300 million Au+Au FXT events at $\sqrt{s_{\text{NN}}} = 3.0$ GeV. With further
 1309 available beam time, we would like to request to collect up to 2 billion events with the same
 1310 setup for the following physics measurements.

1311 Proton correlations higher than 4-th order are useful to study the possible contributions

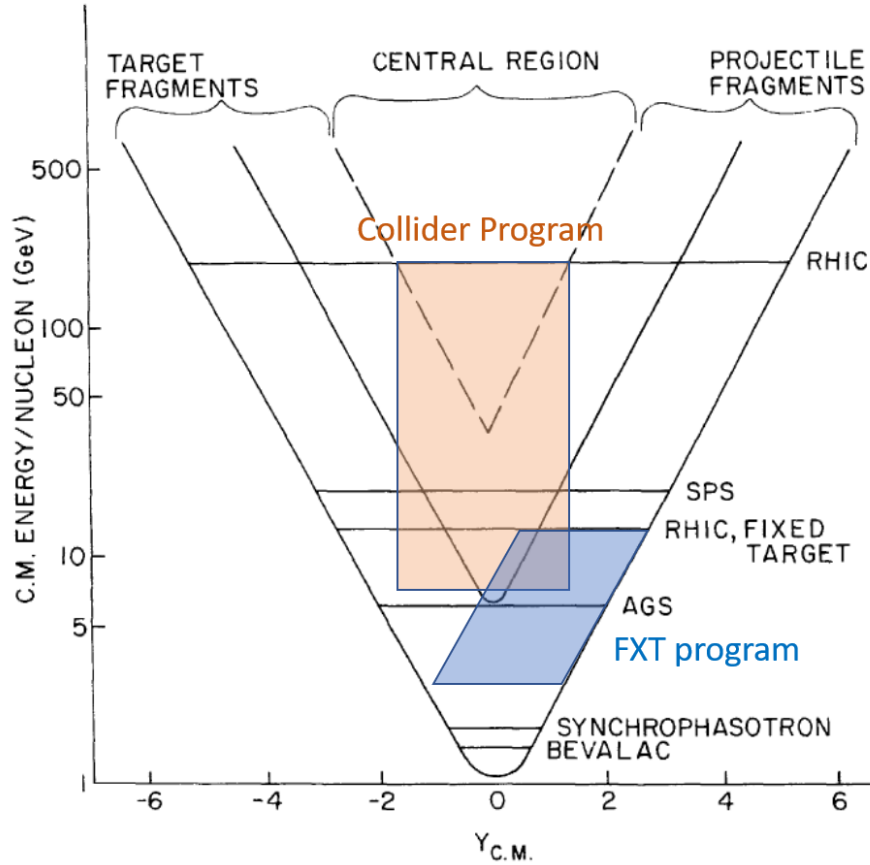


Figure 38: This figure has been modified from a figure in the introduction of the Conceptual Design Report for the RHIC facility. The black lines indicate different regions in the rapidity - center of mass energy space. The ‘V’ shaped region in the top center of the figure which is labeled at the central region have been predicted and demonstrate to be a low baryon chemical potential region characterized by a continuous phase transition between the QG and the hadron gas. The outer ‘V’ shaped region is dominated by the target fragments. Colored regions are overlaid to indicate the coverage of the STAR detector for collider (Orange) and FXT (Blue) modes. For the three higher energies currently being proposed, the FXT acceptance covers the region dominated by target fragments while the collider acceptance covers the equilibrated central region.

1312 of protons from hadronic phase or QGP phase [136]. The requested 2 billion events statistics
 1313 will enable us to perform the analyses of proton moments and cumulants up to 5-th and 6-th
 1314 orders to search for the phase boundary in the high baryon density region. A much larger
 1315 data sample (2 billion events) will enable us to further investigate the centrality/system size
 1316 dependence of ϕ meson production as well as allow us to study the ϕ collective flow behavior
 1317 in these collisions. The 2 billion events would potentially produce $> 25 \frac{5}{\Lambda\Lambda} H$ counts ($\sim 5\sigma$
 1318 discovery potential) within the STAR detector acceptance assuming the predicted yield from
 1319 a thermal model [132].

1320 While there have been tens of hypernuclei measured so far, there are only very few double-

1321 Λ hypernuclei candidates reported from emulsion experiments [137–141]. Their properties
 1322 are directly related to the $\Lambda\Lambda$ interaction. Low energy heavy-ion collisions can be a unique
 1323 environment to copiously produce these light double- Λ hypernuclei. For instance, according
 1324 to the thermal model prediction, the ${}^5_{\Lambda\Lambda}\text{H}$ production yield increases by more than 3 orders
 1325 of magnitude at the low energies compared to that at top RHIC and LHC energies [132].

1326 We performed a Monte Carlo simulation study for the decay chain ${}^5_{\Lambda\Lambda}\text{H} \rightarrow {}^5_{\Lambda}\text{He} + \pi^-$, and
 1327 ${}^5_{\Lambda}\text{He} \rightarrow {}^4\text{He} + p + \pi^-$ within the STAR detector acceptance. Assuming the production yield
 1328 based on the thermal model prediction, with 2 billion Au+Au FXT data at $\sqrt{s_{\text{NN}}} = 3.0$ GeV
 1329 and with the iTPC and eTOF detector, we will have a chance to observe ~ 27 signal counts,
 1330 corresponding to $\sim 5\sigma$ discovery potential assuming negligible background contribution based
 1331 on the two-decay-vertices topology reconstruction.

1332 2.2.2 Au+Au Collisions at $\sqrt{s_{\text{NN}}} = 17.1$ GeV

1333 **Net-proton kurtosis and light nuclei yield ratio from RHIC BES-I** One of the
 1334 main goals of the RHIC Beam Energy Scan (BES) program is the search for the QCD
 1335 critical point (CP), which is a distinct singular feature of the QCD phase diagram. The
 1336 experimental confirmation of the existence of the CP would become a landmark in the
 1337 exploration of the phase structure of hot dense nuclear matter. The characteristic feature of
 1338 the CP is the divergence of the correlation length and density fluctuations. These critical
 1339 phenomena can be probed by measuring event-by-event fluctuations of conserved quantities,
 1340 such as baryon, electric charge, and strangeness numbers. The effect of the CP could show as
 1341 a non-monotonic energy dependence of higher order moments of these conserved quantities
 1342 in close proximity of the critical point during a beam energy scan [121].

1343 In the years 2010-2017 RHIC finished the first phase of the Beam Energy Scan (BES) and
 1344 took data in Au+Au collisions at $\sqrt{s_{\text{NN}}} = 7.7, 11.5, 14.5, 19.6, 27, 39, 54.4, 62.4,$ and 200 GeV.
 1345 With these experimental data STAR measured the higher order fluctuations of net-proton,
 1346 net-charge, and net-kaon multiplicity distributions [122, 143–147]. One striking observation
 1347 is the behavior of the fourth-order cumulants, or kurtosis, of the net-proton fluctuation $\kappa\sigma^2$
 1348 in most central (0-5%) Au+Au collisions as a function of beam energy. As shown on the
 1349 left of Fig. 39, the fourth order net-proton fluctuation is close to unity above 39 GeV but
 1350 deviates significantly below unity at 19.6 and 27 GeV, then approaches or turns above unity
 1351 at lower energies. This behavior may suggest that the created system skims close by the CP,
 1352 and receive positive and/or negative contributions from critical fluctuations. The right of
 1353 Fig. 39 shows the characteristic signature of the critical point for energy dependence of the
 1354 fourth order fluctuations when the system passes through the critical region [121]. Along
 1355 this argument, a peak structure above unity for net-proton kurtosis measurement at lower
 1356 energies could be the signature of the CP. However, it is worth to point out that a first
 1357 order phase transition could also cause a large increase of net-proton kurtosis [136]. When
 1358 entering into the spinodal region (mixed phase), the double peak structure of σ field may
 1359 cause the increase of the fourth order cumulants (C_4).

In addition, STAR has measured light nuclei (deuteron and triton) production in Au+Au collisions at RHIC BES energies. The ratio of these yields are predicted to be sensitive to the

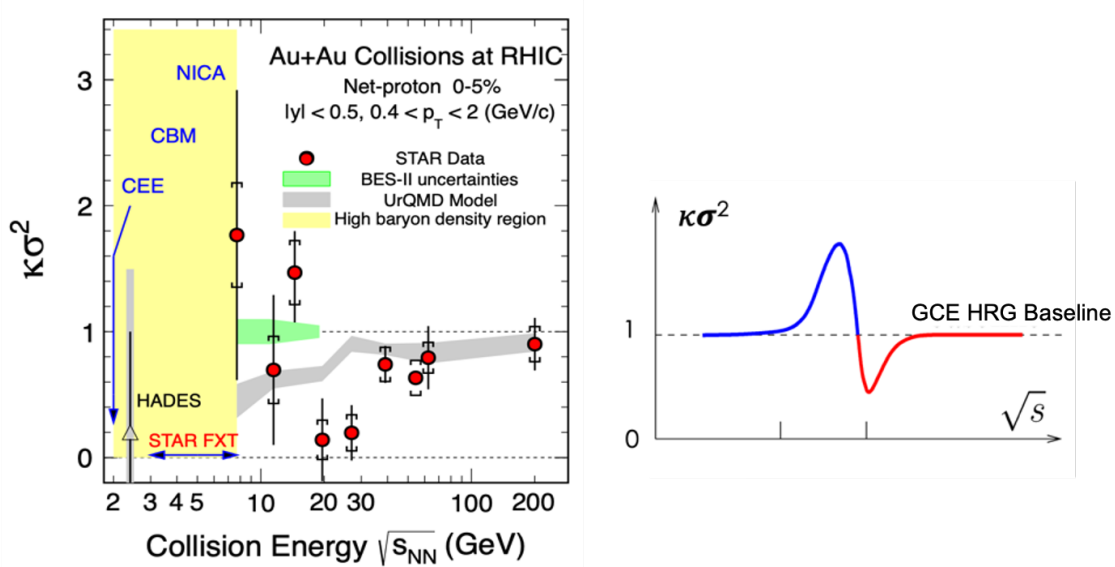


Figure 39: (Left) The fourth order net-proton fluctuations $\kappa\sigma^2$ in most central (0-5%) Au+Au collisions as a function of collision energy from STAR BES-I measurements [122]. (Right) The characteristic signature predicted by the theoretical model for energy dependence of the fourth order fluctuations when the system passes through the critical region [121].

neutron relative density fluctuations at kinetic freeze-out, which is expected to increase near the critical point and/or a first order phase transition [142]. The neutron density fluctuation is defined as $\Delta n = \langle(\delta n)^2\rangle/\langle n\rangle^2$, which can be approximated from:

$$\Delta n = \frac{1}{g} \frac{N_t \times N_p}{N_d^2} - 1,$$

1360 where N_p , N_d and N_t are the proton, deuteron and triton yields, respectively and g is a
 1361 constant factor of 0.29 [142]. In the left panel of Fig. 40, we show the yield ratio $N_t \times N_p/N_d^2$
 1362 in central Au+Au collisions as a function of collision energy. These light nuclei yield ratios
 1363 are obtained by using the feed-down corrected proton yields [148], deuteron yield [149], and
 1364 preliminary triton results [150]. The ratio as a function of energy exhibits a non-monotonic
 1365 energy dependence with a peak around 19.6 GeV. The blue band showing a flat energy
 1366 dependence represents the calculation of the light nuclei yield ratio in Au+Au collisions
 1367 ($b < 3$ fm) from a transport JAM model [151]. Furthermore, the yield ratio shown in Fig. 40
 1368 seems to show a drop between 14.5 and 19.6 GeV. The experimental observation of non-
 1369 monotonic energy dependence in yield ratio may suggest a double peak structure of the
 1370 neutron density fluctuation, indicating that the system goes through the critical region and
 1371 the first order spinodal region, as displayed in Fig. 40 right.

1372 In RHIC 2021 run, we propose to take one more energy point in Au+Au collisions at
 1373 17.1 GeV based on the following two observations presented in Figs. 39 and 40, aiming at
 1374 the QCD critical point search with net-proton kurtosis and light nuclei yield ratio:

- 1375 1. Net-p kurtosis and light nuclei yield ratio, which are both sensitive to the critical

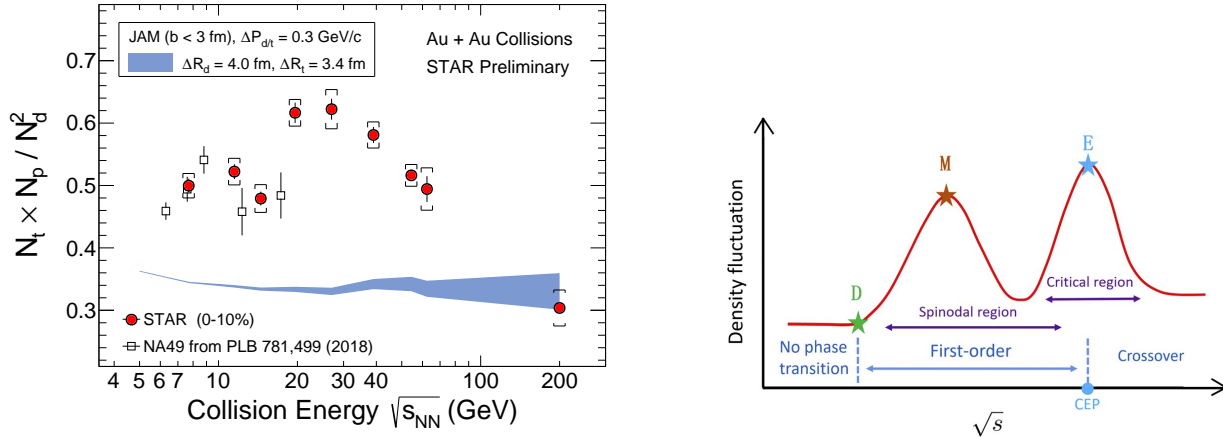


Figure 40: (Left) Collision energy dependence of the light nuclei yield ratio ($N_t \times N_p / N_d^2$) in central Au+Au collisions. The open square data based on NA49 results in central Pb+Pb collisions at $\sqrt{s_{NN}}=6.3$ (0-7%), 7.6 (0-7%), 8.8 (0-7%), 12.3 (0-7%), and 17.3 (0-12%) energies. (Right) Illustration of the density fluctuation as a function of collision energy in the critical region and the spinodal region [142].

1376 fluctuation, show dip and peak structures around 19.6 GeV. These may suggest that
 1377 the system passed through the critical region around 19.6 GeV.

1378 2. We observe sudden changes between 19.6 and 14.5 GeV in the energy dependence of
 1379 net-p kurtosis and light nuclei ratio measurements in the BES-I data measured by the
 1380 STAR experiment. The neutron density fluctuations at low energies below 14.5 GeV
 1381 are consistent with the results from NA49 experiment [142].

1382 These two observations indicate that the critical point may be close to 19.6 GeV. Since
 1383 there are sudden changes in both observables between 19.6 (chemical freeze-out $\mu_B = 205$ MeV)
 1384 and 14.5 GeV ($\mu_B = 266$ MeV), it is important to conduct a finer beam energy scan between
 1385 these two energies. Therefore, we request a run with Au+Au collisions at $\sqrt{s_{NN}}=17.1$ GeV
 1386 ($\mu_B = 235$ MeV), which is just between 19.6 and 14.5 GeV with equal μ_B gap, about 30 MeV,
 1387 on each side.

1388 **Answer to PAC questions from last year** When this proposal was presented last year,
 1389 we have received the following comments: "To make the case for a $\sqrt{s_{NN}} = 17.1$ GeV run, the
 1390 key input will be results from measurements of fluctuation observables from Run-19 data
 1391 taken at $\sqrt{s_{NN}} = 19.6$ and 14.6 GeV. If these measurements, with the smaller error bars
 1392 that are anticipated, show evidence for a possible two-peaked structure in the plot of net
 1393 proton kurtosis or other fluctuation observables as a function of $\sqrt{s_{NN}}$, this could at that
 1394 time become a strong argument for a run at $\sqrt{s_{NN}} = 17.1$ GeV."

1395 To reply the comments from PAC last year, we have analyzed the net-proton fluctuation
 1396 and light nuclei production from the Run-19 fast-offline data of Au+Au collisions at $\sqrt{s_{NN}}$

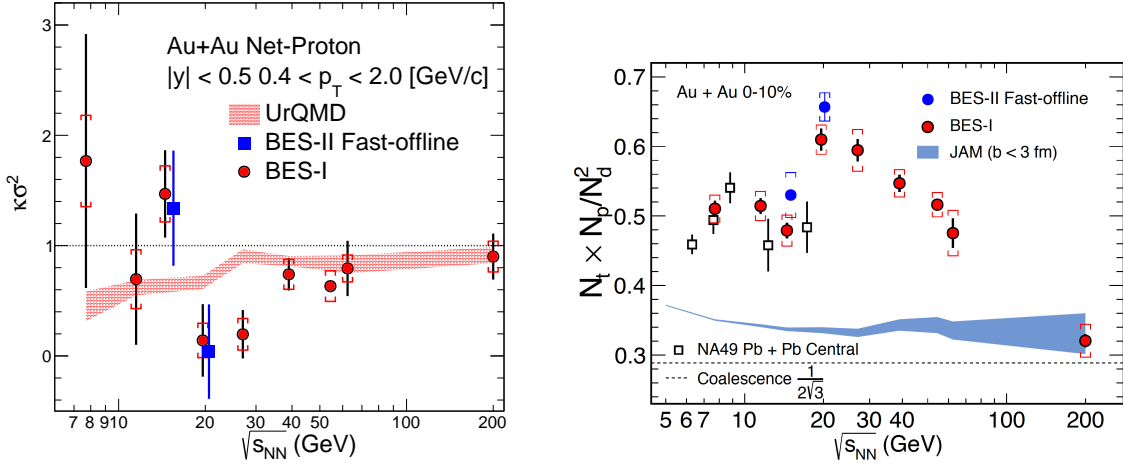


Figure 41: The red solid makers are the results from RHIC BES-I and the blue square represents the results of Run19 fast-offline data. (Left) The fourth order net-proton fluctuations $\kappa\sigma^2$ in most central (0-5%) Au+Au collisions as a function of collision energy. (Right) Collision energy dependence of the light nuclei yield ratio ($N_t \times N_p / N_d^2$) in central Au+Au collisions. The open square data based on NA49 results in central Pb+Pb collisions at $\sqrt{s_{NN}}=6.3$ (0-7%), 7.6 (0-7%), 8.8 (0-7%), 12.3 (0-7%), and 17.3 (0-12%) energies.

1397 = 14.6 and 19.6 GeV. The statistics of the Run-19 fast-offline data of 0-5% Au+Au collisions
 1398 shown in Fig. 41 are about 580k and 750k events for 14.6 and 19.6 GeV, respectively, which
 1399 are roughly about 5% of the full min.-bias statistics of these two data sets. It is found
 1400 that both net-proton fluctuation and light nuclei yield ratios in 0-5% most central Au+Au
 1401 collisions from BES-I are consistent with the results from Run 19 fast-offline data of 14.6 and
 1402 19.6 GeV. For clarity in Fig. 41, the X-axis positions of Run 19 fast-offline data are slightly
 1403 shifted.

Table 8: Event statistics (in millions) needed in a Au+Au run at $\sqrt{s_{NN}} = 17.1$ GeV for fourth order net-proton fluctuations ($\kappa\sigma^2$) and light nuclei yield ratio ($N_t \times N_p / N_d^2$) measurements.

Triggers	Minimum Bias	net-p $\kappa\sigma^2$ (0-5%)	$N_t \times N_p / N_d^2$ (0-10%)
Number of events	250 M	6% error level	3.6% error level

1404 **Required number of minimum bias events and statistical uncertainty level** Ac-
 1405 cording to the previous estimation of the required event statistics for BES-II energies pre-
 1406 sented in Table 6, we need about 250 million minimum-bias events for the net-proton kurtosis
 1407 measurement at 17.1 GeV, which requires 2.5 weeks data taking. It gives us about 12.5 million
 1408 events (250/20) in 0-5% most central collisions. This will ensure that the relative statistical
 1409 error of net-proton $\kappa\sigma^2$ in 0-5% most central Au+Au collisions will reach the 6% level (shown

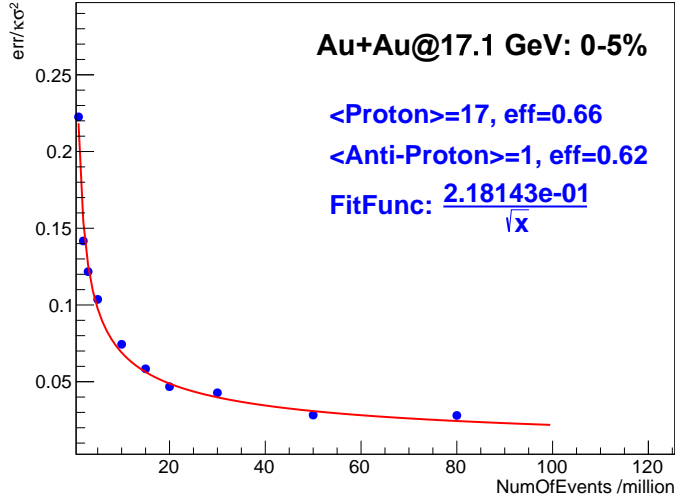


Figure 42: Monte Carlo simulation for the relative statistical errors of net-proton $\kappa\sigma^2$ in 0-5% most central Au+Au collisions at 17.1 GeV. A Skellam distribution for net-proton is assumed; the mean value for protons and anti-protons are 17 and 1, respectively. The average efficiencies for proton and anti-proton are 0.66 and 0.62, respectively.

1410 in Fig. 42). This event statistics will also ensure that the relative statistical error of light nu-
 1411 clei ratio will reach about 3.6% level in 0-10% central Au+Au collisions. In additional to the
 1412 improved statistics, utilizing the iTPC will enable the measurement of lower p_T light nuclei
 1413 and will reduce the systematic uncertainties associated with the low p_T yield extrapolation.

1414 If nature puts the critical point in the QCD phase diagram between 14.5 and 19.6 GeV
 1415 (with μ_B around 200–270 MeV), RHIC has the best chance to discover it !

1416 S

1417 2.2.3 Small System Run: O+O at $\sqrt{s_{NN}} = 200$ GeV

1418 **Introduction** Collective long-range azimuthal correlations in A+A collisions have been
 1419 successfully described as a hydrodynamic response by a fluid-like system to geometric shape
 1420 fluctuations in the initial state. In recent years, observation of similar collective phenomena
 1421 in small-system collisions, such as pp and $p+A$ collisions, has attracted wide interests in
 1422 the community. The interpretation of a fluid-like state formed there has been challenged,
 1423 as the small size and short lifetime might prevent the system from quickly thermalizing
 1424 and evolving hydrodynamically. Instead, collectivity arising either from initial momentum
 1425 correlations motivated by gluon saturation models [152] or via a few scatterings among
 1426 partons (without hydrodynamization) [153–155] has been proposed as alternative source of
 1427 collectivity that may be dominant in small systems. Lots of experimental and theoretical
 1428 efforts have been devoted to the study of collectivity in small-system collisions, with the
 1429 goal of understanding the time-scale for the emergence of collectivity and the mechanism for

1430 early-time hydrodynamization in large collision systems.

1431 One key feature that distinguishes initial momentum correlation models (ISM) from final-
 1432 state interaction models (FSM, including hydrodynamics or a few scatterings) is the connec-
 1433 tion to the initial-state geometry [156]. In FSM, the collectivity is a geometrical response to
 1434 initial shape fluctuations, i.e., v_n is approximately proportional to the n^{th} -order initial-state
 1435 eccentricity ε_n . In ISM, such a geometrical response is expected to be absent [157]. It was
 1436 proposed that a geometry scan of various colliding systems with different spatial eccentricities
 1437 can help distinguish between contributions of these two scenarios [158].

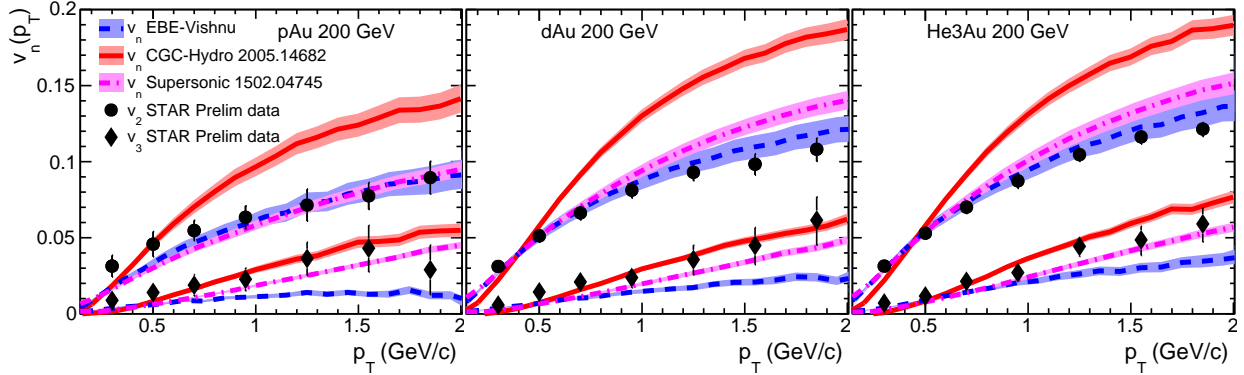


Figure 43: Comparison of v_2 and v_3 in p +Au, d +Au and ${}^3\text{He}$ +Au collisions at $\sqrt{s_{\text{NN}}} = 200$ between STAR data and various model calculations.

1438 Such a small system scan program has been recently carried out at RHIC for a few
 1439 asymmetric small systems including p +Au, d +Au and ${}^3\text{He}$ +Au, where studies of elliptic
 1440 flow (v_2) and triangular flow (v_3) have been performed [43, 159, 160]. In a Glauber model
 1441 that only considers the fluctuations of nucleon positions [158], ε_2 in d +Au and ${}^3\text{He}$ +Au is
 1442 expected to be larger than in p +Au, while ε_3 in p +Au and d +Au are expected to be smaller
 1443 than in ${}^3\text{He}$ +Au. However, once the fluctuations at subnucleonic scales are included [157],
 1444 the ε_3 are expected to be similar among all three systems. Fig. 43 compares the STAR
 1445 v_2 and v_3 results with three hydrodynamic models predictions with different assumptions
 1446 about the initial state. Calculations [161, 162] that include initial momentum anisotropy
 1447 and/or subnucleonic fluctuations indeed describe the STAR v_3 data in all three systems, but
 1448 one of the model [161] overestimates the v_2 data. On the other hand, hydrodynamic model
 1449 based on fluctuations only at nucleonic level [163] fails to describe the v_3 data. This implies
 1450 that the initial state in these asymmetric small collision systems are not well constrained, in
 1451 particular in p +Au and d +Au system (there is reasonable consensus that the flow results in
 1452 ${}^3\text{He}$ +Au is dominated by FSM). The relative importance of FSM vs. ISM for the v_n data in
 1453 small systems is an area of intense ongoing debate [164].

1454 **Physics case for a small A+A scan** So far, both RHIC and the LHC carried out col-
 1455 lisions for either relatively large (Pb+Pb, Au+Au, Xe+Xe, Cu+Cu, ...), which are well
 1456 described by hydrodynamic models, or small asymmetric systems (p +Pb, p +Au, d +Au, and

1457 $^3\text{He}+\text{Au}$), whose initial state are poorly constrained as discussed above. To quantitatively
 1458 understand the initial momentum anisotropy and the role of subnucleonic fluctuations, col-
 1459 lisions of small but symmetric systems, such as O+O, Al+Al and Ar+Ar will be necessary.
 1460 They will also fill the gap between pp and Cu+Cu systems is a crucial unexplored frontier¹,
 1461 where a transition from ISM to FSM dominated collectivity may be observable. The list of
 1462 key open questions related to collectivity in small systems includes:

- 1463 • How much do initial-state correlations vs. geometry-driven final-state interactions con-
 1464 tribute to the observed collectivity? Can we unambiguously establish experimental
 1465 evidence of initial-state correlations?
- 1466 • For final-state scenarios, to what extent does the collectivity arise from a hydrodynamic
 1467 fluid-like QGP, as opposed to an off-equilibrium system with only a few scatterings per
 1468 parton?
- 1469 • What is the role of subnucleonic fluctuations in determining the initial-state geometry?
- 1470 • Can we observe jet quenching in small systems?

1471 A new comprehensive scan of colliding ion species at RHIC by systematically varying the
 1472 system size and geometry between pp and Cu+Cu collisions, will provide a unique lever-
 1473 arm to dial contributions from various mechanisms and impose strong constraints on both
 1474 ISM and FSM. Since the last RHIC $p/d/\text{He}+\text{Au}$ scan, the STAR experiment has completed
 1475 several detector upgrades that extend p_T and particle identification to $|\eta| < 1.5$, and provide
 1476 centrality and event plane determination in $2 < |\eta| < 5$ [17, 165, 166]. An ongoing forward
 1477 upgrade to instrument the $2.5 < \eta < 4$ region with tracking detectors and calorimeters
 1478 will be completed prior to 2021 run [167]. The extended detector capability will allow a
 1479 full exploration of collectivity using all the observables and methods developed for large
 1480 systems at RHIC/LHC. We will have better control of the non-flow systematics, leading to
 1481 a better understanding of the multi-particle nature of the collectivity and the longitudinal
 1482 correlations to constrain the full 3D initial conditions. As an illustration, model studies of
 1483 v_2 and v_3 in a series of small systems including symmetric (C+C, O+O, Al+Al, Ar+Ar) and
 1484 asymmetric ($p+\text{Au}$, $d+\text{Au}$, $^4\text{He}+\text{Au}$) collisions using the AMPT model are shown in Fig. 44.
 1485 AMPT belongs to the category of final-state interaction models, where v_n is largely driven by
 1486 the geometry of initial nucleon distributions. The v_2 values from asymmetric systems follow
 1487 different trends: the v_2 in $d/{}^4\text{He}+\text{Au}$ increases with N_{ch} , while it is relatively constant in
 1488 $p+\text{Au}$. The v_3 values show a similar N_{ch} dependence as symmetric systems, except for $d+\text{Au}$
 1489 which deviates from the common trend at large N_{ch} . This study demonstrates that, in a
 1490 scenario driven by final-state interactions, a clear difference is expected between $d/{}^4\text{He}+\text{Au}$
 1491 and A+A for v_2 , while a relatively similar behavior should be observed for v_3 . Contributions
 1492 from other sources, especially ISM, are expected to follow a drastically different behavior;
 1493 as the system size increases, the ISM contribution will gradually become subdominant.

¹RHIC has no limitation on small A+A systems, based on private communication with Wolfram Fischer

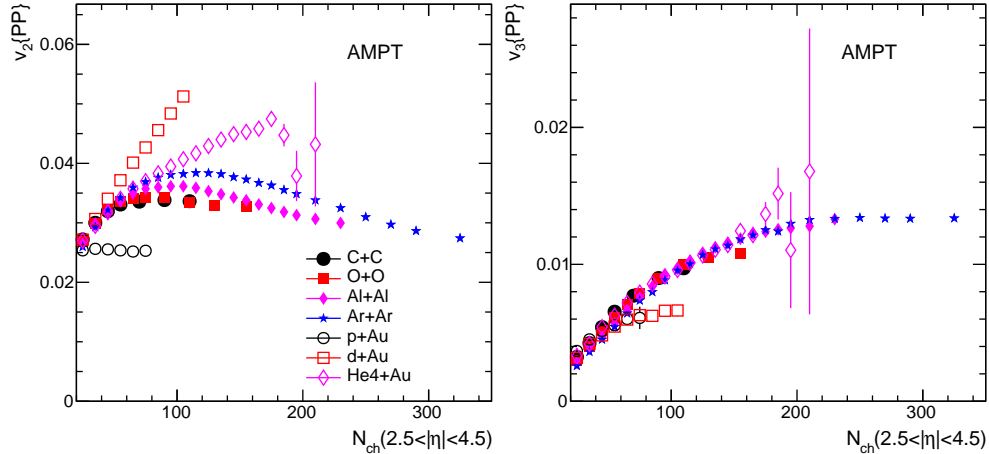
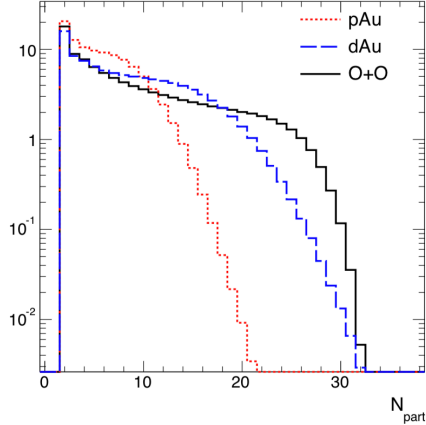


Figure 44: (Left) AMPT predictions for v_2 and (Right) v_3 as a function of N_{ch} in four symmetric and three asymmetric small collision systems.

1494 **Arguments for a short O+O run in 2021** In this BUR, we propose a O+O run at $\sqrt{s_{\text{NN}}} = 200$ GeV towards the end of the BES-II in 2021, to be followed up with a comprehensive scan
 1495 of symmetric and asymmetric small collision systems using the STAR forward upgrade after
 1496 2021, possibly in collaboration with sPHENIX. The choice of O+O collisions as the starting
 1497 point is motivated by the following reasons: 1) O+O has an N_{part} coverage comparable to
 1498 p +Au and d +Au but with a much flatter distribution (see Fig 45), which allows much better
 1499 control of initial geometry and centrality bias, 2) the Oxygen is a reasonably sized system
 1500 for which the both the nucleonic and subnucleonic DOF are important, which together
 1501 with p/dAu data can be used disentangle these contributions, 3) a strong synergy with the
 1502 proposed higher-energy O+O run at the LHC around around 2023–2024 to enable a direct
 1503 comparison of the same small-system collision species at drastically different energies. More
 1504 details, including hydrodynamic model predictions, are presented and discussed below.

1506 The recent yellow report on the future LHC heavy-ion physics program discusses the
 1507 possibility for smaller A+A collisions [168]. This includes a proposal of an O+O run at
 1508 $\sqrt{s_{\text{NN}}} = 2.76\text{--}7$ TeV in 2022², and other light-ion species such as Ar+Ar beyond 2028. As
 1509 mentioned earlier, one big advantage of the O+O system is that it allows a better control
 1510 of N_{part} and ε_n , compared to peripheral Pb+Pb collisions [168]. An O+O run at RHIC
 1511 right after the BES-II would provide a timely comparison of the same small system at very
 1512 different collision energies (0.2 TeV vs. 2.76–7 TeV). This “RHIC-LHC energy scan” provides
 1513 a unique opportunity to study systems with nearly identical initial nucleon geometry but
 1514 very different subnucleon fluctuations and different saturation scales. The large lever-arm
 1515 in collision energy should provide new insights on the onset behavior of collectivity, jet
 1516 quenching, or any other final-state effects in small systems: any model has to describe
 1517 results at both energies, which naturally leads to a better understanding of results at each
 1518 energy.

²According the latest schedule of the LHC run 3, O+O run will most likely be scheduled in 2023.



	pAu	dAu	$^{16}\text{O}+^{16}\text{O}$
$\langle N_{\text{part}} \rangle$	5.8	8.8	9.5

Figure 45: The N_{part} distribution in O+O collisions compared with p +Au and d +Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV estimated from Glauber model. The table to the right shows the average N_{part} values in the three systems.

1519 Figure. 46 compares the $v_n(p_T)$ data and hydrodynamic calculations for $n = 2$ and 3 at
 1520 two energies in a large A+A system (left) and in a p +A system (right). It is well-known that
 1521 $v_n(p_T)$ for charged hadrons in large systems has very little $\sqrt{s_{\text{NN}}}$ dependence from RHIC
 1522 to LHC [169], as well as from 39 to 200 GeV at RHIC [170, 171]. This is confirmed by the
 1523 left panel which compares Pb+Pb [172] with Au+Au [173] data at 30–40% centrality, as
 1524 well as calculations from the CGC-Hydro model. However, a comparison of $v_n(p_T)$ between
 1525 p +Pb [174] and p +Au [160] central data suggests a small difference in v_2 , while the v_3 data
 1526 are nearly identical. In the FSM picture, this suggests that the contributions of subnucleonic
 1527 fluctuations to the initial eccentricities are very different between the two collision energies.
 1528 In the ISM picture, it may be the result of an energy dependence of initial momentum
 1529 anisotropy. It would be exciting to see whether the $\sqrt{s_{\text{NN}}}$ dependence for v_2 and v_3 in p +A
 1530 collisions also persists in small A+A systems such as O+O collisions between RHIC and
 1531 LHC. The CGC-Hydro model calculations of v_2 and v_3 in O+O collisions at RHIC and
 1532 the LHC energies are shown in Fig. 46 (middle), where a split in both v_2 and v_3 between
 1533 two energies is predicted. These rather non-trivial \sqrt{s} dependence across different collision
 1534 systems reflects the rich physics mechanisms behind origin of collectivity.

1535 We propose a one-week O+O program in 2021 right after BES-II. Assuming a total
 1536 interaction rate of ~ 10 – 15 kHz (based on recent isobar runs), the STAR DAQ rate of 2 kHz
 1537 and the RHIC uptime of 50% (12 hour/day), tentative numbers of events we expect to
 1538 record for different triggers are summarized in Table 9 for one week, default run plan, and
 1539 two weeks as a more optimal running scenario. Note that we do not have an estimation of
 1540 minimum-bias trigger efficiency at this point, and assumed it to be $\sim 100\%$.

1541 The event statistics listed in Table 9 should allow precision measurements of many types
 1542 of two-particle correlations, including the N_{ch} dependence of integral v_n , p_T dependence of v_n
 1543 in 0-5% for identified particles (π , K, p and ϕ) to test the NCQ-scaling. The non-flow effects
 1544 for these observables can be studied in detail thanks to the large acceptance of iTPC and
 1545 EPD. Based on a Glauber model estimation, the $\langle N_{\text{part}} \rangle$ value is 9.5 and 26 for minimum-bias

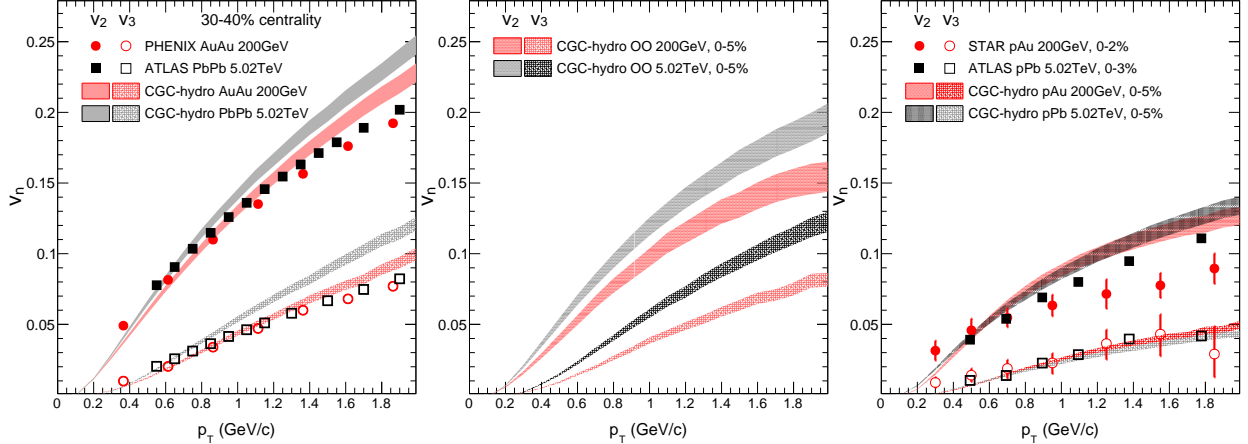


Figure 46: Comparison of measured v_2 and v_3 between Pb+Pb and Au+Au 30–40% centrality events (Left) and high-multiplicity p +Pb and p +Au data (Right) at RHIC and the LHC energies. The CGC-Hydro model calculations are also shown for Au+Au and Pb+Pb (Left), p +Au and p +Pb (Right), and O+O as a prediction (Middle) at both energies.

Table 9: Number of events (in millions) needed in an O+O run at $\sqrt{s_{NN}} = 200$ GeV for various triggers for one week (default) and two weeks (optimistic) running scenarios.

Triggers	Minimum bias	0–5% centrality
Events (1 week)	400 M	200 M
Events (2 week)	800 M	400 M

1546 and 0-5% central O+O collisions, respectively.

1547 Figure 47 shows the projection of the statistical precision for the ϕ meson $v_2(p_T)$ in 0–5%
 1548 centrality O+O collisions. Under the assumption that its v_2 in O+O is similar to that of
 1549 a charged hadron in p +Au around $p_T \sim 2$ –3 GeV/ c , the estimation scales the ϕ $v_2(p_T)$ in
 1550 peripheral Au+Au collisions [175] to approximately match the charged hadron v_2 in p +Au
 1551 collisions in Fig. 46, accounting for differences in $\langle N_{part} \rangle$, event plane resolution, and event
 1552 statistics. A decent measurement of ϕ meson v_2 can be achieved with one week of running.

1553 In fact, the statistics requirement in Table 9 is mainly driven by multi-particle correla-
 1554 tions, for example four-particle cumulants for single harmonics $c_2\{4\} = \langle v_n^4 \rangle - 2 \langle v_2^2 \rangle^2$, four-
 1555 particle symmetric cumulants $SC(2, 3) = \langle v_2^2 v_3^2 \rangle - \langle v_2^2 \rangle \langle v_3^2 \rangle$ and three-particle asymmetric
 1556 cumulants $AC(2, 4) = \langle v_2^2 v_4 \cos 4(\Phi_2 - \Phi_4) \rangle$ (Φ_n is the event plane). These observables are
 1557 sensitive to event-by-event fluctuations of collectivity, and measurements of them at LHC in
 1558 pp , p +Pb and Pb+Pb collisions have led to high impact results which provide evidence for
 1559 geometry response in small systems [176–179].

1560 Figure 48 shows the projection of the statistical precision for the $c_2\{4\}$ measurement.
 1561 The projected precision should allow a measurement of $c_2\{4\}$ signal, assuming a $v_2\{4\}$ value

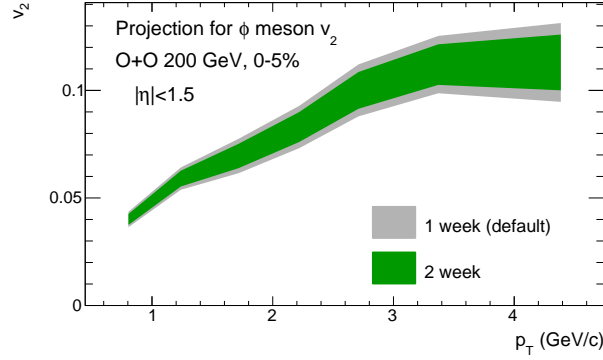


Figure 47: Projected statistical error on ϕ meson $v_2(p_T)$ in central O+O collisions within the TPC acceptance.

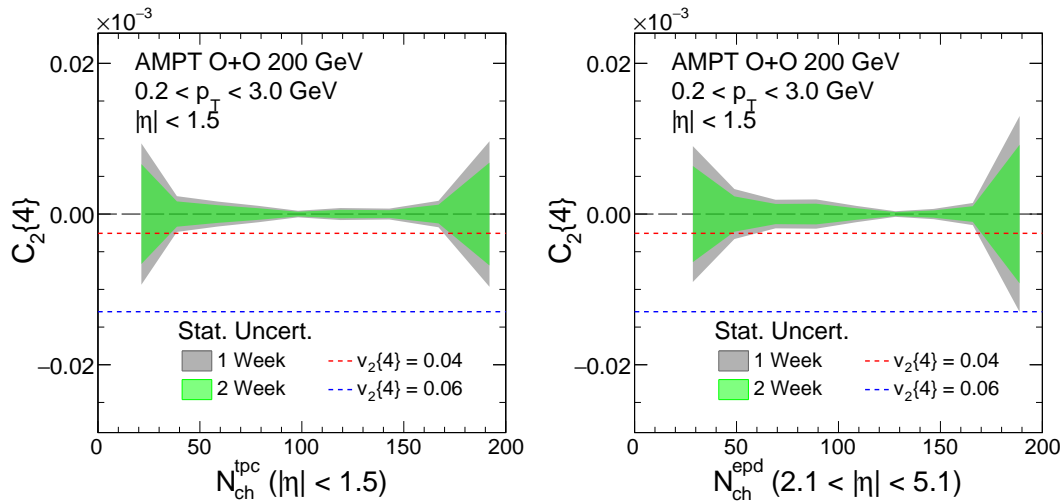


Figure 48: (Left) The projected statistical error bar on $c_2\{4\}$ in 0.2-3 GeV/c in the TPC acceptance as a function of number of charged particles in TPC acceptance and (Right) EPD acceptance.

1562 to be between 4–6%³.

1563 Figure 49 shows the projection of the statistical precision for the charged hadron R_{AA}
 1564 measurement for minimum bias O+O collisions (assume 400 Million). This calculation
 1565 includes the state-of-art knowledge of nPDF effects and jet quenching modeling of Refs.
 1566 [181, 182]. A significant suppression of $R_{AA} = 0.85 - 0.9$ is expected which should be mea-
 1567 surable with decent statistical uncertainty out to 15 GeV.

1568 **Answer to PAC questions from last year** When this proposal was presented last year,
 1569 we have received the following comments: "With regards to an O+O run, the case for this
 1570 could become persuasive if, between now and next year, theorists with expertise in hydro-
 1571 dynamics can provide some simulations that demonstrate what hydrodynamics predicts for
 1572 v2 and v3 behavior in O+O collisions, and how this compares to results from p+A, Cu+Cu,

³The p_T integrated $v_2\{4\}$ in $d+Au$ from PHENIX [180] at forward rapidity is about 4%

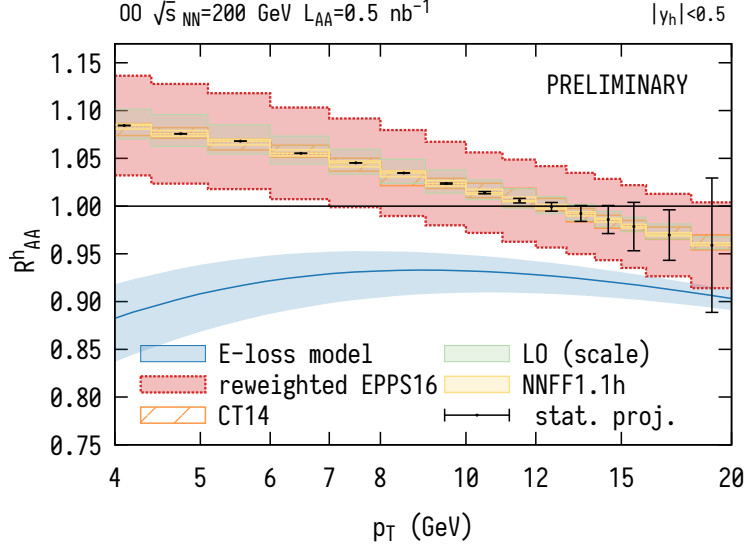


Figure 49: Prediction of minimum bias hadron nuclear modification factor for $\sqrt{s_{NN}} = 200$ GeV O+O collisions following Refs. [181, 182]. A particular parton energy loss model predictions (blue line) is overlaid with the baseline in the absence of parton rescattering. The blue band represents model uncertainty only due to experimental uncertainties in $\sqrt{s_{NN}} = 5.02$ TeV Pb+Pb collisions used to fit a free model parameter. The red band shows nPDF uncertainties reweighted with additional CMS pPb dijet data. Proton PDF (orange), leading order scale (green) and fragmentation function (yellow) uncertainties are fully correlated and cancel. Error bars illustrate statistical uncertainties for OO mock data at 100% efficiency.

1573 and Au+Au collisions. We also suggest that these calculations should be undertaken for
 1574 $\alpha + \alpha$, Be+Be, Al+Al and Ar+Ar collisions also, as well as for O+Au and other asymmetric
 1575 small+large nuclear collision options, so as to be able to make the case that O+O is the
 1576 optimal physics choice, most likely to yield new or substantially improved understanding of
 1577 questions relating to how small droplets of QGP equilibrate and what is the smallest droplet
 1578 of QGP that is possible to be formed in collisions at 200 GeV".

1579 We have prepared the following answers to these comments:

- 1580 • Why O+O? 1) O+O collisions cover similar N_{part} range as $p+\text{Au}/d+\text{Au}$ (see Fig-
 1581 ure 45) where the collectivity debate is ongoing, 2) O+O has similar N_{part} but differ-
 1582 ent nucleon/sub-nucleon fluctuations, 3) leverage similar measurement at LHC for new
 1583 insight and precision.
- 1584 • Are there theoretical calculations? Many model studies on O+O exist by now, which re-
 1585 flects the community interests: 1812.08096,1904.10415,1908.06212, 1910.09489, 2003.06747,
 1586 2005.14682. Figure 46 shows the new prediction on O+O taken from 2005.14682,
- 1587 • Why not other collision systems? Analyzing power for 2k-particle cumulants $v_n\{2k\}$
 1588 scales as $N_{\text{events}} \times N_{\text{part}}^{2k}$, system smaller than O+O, such as C+C require much longer
 1589 running time and also difficult to setup the high-multiplicity triggers due to steeply

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falling N_{ch} distributions. Also these systems will not have compatible N_{part} coverage as p/d+Au systems. One may suggest to repeat the p/d+Au. But this will require long running time, since previous p+Au (d+Au) data was taken over 5 (1.5) weeks period in run15 (run16).

- Why not larger small systems? Larger asymmetric system we already have Cu+Au and $^3\text{He}+\text{Au}$. Results from both are consistent with final-state interpretation from both experiments and theory. Any system in between such as O+Au with $\langle N_{part} \rangle=60$ is expected to be dominated by final state effects.

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2.3 Future Possibilities

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2.3.1 Exploring the Nuclear Equation-of-State (EoS) with Heavy Ion Collisions

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In the interior of the fireball created in HI collisions, nuclear densities of up to $10 \rho_0$ can be achieved depending on the energy of the colliding nuclei [183]. Similar densities are predicted to be present in the core of neutron stars (NS). However, the composition and maximal mass of NS is highly dependent on the nuclear equation-of-state (EoS) which is close related to the compressability of nuclear matter. Therefore HI collisions are considered as a ideal tool to study the EoS at high nuclear densities and establish a bridge between astrophysics and nuclear physics.

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Already in the early 80s several observables probing the EoS (i.e. which are sensitive on the density and pressure of the system) like particle production [184, 185], transverse momentum analysis [186], directed and elliptic flow [187, 188] were proposed. At low energies and densities up to about $2.5 \rho_0$ elliptic flow analysis favour a soft EoS [189] while at AGS energies (between 2.5 and $4.5 \rho_0$) no clear picture could be established yet. Transverse flow measurements hint to a soft ($K \leq 210$) EoS while elliptic flow measurements indicate a stiff ($K \approx 300$) EoS [188]. Another method to probe the equation-of-state of nuclear matter was proposed already 1985 by Aichelin and Ko [190] by measuring the sub-threshold production of strange particles which are created in the dense medium early in the reaction. Especially the multi-strange particle production at sub-threshold energies should be highly sensitive to the density of the medium since multiple-collision processes are required. So the information on the EoS is stored in the yield and slope of the excitation function of the sub-threshold produced particles which can be compared to various theoretical models. It is also clear that the abundance of produced particles is highest at heavy systems like $Au + Au$ since the reaction volume is big. Therefore comparison measurements with light system like $C + C$ can be used as a reference where the influence of the EoS is small. In addition, systematic errors both in experiment and theory cancel out to a large extent by taking the yield ratio of produced (multi-) strange particles at both systems.

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Figure 50 shows the normalized yield ratio $((M_{K^+}/A)_{Au+Au}/(M_{K^+}/A)_{C+C})$ of sub-threshold produced K^+ as a function of 4 different beam energies [191]. A comparison to different models [192, 193] predict a soft EoS at low bombarding energies around $E_{beam} \approx 1.5 \text{ GeV}$. On the other hand the finding of neutron stars with masses above 2 solar masses can only be ex-

1629 plained by a stiff EoS [194] which should be measurable in flow observable at higher incident
 1630 energies.

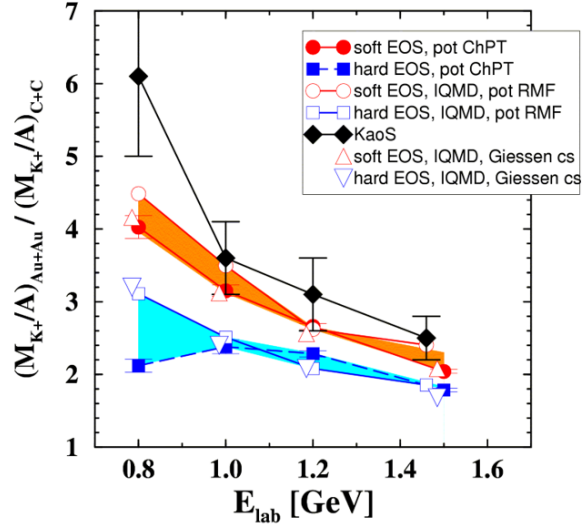


Figure 50: Excitation function of the ratio of K^+ multiplicity obtained in $Au + Au$ over $C + C$ reactions [191] together with model calculations [192,193] indicating a soft EoS at this bombarding energies.

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An independent observable at these higher energies are the production rates of multi-strange baryons, of which several have their NN production threshold in the reach of BES II: Ξ^- (production threshold $\sqrt{s_{NN}}=3.247$ GeV), Ω^- (production threshold $\sqrt{s_{NN}}=4.092$ GeV), Ξ^+ (over $\Xi^+\Xi^-$ -channel, production threshold $\sqrt{s_{NN}}=4.520$ GeV) and Ω^+ (over $\Omega^+\Omega^-$ -channel, production threshold $\sqrt{s_{NN}}=5.222$ GeV). A beam energy scan in small steps (with 3 to 4 points with distance of about 200 to 300 MeV) below the production threshold energy for the various particles species for $Au + Au$ collisions could give access to the properties of the EoS as function of colliding energy and density. In fact a big fraction of the measurements were already finished during the BESII campaign and only a few additional fixed target points have to be performed which depends on the number of investigated particles used as probes. The measurement of the pressure as function of energy in small steps could even allow for the discovery of a 1st order phase transition in the covered energy range [195]. The measurements at the heavy system could be accompanied by a reference measurement of a light system (for example $C + C$) for at least one particle species and extrapolated to the others. However, since the production yield scales with the number of participants the time needed for the light system is significantly longer (few weeks) as for the heavy system which needs typically btw. 12 h and a day.

Nevertheless the proposed measurements have a high discovery potential and could be an opportunity for a future fixed target HI program at STAR.

2.4 Exploring the Microstructure of the QGP (Run-23 and Run-25 Au+Au)

The completion of the RHIC's scientific mission involves the two central goals [196] of (i) mapping out the phase diagram of the QCD, and (ii) probing the inner workings of the QGP by resolving its properties at short length scales. The complementarity of the RHIC and LHC facilities to study the latter is scientifically as essential as having more than one experiment independently study the microstructure of the QGP. With several years of operating the recently installed iTPC upgrade and the soon-to-be installation and operation of STAR's forward detectors, the STAR collaboration will be in an excellent position to take advantage of its vastly improved detection capabilities. Combine this with the prospect of a substantial increase in beam luminosities and RHIC will be uniquely positioned to fully engage in a detailed exploration of the QGP's microstructure. Through careful discussions in its physics working groups, the STAR collaboration has identified a number of topics that together make a compelling case to take data during Runs 23-25 alongside sPHENIX, and successfully complete RHIC's scientific mission. In this section, we present a selection of those topics that will take full advantage of both STAR and RHIC's unique capabilities and address the following important questions about the inner workings of the QGP.

- What is the precise temperature dependence of the shear η/s , and bulk ζ/s viscosity?
- What is the nature of the 3-dimensional initial state at RHIC energies? How does a twist of the event shape break longitudinal boost invariance and decorrelate the direction of an event plane?
- How is global vorticity transferred to the spin angular momentum of particles on such short time scales? And, how can the global polarization of hyperons be reconciled with the spin alignment of vector mesons?
- What is the precise nature of the transition near $\mu_B = 0$, and where does the sign-change of the susceptibility ratio χ_6^B/χ_2^B take place?
- What is the electrical conductivity, and what are the chiral properties of the medium?
- What can we learn about confinement and thermalization in a QGP from charmonium measurements?
- What are the underlying mechanisms of jet quenching at RHIC energies?

The event statistics projections that are used in this section will rely on the CAD's recently update 2023E and 2025E Au+Au luminosities [197] and are listed in Table 10. For each year we presume 24 weeks of RHIC operations, and based on past run operations an overall average of 85% \times 60% (STAR \times RHIC) uptime, respectively. The minimum-bias rates assume a conservative 1.5 kHz DAQ rates which will allow sufficient bandwidth for specialized triggers which are listed as integral luminosities.

year	minimum bias [$\times 10^9$ events]	int. luminosity [nb^{-1}]		
		all vz	$ vz < 70\text{cm}$	$ vz < 30\text{cm}$
2014 2016	2	26.5	19.1	15.7
2023	10	43	38	32
2025	10	58	52	43

Table 10: STAR minimum bias event statistics and luminosity projections for the 2023 and 2025 Au+Au runs. For comparison the 2014/2016 event statistics and luminosities are listed as well.

1686 At RHIC it is possible to build detectors that can span from mid-rapidity to beam rapidity
1687 – with the two recent upgrades STAR is able to achieve this unique capability. STAR’s BES-
1688 II upgrade sub-systems comprised of the inner Time Projection Chamber (iTPC, $1.0 < |\eta| <$
1689 1.5), endcap Time Of Flight (eTOF, $1 < \eta < 1.5$) and Event Plane Detector (EPDs, $2.1 <$
1690 $|\eta| < 5.1$), that are all commissioned and fully operational since the beginning of 2019 [16–18].
1691 As will be discussed in Sect. 4, the STAR collaboration is constructing a forward rapidity (2.5
1692 $< \eta < 4$) upgrade that will include charged particle tracking and electromagnetic/hadronic
1693 calorimetry [198]. For charge particle tracking the aim is to construct a combination of
1694 silicon detectors and small strip thin gap chamber detectors. The combination of these two
1695 tracking detectors will be referred to as the forward tracking system (FTS). The FTS will
1696 be capable of discriminating the hadron charge sign. It should be able to measure transverse
1697 momentum of charged particles in the range of $0.2 < p_T < 2$ GeV/ c with 20–30% momentum
1698 resolution. In what follows, we will refer to the combination of the existing TPC ($|\eta| < 1$)
1699 and the iTPC upgrade as iTPC ($|\eta| < 1.5$) for simplicity.

1700 The impetus for running STAR during the year of 2023-2025 in terms of bulk correlation
1701 measurements in Au+Au 200 GeV collisions comes from the gain in : i) extended accep-
1702 tance and ii) enhanced statistics. In the first subsections, we briefly describe how these two
1703 opportunities can be exploited to perform correlations measurements that are unique to the
1704 physics goals of the RHIC heavy ion program.

1705 Next, thanks to a reduced material budget between the beam and the iTPC, STAR will be
1706 uniquely positioned to perform dielectron measurements which we propose to probe degrees
1707 of freedom of the medium and its transport properties. For that we will use high precision
1708 dilepton excess yield, i.e. l^+l^- invariant mass distribution after subtraction of dilepton
1709 sources produced after freeze-out, and contributions from the initial collisions such as Drell-
1710 Yan and correlated charm-anticharm pairs. Furthermore, we propose to study the virtuality,
1711 Wigner function and final-state magnetic field in QGP. For the latter the photon-photon
1712 collisions in ultra-peripheral, peripheral, and midcentral reactions and p+A (all centralities)
1713 in both channels e^+e^- , $\mu^+\mu^-$ will be measured with high accuracy.

1714 In the last subsections, we address our proposed charmonium measurements and motivate
1715 the importance of STAR’s proposed program of precise jet measurements to explore the
1716 micro-structure of the QGP.

1717 **2.4.1 Correlation Measurements Utilizing Extended Acceptance**

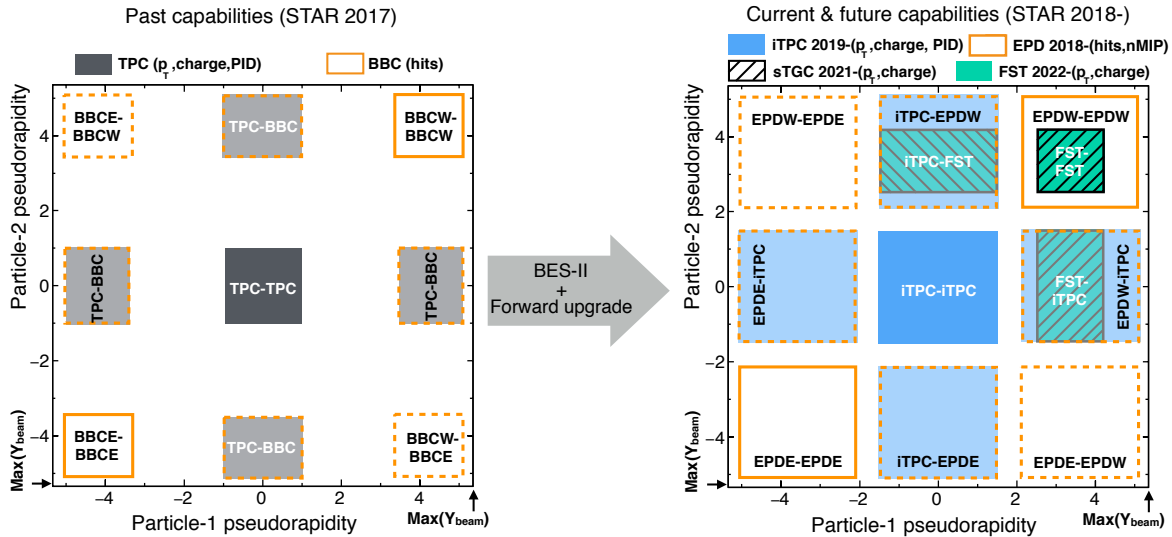


Figure 51: A visual representation of two-particle phase space in pseudorapidity covered by STAR detectors with respect to the region allowed by maximum beam rapidity ($Y_{beam}=5.36$ at 200 GeV Au+Au collisions) of RHIC. Left and right panels show the capabilities before and after BES-II and forward upgrade of the STAR detector, respectively.. Note that in addition to a larger pair acceptance, the EPD granularity is over an order of magnitude larger than that of the BBC, and individual EPD tiles are shown to be separable into 1, 2, 3 MIP responses.

1718 Figure 51 demonstrates how STAR with the BES-II and forward upgrades will extend the
 1719 two-particle phase-space (in terms of η_1 and η_2 with respect to beam rapidity) many times
 1720 enabling us to perform correlation measurements over a wide window of relative pseudorapid-
 1721 ity. Since many of the important correlation measures are based on two-particle correlations,
 1722 this enhanced phase-space will provide STAR with many advantages: 1) increase the number
 1723 of pairs to bring better precision, 2) reduction in different sources of the non-flow background
 1724 by increasing pseudorapidity separation. Many multi-particle correlations will also get bene-
 1725 fitted due to increase in triplets, quadruplets and so on due to overall increase in acceptance.
 1726 With this unique extended pseudorapidity reach offered by the BES-II and forward upgrade
 1727 of the STAR detector, our goal is to perform correlation measurements aimed towards a
 1728 deeper understanding of the largely unexplored three-dimensional structure of the initial
 1729 state and temperature dependent transport properties of the subsequent fluid-like medium
 1730 produced in heavy ion and small system collisions at RHIC.

1731 Two key sets of measurements are of interests: 1) pseudorapidity dependence of azimuthal
 1732 correlations, 2) pseudorapidity dependence of global hyperon polarization.

1733 **Pseudorapidity-dependent azimuthal correlation to tightly constrain tempera-**
 1734 **ture dependence of viscosity** The idea of tightly constraining the temperature viscosity

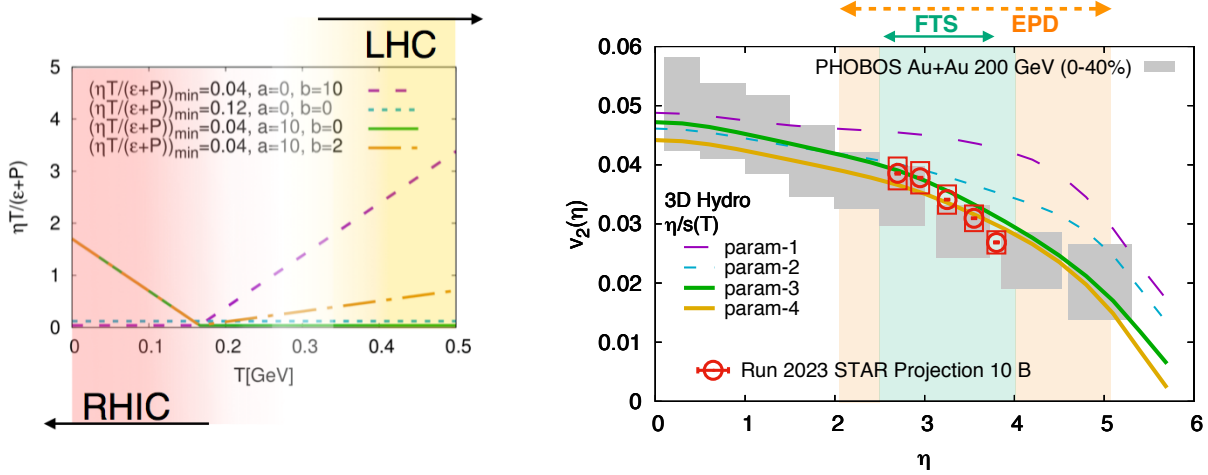


Figure 52: (Left) Different parameterizations of temperature dependence of shear viscosity to entropy η/s (T) (at zero chemical potential) used in the hydrodynamical simulation of Ref [199]. Interestingly, it has been demonstrated in Ref [200] that the region of lowest η/s is the one that can be probed at RHIC. (Right) Effects on the elliptic flow co-efficient v_2 due to the different parameterizations of the viscosity parameter indicating better constraints on η/s (T) can only be performed by measurements at forward rapidities at RHIC. The interpretation of the existing PHOBOS data is limited by the large uncertainties. The projection for STAR measurements are shown on the same plot.

1735 of the QGP was envisioned in the 2015 Long Range Plan for Nuclear Science [196]. The
 1736 QCD matter formed at RHIC shows nearly perfect fluidity characterized by the smallest
 1737 viscosity to entropy ratio η/s known in nature. The temperature dependence of η/s and
 1738 other transport parameters has not been fully constrained. One major aim is to perform
 1739 precision measurements to contain the temperature dependence of shear η/s (T) and bulk
 1740 ζ/s (T) viscosity. Hydrodynamic simulations have demonstrated that since the temperature
 1741 of the produced fireball in HICs vary with the rapidity, the measurement of the rapidity
 1742 dependence of flow harmonics has the potential to constrain η/s (T) and ζ/s (T) [199]. For
 1743 this, RHIC measurements have advantage over LHC since smaller beam rapidity at RHIC
 1744 provides stronger variations of the temperature with rapidity. The beam energy scan at
 1745 RHIC provides an additional handle on temperature to map η/s (T), and ζ/s (T) over a
 1746 wide range of temperature. Indeed, the hydrodynamic simulation of Ref [199] indicates that
 1747 η/s (T) at lower temperatures, near its possible minimum ($T = T_c$), can be better con-
 1748 strained by RHIC measurements. Results from such simulations are shown in Fig. 52. In
 1749 this simulation, a number of QCD-motivated parameterizations of the temperature depen-
 1750 dence of the shear viscosity was assumed, as shown in Fig. 52 (left). Existing data from the
 1751 PHOBOS collaboration suffer from large uncertainties, therefore only limited constrain on
 1752 the temperature dependence of the transport parameters can be achieved. The BES-II and
 1753 the forward upgrade of STAR will provide precise estimations of different azimuthal correla-
 1754 tion observables: $v_n(\eta)$ and other higher-order ($n > 2$) flow coefficients $v_n(\eta)$, its fluctuations

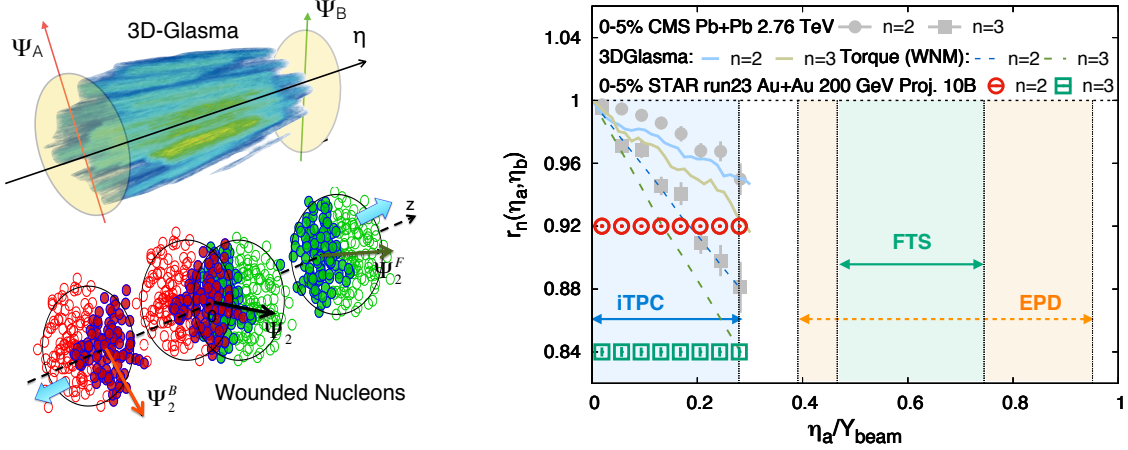


Figure 53: (Left) Cartoon to demonstrate the de-correlation of event planes in the longitudinal direction of collision from gluon saturation based 3D-Glasma model [201] and wounded nucleon model (WNM) [202, 203]. (Right) The longitudinal de-correlation of elliptic anisotropy plane as a function of the pseudorapidity in units of beam rapidity. CMS results are compared predictions from two models in the left with STAR projection for Run 2023 (using preliminary Run 19 results) for anticipated 10 B min-bias events. The colored regions show that the current and future capabilities at STAR (with iTPC+EPD+FTS) can extend such measurements with good precision by covering a large fraction of the beam rapidity at 200 GeV – this demonstrates the unique strength to STAR to study the physics of 3D initial state.

1755 $\sigma(v_n)/v_n$ that have never been measured at forward rapidity and are essential in terms of
 1756 constraining η/s (T) near its possible minimum. While transverse momentum integrated
 1757 quantities can already constrain the shear viscosity, additional information of transverse mo-
 1758 mentum is essential to constrain the bulk viscosity ζ/s . With the forward tracking systems
 1759 it will be possible to measure the p_T dependence of v_n – in particular the information of $\langle p_T \rangle$
 1760 is essential to constrain the bulk viscosity ζ/s (T). This can be done with a possible $A + A$
 1761 collisions with the forward upgrade and running of STAR during the year 2023.

1762 **Pseudorapidity-dependent azimuthal correlation to constrain the longitudinal**
 1763 **structure of the initial state** Initial-state longitudinal fluctuations and fluid dynam-
 1764 ical response of the medium formed in heavy ion collisions can lead to de-correlations of
 1765 the direction of the reaction planes Ψ_n (which determines the orientation of the harmonic
 1766 anisotropies) with pseudorapidity (see Fig. 53). Such effects are often referred to as torque
 1767 or twist of the event shape [203–205] that eventually leads to a breaking of longitudi-
 1768 nal/boost/rapidity invariance. The magnitude of the de-correlation is determined by the
 1769 details of the dynamics of initial state, the distribution of nucleons and partons inside the
 1770 colliding nuclei. Several promising observables have been proposed to study this effect,
 1771 Fig. 53 shows one which can be expressed as $r_n(\eta_a, \eta_b) = V_{n\Delta}(-\eta_a, \eta_b)/V_{n\Delta}(\eta_a, \eta_b)$, where
 1772 $V_{n\Delta}(\eta_a, \eta_b)$ is the Fourier coefficient calculated with pairs of particles taken from three differ-
 1773 ent pseudorapidity regions $-\eta_a$, η_a and η_b . The observable $r_n(\eta_a, \eta_b)$ was originally introduced

1774 and measured by CMS collaboration in Ref [206] and also been measured by the ATLAS
 1775 collaboration in [207]. An observable using three-particle correlations that is sensitive to
 1776 this effect is [208] the relative pseudorapidity dependence of the three-particle correlator
 1777 $C_{m,n,m+n}(\eta_a, \eta_b, \eta_c) = \langle \cos(m\phi_1(\eta_a) + n\phi_2(\eta_b) - (m+n)\phi_3(\eta_c)) \rangle$. Also, another one very sim-
 1778 ilar to r_n in term of design but involves four-particle correlations is: $R_{n,n|n,n}(\eta_a, \eta_b)$ is also
 1779 very useful to study this effect [209]. As shown in Fig. 53 CMS measurements of r_n show a
 1780 strongest de-correlation ($\sim 16\%$ for $n=3$, $\sim 8\%$ for $n=2$) in central events within the range of
 1781 their acceptance. Initial state, described by gluon saturation, as simulated by the 3D-Glasma
 1782 model, the breaking of boost invariance is determined by the QCD equations which predict
 1783 the evolution of gluons in saturation regime with Bjorken-x. At the LHC such models predict
 1784 weaker de-correlation as compared to initial state described by wounded nucleon models and
 1785 does a good job in explaining the r_2 data from CMS collaboration [201] but over-predicts r_3
 1786 results. One expect the nature of the initial state to change from LHC to RHIC, in particular
 1787 the region of Bjorken-x probed is very different at RHIC. It is therefore extremely important
 1788 to utilize the enhanced acceptance of the STAR detector with the Au+Au 200 GeV run to
 1789 study this effect. In Fig.53 STAR projection using preliminary Run 19 results for 10 B events
 1790 is shown for the measurement of r_n within the acceptance $|\eta| < 1.5$. The colored regions
 1791 show that the current and future capabilities at STAR (with iTPC+EPD+FTS) can extend
 1792 such measurements using observables $r_n, C_{m,n,m+n}, R_{n,n|n,n}$ with good precision by covering
 1793 either equal (iTPC only) or larger (iTPC+FTS+EPDs) fraction of the beam rapidity at 200
 1794 GeV compared to the LHC measurements. This unique measurement capability will help
 1795 pin down the nature of the 3-dimensional initial state of heavy ion collisions. It will also
 1796 help constrain different models of QCD that predict the rapidity (or Bjorken-x) dependence
 1797 of valance quark and gluon distribution inside colliding nuclei that has been demonstrated
 1798 by theoretical calculations in Ref. [201,210].

1799 **Pseudorapidity dependence of global hyperon polarization** The global polarization
 1800 of hyperons produced in Au+Au collisions has been observed by the STAR collaboration [20].
 1801 The origin of such a phenomenon has hitherto been not fully understood. Several outstanding
 1802 questions remain. How exactly the global vorticity is dynamically transferred to the fluid like
 1803 medium on the rapid time scales of collisions? Then, how does the local thermal vorticity
 1804 of the fluid gets transferred to the spin angular momentum of the produced particles during
 1805 the process of hadronization and decays? In order to address some of these questions one
 1806 may consider measurement of the polarization of different particles that are produced in
 1807 different spatial parts of the system, or at different times. A concrete proposal is to: 1)
 1808 measure the $\Lambda(\bar{\Lambda})$ polarization with pseudorapidity and 2) measure it for different particles
 1809 such as Ω and Ξ . Both are limited by the current acceptance and statistics available to the
 1810 STAR collaboration. However, as shown in Fig.54 with the addition of iTPC, FTS and with
 1811 high statistics data from run 2023 it will be possible to preform such measurements with
 1812 a reasonable significance. iTPC (+TPC) has excellent PID capability to measure all these
 1813 hyperons. Although FTS has no PID capability we can do combinatorial reconstruction of
 1814 $\Lambda(\bar{\Lambda})$ candidates via displaced vertices. A similar analysis was performed and published by

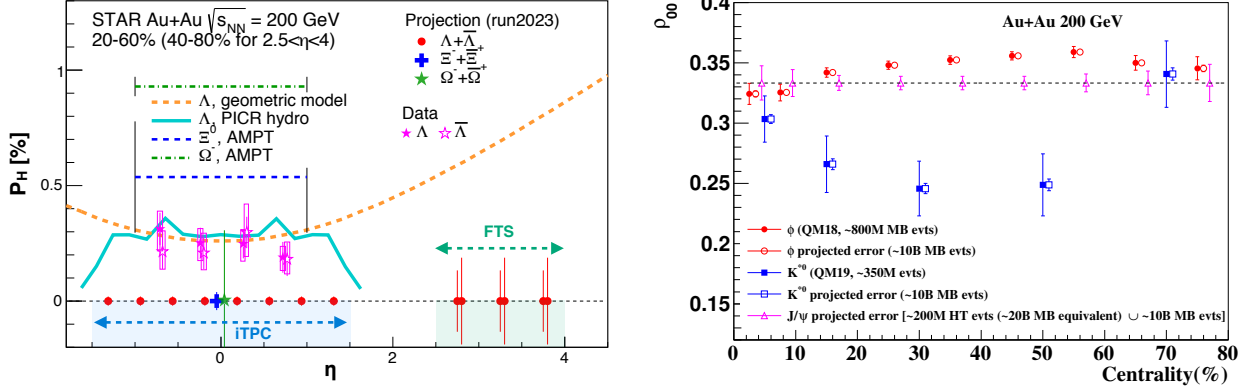


Figure 54: (Left) Projections (along with preliminary data) for differential measurement of $\Lambda(\bar{\Lambda})$ polarization over the extend range of pseudorapidity with the iTPC and FTS detectors of STAR that will help resolve tension between different theoretical model predictions (shown by curves) of polarization with η . In addition, projections for the measurements of spin-1/2 Ξ and spin-3/2 Ω particles is also shown. (Right) Spin alignment co-efficient ρ_{00} as a function of centrality, with projected errors based on ~ 10 billion events. The enhanced statistics run 2023 combined with excellent dilepton capability of STAR will enable us to measure J/ψ alignment along with increasing the significance of ϕ and K^{*0} measurements.

1815 the STAR collaboration using the FTFC detector of STAR in Ref [211]. In order to make
 1816 a conservative projection we assume similar momentum resolution of 10 – 20% for single
 1817 charged tracks, similar overall tracking efficiency, charge state identification capability for
 1818 FTS and FTFC (see forward upgrade section for exact numbers). We also assume the FTS
 1819 with it’s novel-tracking framework will be able to measure a minimum separation of 20 cm
 1820 between the all pairs of one positive and one negative track (a possible decay vertex) from the
 1821 main vertex of the event. This will give rise to about 5% efficiency of $\Lambda(\bar{\Lambda})$ reconstruction
 1822 with about 15 – 20% background contribution from $K_S^0 \rightarrow \pi^+ + \pi^-$ [211] for which we
 1823 can make projections for polarization measurement in Au+Au 200 GeV 40 – 80% assuming
 1824 10 Billion minimum-bias events as shown in Fig. 54. Currently theoretical models predict
 1825 contradictory trends for the pseudorapidity dependence Λ polarization. If the initial local
 1826 orbital angular momentum driven by collision geometry [212] play dominant role it will lead
 1827 to increases of polarization with pseudorapidity. On the other hand if the local thermal
 1828 vorticity and hydrodynamic evolution [213] play a dominant role it will predict decreasing
 1829 trend or weak dependence with pseudorapidity. Such tensions can be easily resolved with
 1830 the future proposed measurement during the run 2023.

1831 2.4.2 Correlation Measurements Utilizing the Enhanced Statistics

1832 Over the past years the STAR collaboration has pursued dedicated measurements at Au+Au
 1833 collisions at $\sqrt{s_{NN}}= 200$ GeV that have major discovery potential but are intrinsically statis-
 1834 tics hungry. In the past, attempts have been made to combine datasets from several years to
 1835 increase the significance of such measurements. This results in additional uncorrectable sys-

1836 tematic uncertainties in the measurements, mostly due to run-to-run variation of detector
 1837 response and collision conditions. A single stable long run with similar detector condi-
 1838 tions, as anticipated during Run-23 will not only reduce the statistical uncertainty but will
 1839 also bring the systematics under control. In the following section and also in section 1.2.5
 1840 we propose correlation measurements that will utilize the enhanced statistics from Run-
 1841 23 and that can lead to high-impact results. To start with we can assume STAR DAQ
 1842 to collect data at the rate of 1.5 kHz and a combined RHIC×STAR uptime of 50% (12
 1843 hour/day) for 24 weeks of running during Run-23. This will lead to the accumulation of
 1844 about $24 \times 7 \times 86400 \times 0.5 \times 1500 \approx 10$ billion events.

1845 **Global spin alignment of J/ψ :** Surprisingly large signals of global spin alignment of
 1846 vector mesons such as $\phi(1020)$ and $K^{*0}(892)$ have been measured via the angular distribu-
 1847 tion of one of their decay products. By far the experimental observation of vector meson
 1848 spin alignment have not been interpreted satisfactorily by theory calculations. It has been
 1849 realized that the mechanism driving the global polarization of hyperons can have its im-
 1850 print on vector mesons spin alignment albeit the observed strength of signals for the two
 1851 measurements cannot be reconciled. In fact the large quantitative difference between the
 1852 measurements of $\phi(1020)$ and $K^{*0}(892)$ spin alignment as shown in Fig. 54 (right) cannot
 1853 be simultaneously explained by conventional mechanisms of spin-orbit coupling, driven by
 1854 angular momentum, without invoking the strong force fields. It is argued that the strong
 1855 force field makes a dominant contribution to the spin-alignment coefficient ρ_{00} of ϕ , while
 1856 for K^{*0} , the contribution is diminished due to the mixing of quark flavors (averaging-out
 1857 of different meson fields) [214, 215]. Therefore, the current preliminary experimental data
 1858 from STAR (Fig. 54, showing $\rho_{00}(\phi) > \rho_{00}(K^{*0})$) support the role of strong force field as
 1859 a key mechanism that leads to global spin alignment. However, a stringent test of such a
 1860 prediction can be performed by measuring the spin alignment of J/ψ . This is because the
 1861 similar argument applies for both ϕ and J/ψ , i.e. like s and \bar{s} , the strong field component
 1862 also couples to c and \bar{c} quarks leading to larger ρ_{00} for J/ψ . In Fig. 54(right) we present the
 1863 projected errors for ρ_{00} of J/ψ estimated for various centralities assuming 200 million events
 1864 (24 weeks running) anticipated in Run-23 by implementing High Tower (HT) triggers with
 1865 the Barrel Electromagnetic Calorimeter. It is worth to mention that apart from J/ψ spin
 1866 alignment, such a large statistics data set will also allow addition differential study of global
 1867 spin alignment of ϕ and K^{*0} and help to further elucidate the mechanism behind vector
 1868 meson spin alignment.

1869 **Sixth order cumulant of net-proton distributions** LQCD calculations [216, 217] pre-
 1870 dict a sign change of the susceptibility ratio χ_6^B/χ_2^B with temperature (T at $\mu_B = 0$) taking
 1871 place in the range of 145-165 MeV. The observation of this ratio going from positive to
 1872 negative values is considered to be a signature of crossover transition. As described in the
 1873 previous section, the cumulants of net-proton distribution are sensitive to the chiral crossover
 1874 transition at vanishing baryon chemical potential. Interestingly, as reported in the last BUR
 1875 and in the recent Quark Matter 2019, the preliminary results from STAR [218] observed

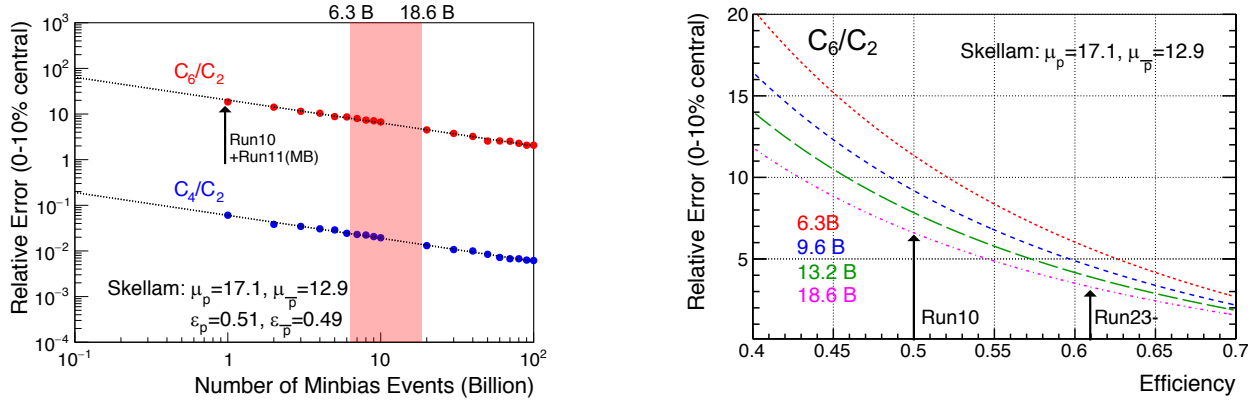


Figure 55: Projection for measurement of ratio of sixth order over second order cumulants of net-proton distribution.

1876 $C_6/C_2 > 0$ in 54.4 GeV while $C_6/C_2 < 0$ in 200 GeV in central Au+Au collisions. The
 1877 observation of positive C_6/C_2 at lower energies can be further confirmed by higher statistics
 1878 data sets from the BES-II program over the energy range of 7.7-19.6 GeV, which also include
 1879 the increased acceptance iTPC. The observation of negative C_6/C_2 is intriguing and by far
 1880 only seen at 200 GeV and based on the current STAR data has less than 2.5σ significance
 1881 in terms of statistical uncertainties. The currently systematic uncertainty is of similar order
 1882 as statistical uncertainty mainly due to combining data sets from Runs 10 and 11. As
 1883 shown in the projection plot of Fig. 55 it is possible to establish definitive observation of
 1884 negative C_6/C_2 at 200 GeV with nearly 10 billion minimum-bias events collected during the
 1885 Run-23 with 15% increase in the reconstruction efficiency and enhanced acceptance of the
 1886 iTPC detector upgrade. A similar measurement can be performed at the LHC however only
 1887 STAR measurements can pinpoint the region of T and μ_B where this phenomenon occurs. In
 1888 other words it can establish if that the sign change occurs somewhere between 54.4 GeV and
 1889 200 GeV. Such measurement has the potential to establish the first experimental observation
 1890 of QCD chiral crossover transition.

1891 2.4.3 Electromagnetic Probes

1892 **Probing the degrees of freedom of the medium and its transport properties.** As
 1893 discussed in Sect. 1.1.3, at $\mu_B \sim 0$ Lattice QCD works and can be directly tested against
 1894 experimental results. This will help to disentangle the ρ melting from other explanations
 1895 such as collision broadening. In case the measured in-medium spectral function merges into
 1896 QGP description this would indicate a transition from hadrons into a structure-less quark-
 1897 antiquark continuum and thus providing the manifestation of chiral symmetry restoration.
 1898 To study this, we will continue to search for a direct signature for chiral symmetry restoration
 1899 via chiral ρ - a_1 mixing. The signal is predicted to be detectable in the dilepton intermediate
 1900 mass range. Difficulties are related to the fact that correlated charm-anticharm and QGP
 1901 saturate invariant mass region of 1.1 — 1.3 GeV/ c^2 . Therefore an accurate measurement

1902 of the excess dilepton yield, i.e. dilepton yield after subtraction of the cocktail of contribu-
 1903 tions from final-state decays, Drell-Yan and those from correlated heavy-flavor decays, up
 1904 to invariant mass of $2.5 \text{ GeV}/c^2$ is required. The challenging analysis on charmed-decayed
 1905 dielectron is ongoing from the data sets taken with the Heavy Flavor Tracker at STAR [?].
 1906 Thus deeper understanding of origin of thermal radiation in Au+Au collisions at $\sqrt{s_{NN}}=$
 1907 200 GeV from \sim zero mass up to $2.5 \text{ GeV}/c^2$ will become possible with rigorous theoretical
 1908 efforts and improved dielectron measurements. Figure 56 shows the expected statistical and
 1909 systematic uncertainties of dielectron excess mass spectrum with all the detector upgrades
 1910 and for the anticipated total Run-23/Run-25 statistics of 20×10^9 events.

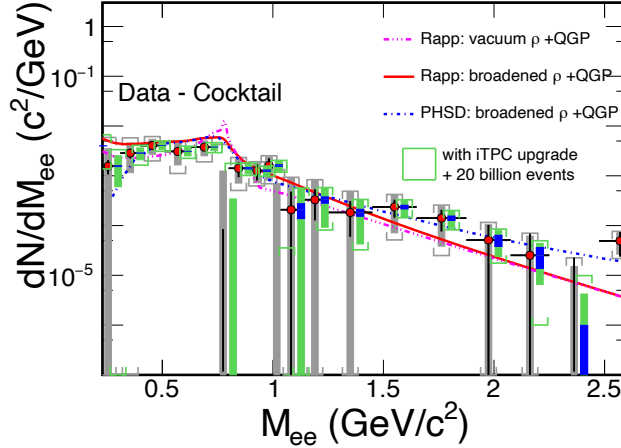


Figure 56: The expected statistical and systematic uncertainties of dielectron excess mass spectrum with the iTPC upgrade compared to the current TPC case. The data are from our measurements in $\sqrt{s_{NN}}= 200 \text{ GeV}$ Au+Au collisions [58]. Model comparisons are also shown. The boxes represent systematic uncertainties from data and the brackets represent the total systematic uncertainties including those from cocktails. The grey ones are for the current case while the green ones are for the expected case. The blue bands represent statistical uncertainties from 20 billion minimum-bias events with the iTPC upgrade.

1911 Another application of dileptons is to use them to measure transport coefficients. The
 1912 electrical conductivity can be directly obtained as the low-energy limit of the EM spectral
 1913 function. We aim to extract such information by studying excess dielectron yield at the
 1914 low-energy regime of dilepton spectra and the conductivity peak at small invariant masses,
 1915 i.e. at low invariant mass and low p_T^{ee} . Low field run could be profitable, however already
 1916 now dileptons with pair p_T^{ee} down to $60 \text{ MeV}/c$ could be measured. Measurement of Drell-
 1917 Yan in p+A collisions at low p_T would provide an important reference to constrain dilepton
 1918 cocktail.

1919 To gain deeper understanding of the microscopic origin of the excess radiation, we will

- 1920 • separate early from later time radiation by measuring dilepton elliptic flow (v_2) as a
- 1921 function of dilepton mass;
- 1922 • identify the source of dilepton radiation by studying dilepton polarization versus in-
- 1923 variant mass (helicity angle);

- 1924 • measure precisely the lifetime of the interacting fireball. As an observable we will use
1925 integrated low-mass yield but also compare explicit model calculations with various
1926 $\tau_{fireball}$;
- 1927 • extract an average radiating source temperature from the fit of Boltzmann distribution
1928 to the invariant mass slope in the range 1.1 - 2.5 GeV/ c^2 spectrum. The higher the
1929 invariant mass, the stronger the QGP contribution to the spectrum, the higher the
1930 chance to measure temperature of QGP.

1931 Last, but not least, concerning direct-photon emission, the existing difference, on the
1932 order of a factor of two, between the low momentum spectra from PHENIX and STAR in
1933 200 GeV Au+Au collisions, has to be resolved. In order to clarify the direct photon puzzle
1934 we will measure direct virtual photon yield as well as its elliptic flow coefficient. We will
1935 particularly focus on low p_T η measurement which might be instrumental in clarifying this
1936 long standing question.

1937 **Studying virtuality, the Wigner function and final-state magnetic fields in the**
1938 **QGP.** The unsuccessful description of STAR data by the STARLight model led to the
1939 attribution of the broadening to the possible residual magnetic field trapped in an electri-
1940 cally conducting QGP, which is the key information to study the chiral magnetic effect.
1941 Similarly, the ATLAS collaboration qualified the effect via the acoplanarity of lepton pairs
1942 in contrast to the measurements in UPC and explained the additional broadening by the
1943 multiple electromagnetic scatterings in the hot and dense medium, which is analogous to the
1944 medium P_{\perp} -broadening effects for jet quenching. These descriptions of the broadening effect
1945 in hadronic collisions are based on the assumption that there is no impact parameter depen-
1946 dence of the transverse momentum distribution for the electromagnetic production. Recent
1947 lowest-order QED calculation, in which the impact parameter dependence is recovered, could
1948 reasonably describe the broadening observed by STAR and ATLAS without any in-medium
1949 effect. To solve the puzzle, we propose to precisely study the initial P_{\perp} -broadening for the
1950 dilepton pair in ultra-peripheral collisions. Different neutron emission tag serve as the cen-
1951 trality definition, and will allow us to explore the broadening baseline variation with impact
1952 parameter. Furthermore, the differential spectrum as a function of pair P_{\perp} , rapidity, and
1953 mass enable us to study the virtuality and Wigner function of the initial electromagnetic
1954 field, which provide the information to extract the momentum and space correlation of EM
1955 field. As shown in Fig. 57, comparing with the latest QED calculation, there still exists ad-
1956 ditional broadening in peripheral collisions, although the significance is about 1σ , which still
1957 leave room for the medium effect. In Runs 23/25, as projected in the figure, we could judge
1958 the existence of additional broadening with much higher precision and further constrain the
1959 strength of final-state magnetic field in QGP.

1960 2.4.4 Deconfinement and Thermalization With Charmonia Measurements

1961 Measurements of charmonia in heavy-ion collisions provide important information about
1962 the thermodynamic properties of the created medium. Production of J/ψ mesons in Au+Au

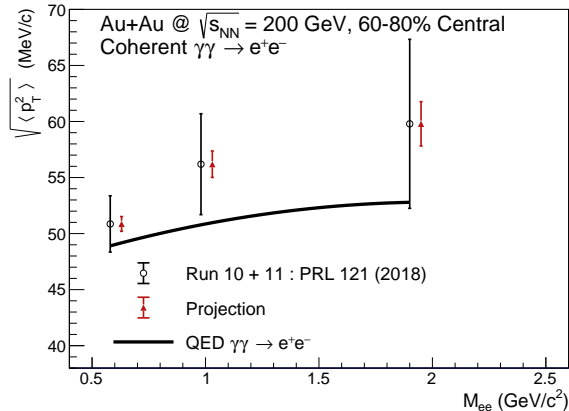


Figure 57: (Color online) The $\sqrt{\langle p_T^2 \rangle}$ of di-electron within the fiducial acceptance as a function of average pair mass, $\langle M_{ee} \rangle$, for 60-80% centrality class in Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV.

1963 collisions at RHIC was found to be suppressed compared to the production in proton+proton
 1964 collisions. The suppression of charmonium states is due to a screening of the $c\bar{c}$ potential by
 1965 the medium color charges. In addition, the J/ψ production can be affected by recombination
 1966 of charm quarks in a later stage of the collision evolution. The regeneration mechanism is
 1967 expected to contribute mostly at the low J/ψ transverse momentum range.

1968 In particular, STAR proposes to utilize the Run-23/25 RHIC heavy-ion runs to measure:
 1969 (i) low transverse momentum J/ψ elliptic flow (v_2) in order to study in more details the
 1970 recombination mechanism (ii) J/ψ directed flow (v_1) that will allow us to study the initial
 1971 tilt of the bulk medium (iii) suppression of the loosely bounded $\psi(2S)$ state to explore the
 1972 temperature profile of the medium.

1973 An important observable for studying the properties of the deconfined medium is the
 1974 second order flow harmonic of the Fourier expansion of the azimuthal distribution of the
 1975 produced hadrons, the elliptic flow coefficient v_2 . Similarly as in the case of light hadrons,
 1976 a positive v_2 of D -mesons and electrons from heavy-flavor hadron decays was observed at
 1977 RHIC energies of 54.4 and 200 GeV. Which suggests that the charm quarks may (partially)
 1978 thermalize and participate in the bulk medium collective evolution. On the other hand, the
 1979 v_2 of heavier J/ψ reported by STAR based on the 2010 Au+Au 200 GeV data sample was
 1980 found to be consistent with zero, albeit within large statistical uncertainties and systematic
 1981 uncertainties due to non-flow effects. The precision of the measurement was also not enough
 1982 to distinguish between theoretical model calculations that assume only primordial J/ψ pro-
 1983 duction and the ones that include additional J/ψ production via the recombination. This
 1984 calls for a larger sample of heavy-ion data at 200 GeV, as will be provided by RHIC in 2023
 1985 and 2025, in order to observe a possible non-zero J/ψ v_2 at RHIC energies and put more
 1986 constraints on the J/ψ production models especially regarding its regeneration. Particularly
 1987 important for these studies will be the STAR potential to measure low transverse momentum
 1988 J/ψ with a very good precision. This excellent low- p_T performance at STAR can be achieved
 1989 thanks to its low material budget and great particle identification capabilities.

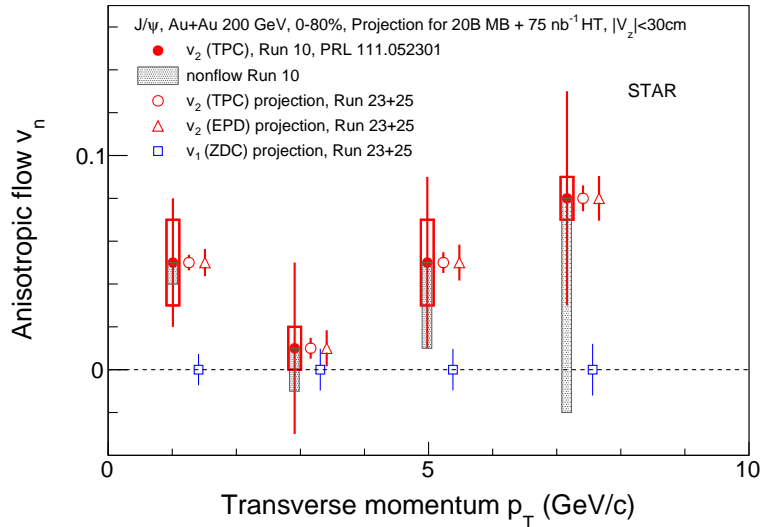


Figure 58: Projections for the J/ψ ($J/\psi \rightarrow e^+e^-$) directed (v_1) and elliptic (v_2) flow vs J/ψ p_T in 0-80% Au+Au collisions at 200 GeV, assuming 20B MB events and HT triggered events corresponding to an integrated luminosity of 75 nb^{-1} with $|V_Z| < 30 \text{ cm}$.

1990 Moreover, the second order Event Plane (EP) can be reconstructed using the new Event
 1991 Plane Detectors (EPD) installed before the 2018 run. It is expected that using the EPD,
 1992 that are forward detectors, will significantly decrease the contribution from the non-flow
 1993 effects and consequently the measurement systematic uncertainties. Also, an inverse of the
 1994 EP resolution enters directly the J/ψ v_2 uncertainty calculation. Thanks to EPD, the res-
 1995 olution of the EP reconstruction at forward rapidity for the J/ψ v_2 measurement at STAR
 1996 will improve. Figure 58 presents statistical projections for the J/ψ v_2 measurement in 0-80%
 1997 central Au+Au collisions assuming 20B MB events and HT triggered events corresponding
 1998 to an integrated luminosity of 75 nb^{-1} . Both cases of the second order EP reconstruction,
 1999 using forward EPD and mid-rapidity TPC detectors, are considered and shown. However,
 2000 measurements for which the TPC is utilized for the EP reconstruction suffer from a sub-
 2001 stantial non-flow contribution which would be greatly reduced by reconstructing the second
 2002 order Event Plane with the EPD. A clear significant improvement in the precision of the
 2003 J/ψ v_2 can be seen across the whole experimentally accessible J/ψ p_T coverage of the previ-
 2004 ous measurement. In addition, the new larger dataset would allow to extend the measured
 2005 p_T range beyond 10 GeV/ c .

2006 Studies of the directed flow v_1 as a function of rapidity provide crucial information to
 2007 understand the initial tilt of the medium produced in heavy-ion collision. The heavy quarks
 2008 are produced in the early stage of a heavy-ion collision and thus are of a particular interest
 2009 for the medium initial asymmetry studies. STAR recently reported the first measurement of
 2010 D-meson v_1 in Au+Au collisions at 200 GeV where the magnitude of the heavy-flavor meson
 2011 v_1 is about 25 times larger than the v_1 for charged kaons. With the 2023-2025 data, STAR
 2012 would have a unique opportunity to also study the v_1 of a bound $c\bar{c}$ state, the J/ψ mesons,

2013 for which even larger directed flow can be expected [219]. In addition to the STAR excellent
 2014 capability to reconstruct low- p_T J/ψ , as discussed above, the iTPC detector completed in
 2015 2018 will improve the momentum resolution and extend the STAR pseudorapidity coverage
 2016 around the mid-rapidity. This will provide a better precision for the slope extraction of the
 2017 v_1 vs y measurement, that quantifies the strength of directed flow. The expected precision
 2018 of a J/ψ v_1 measurement vs p_T at STAR in 2023-2025, assuming 20B MB events and HT
 2019 triggered events corresponding to an integrated luminosity of 75 nb^{-1} , in 0-80% central
 2020 Au+Au collisions at 200 GeV is shown in Fig. 58. Together with the J/ψ v_2 measurements,
 2021 v_1 would provide a more complete picture of the J/ψ production mechanism as well as the
 2022 medium properties in heavy-ion collisions at RHIC.

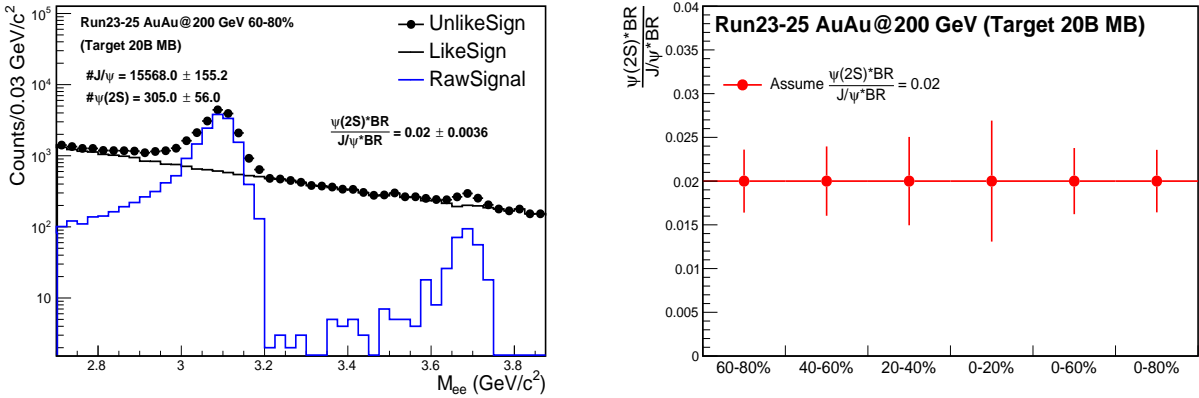


Figure 59: Projections for the J/ψ and $\psi(2S)$ signals in 60-80% Au+Au collisions at 200 GeV and the yield ratio in various centrality bins. The improvement of momentum and dE/dx resolution thanks to the STAR iTPC upgrade have not been taken into account.

2023 $\psi(2S)$ is the most loosely bounded quarkonium state accessible in the heavy-ion collision
 2024 experiments. Its dissociation temperature is predicted to be around or below the critical
 2025 temperature and is much less than that of J/ψ and Υ states. It is more likely to be dis-
 2026 sociated in the early stage and in the core of the fireball and the $\psi(2S)$ survived may have
 2027 significant contribution from the regeneration at later stage in the evolution of the fireball.
 2028 The relative suppression of $\psi(2S)$ and J/ψ is sensitive to the the temperature profile of the
 2029 fireball produced in heavy-ion collisions and its space-time evolution. It is also argued that
 2030 the charmonium formation process from a $c\bar{c}$ pair may be affected by the QGP or the ini-
 2031 tial strong external magnetic field, altering the relative yields among different charmonium
 2032 states [220, 221]. The measurement of $\psi(2S)$ is much more difficult than that of J/ψ due to
 2033 a much smaller production cross-section and dilepton decay branching ratio, resulting in a
 2034 very low signal-to-background ratio. The ALICE Collaboration successfully measured the
 2035 relative suppression of $\psi(2S)$ and J/ψ in Pb+Pb collisions at forward rapidity [222], and the
 2036 ATLAS and CMS Collaboration published the relative suppression in Pb+Pb collisions at
 2037 mid-rapidity and high p_T [223, 224]. The attempt of measuring $\psi(2S)$ suppression in heavy-
 2038 ion collisions at RHIC has no success as to date. The low material budget and excellent

2039 particle identification capability of the STAR detector together with the large data sample
 2040 in 2023 and 2025 will provide a unique opportunity to measure the suppression of $\psi(2S)$ at
 2041 low p_T and mid-rapidity in heavy-ion collisions. Figure 59 shows the projections of $\psi(2S)$
 2042 signal and the yield ratio of $\psi(2S)$ and J/ψ in Au+Au collisions assuming 20B MB events.
 2043 The improvement of momentum and dE/dx resolution thanks to the STAR iTPC upgrade
 2044 will further enhance the signal-to-background ratio and the significance of $\psi(2S)$ signal.

2045 **2.4.5 Jet Probes**

2046 Precise jet quenching measurements with reconstructed jets over a broad kinematic range
 2047 at RHIC are essential to meet the goal of the NSAC 2015 Long Range Plan, to “probe the
 2048 inner workings of the QGP” [196]. For example, the dependence of jet energy loss on the jet
 2049 p_T and the resolution or angular scale tagged by jet substructure observables are key tools
 2050 to discriminate jet quenching mechanisms [225–228]. In addition, the measurement of jet
 2051 acoplanarity as a probe of in-medium jet scattering is most sensitive at low jet p_T to a given
 2052 momentum transfer and to medium-induced radiative effects [229], and is least affected by
 2053 background due to vacuum Sudakov radiation [230].

2054 The highest-statistics STAR Au+Au collision datasets currently available were recorded
 2055 in 2014 and 2016, with the integrated luminosity sampled by STAR BEMC triggers shown
 2056 in Table 10. Preliminary jet analyses using the 2014 dataset are discussed in section 1.1.1
 2057 and are moving towards publication. STAR will continue to exploit these rich datasets to
 2058 carry out high-precision measurements with fully reconstructed jets over the full RHIC phase
 2059 space.

2060 The 2023 and 2025 runs will generate another significant increase in sampled integrated
 2061 luminosity, enabling a third generation of STAR heavy-ion jet measurements that are yet
 2062 more differential and precise. STAR’s open geometry near the beam pipe allows it to utilize
 2063 a wide range in the vertex position along the beam direction (z_{vertex}) for jet analyses, thereby
 2064 utilizing the RHIC luminosity efficiently. Optimization of the z_{vertex} range used in an analysis
 2065 entails a balance between statistical precision and complexity of corrections, with the latter
 2066 influencing the systematic uncertainty of the measurement. Recent STAR jet measurements
 2067 in Au+Au collisions have employ two different z-vertex cuts: the inclusive charged-particle
 2068 jet analysis [7] utilizes $|z_{\text{vertex}}| < 30$ cm, whereas the $\gamma_{\text{dir}} + \text{jet}$ analysis [10] utilizes $|z_{\text{vertex}}| < 70$
 2069 cm. With the success of the $\gamma_{\text{dir}} + \text{jet}$ analysis in analyzing this broad z_{vertex} range with good
 2070 systematic precision, we will re-examine this cut for future jet measurements, including the
 2071 inclusive jet analysis. In section 2.4 we present the sampled integrated luminosity in 2023
 2072 and 2025 for both the 30 cm and 70 cm z_{vertex} cuts. The following physics performance
 2073 projections are based on the 70 cm cut, using the cumulative sampled integrated luminosity
 2074 for Runs 2014, 2016, 2023 and 2025 together. For $|z_{\text{vertex}}| < 70$ cm, this total is 110 nb^{-1} ,
 2075 which is a factor 11 increase in trigger statistics relative to the current analyses based on
 2076 Run 14 data.

2077 To quantify the effect of this marked increase in integrated luminosity, we utilize two
 2078 mature jet measurements currently in progress and discuss their expected improvement
 2079 with enhanced integrated luminosity. These analyses are the semi-inclusive distribution

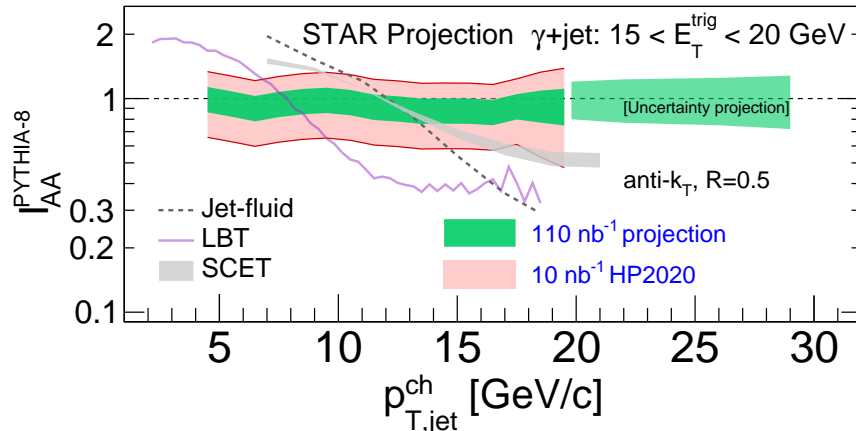


Figure 60: Ratio of semi-inclusive distributions of charged-particle jets (anti- k_T , $R = 0.5$) recoiling from a direct-photon trigger with $15 < E_T < 20$ GeV for central Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV measured by STAR (numerator) and p+p collisions simulated by PYTHIA (denominator). The pink band shows the cumulative uncertainty for the current analysis based on 10 nb^{-1} [10], while the green band shows the projected uncertainty for 110 nb^{-1} . Theory calculations are discussed in [10].

2080 of charged-particle jets recoiling from a high- E_T direct-photon trigger ($\gamma_{\text{dir}} + \text{jet}$); and the
 2081 differential measurement of energy loss for jet populations selected by varying a substructure
 2082 metric. Since these analyses are mature, their analysis methodologies and correction schemes
 2083 are optimized, so that their projections based on increased statistics are meaningful. We do
 2084 not imply that these will be the only flagship measurements that STAR will make with the
 2085 ‘23/‘25 datasets; in addition we will focus for instance on fully reconstructed jets and uti-
 2086 lize additional substructure observables, including those not yet developed. However, these
 2087 analyses are most mature at present, and therefore provide the most accurate projections of
 2088 gain in precision.

2089 **Semi-inclusive $\gamma_{\text{dir}} + \text{jet}$ measurement**

2090 Figure 60 shows I_{AA} for fully-corrected semi-inclusive distributions of charged-particle jets
 2091 (anti- k_T , $R = 0.5$) recoiling from a direct-photon trigger with $15 < E_T < 20$ GeV in central
 2092 Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV, for the current analysis based on 10 nb^{-1} [10] within
 2093 $|z_{\text{vertex}}| < 70$ cm, and projected for 110 nb^{-1} . Significant reduction in the uncertainty band
 2094 is seen to result from the increase in integrated luminosity, together with significant increase
 2095 in kinematic reach.

2096 Note that the projection to 110 nb^{-1} only takes into account the increase in statistical
 2097 precision, and assumes that the systematic uncertainty remains the same. The reduction in
 2098 width of the uncertainty band is therefore less than the factor $1/\sqrt{11}$ expected from statistical
 2099 considerations alone, indicating the magnitude of the systematic component. Experience
 2100 shows that systematic uncertainty can also be improved by an increase in statistical precision,
 2101 since additional and more differential systematic studies can be carried out. However, it is

2102 not possible to project that improvement with confidence at present; thus Fig. 60 should
 2103 therefore be regarded as a conservative estimate of the improvement in precision of this
 2104 measurement channel with the projected integrated luminosity increase.

2105 Broadening of the back-to-back di-jet angular distribution due to jet scattering from con-
 2106 stituents of the QGP (medium-induced acoplanarity) was proposed over three decades ago
 2107 as a diagnostic probe of the QGP [231, 232]. While the physical picture of this process is
 2108 intuitive and compelling, such measurements are extremely challenging, because of both the
 2109 large jet backgrounds in heavy ion collisions, and the large contribution of vacuum QCD pro-
 2110 cesses (Sudakov radiation) to the di-jet angular distribution [230]. In addition, minimization
 2111 of these two effects nominally drives the experimentalist in opposite directions: minimization
 2112 of background effects prefers larger $p_{T,\text{jet}}$, whereas minimization of Sudakov broadening and
 2113 higher sensitivity to medium-induced momentum transfer prefers lower $p_{T,\text{jet}}$ [230]. These
 2114 contradictory requirements were resolved only with the development of absolutely normalized
 2115 semi-inclusive jet measurements in heavy-ion collisions, with statistically-based background
 2116 corrections that enable measurements at low $p_{T,\text{jet}}$ and large R [1, 233].

2117 The first generation searches for medium-induced acoplanarity using this approach did
 2118 not exhibit a significant signal above background [1, 233], though with limited statistical
 2119 precision. Higher-precision measurements of medium-induced acoplanarity over a broad $p_{T,\text{jet}}$
 2120 range – including low $p_{T,\text{jet}}$ – are clearly of great interest at both RHIC and the LHC. Such
 2121 measurements may provide a direct probe of \hat{q} [230], or evidence of large-angle jet scattering
 2122 off of quasi-particles in the QGP [234]. Consideration of higher-order processes suggests that
 2123 the contribution of radiative corrections to this distribution may be negative [229], thereby
 2124 narrowing rather than broadening the recoil jet azimuthal distribution; a recent preliminary
 2125 measurement by the ALICE Collaboration at the LHC may indeed have observed such an
 2126 effect [235]. Complementary measurements of medium-induced acoplanarity over wide phase
 2127 space by STAR at RHIC and ALICE at the LHC, using similar instrumentation and similar
 2128 analysis techniques, promise to provide strong constraints on theoretical descriptions of this
 2129 fundamental process, providing new insight into the inner workings of the QGP [196].

2130 Figure 61 shows the semi-inclusive distribution of the azimuthal separation between a
 2131 direct-photon trigger with $15 < E_T < 20$ GeV and a charged-particle jet (anti- k_T , $R = 0.5$)
 2132 with $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/ c , in central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV with only
 2133 statistical error bars, based on the analysis described in [10]. Azimuthal smearing of this
 2134 observable due to uncorrelated background is small, and such acoplanarity measurements
 2135 are therefore strongly statistics-dominated [1, 233]. The grey points are from the current
 2136 analysis based on 10 nb^{-1} [10], whereas the red points correspond to 110 nb^{-1} . A marked
 2137 increase in measurement precision is projected, with corresponding qualitative increase in
 2138 physics impact. Similar improvements in precision for this observable are expected at the
 2139 LHC in Run 3, due to detector upgrades and enhanced machine luminosity [168].

2140 **Jet measurement with a varying substructure metric**

2141 An important facet of the third generation of STAR jet measurements is a systematic
 2142 exploration of parton energy loss based on controlled variation of the jet shower topology.

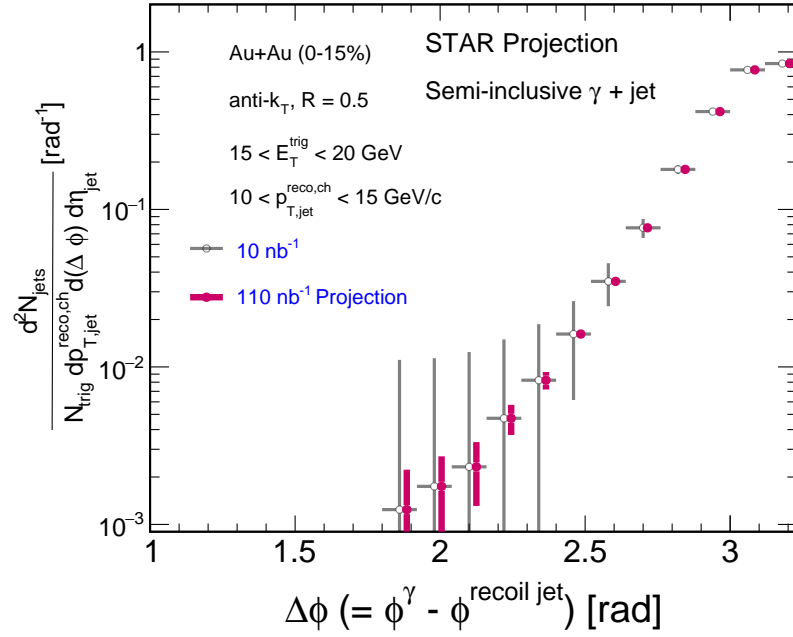


Figure 61: Semi-inclusive azimuthal distribution of charged jets (anti- k_T , $R=0.5$) with $10 < p_{T,\text{jet}}^{\text{ch}} < 15$ GeV/ c recoiling from a direct photon trigger with $15 < E_T^{\text{trig}} < 20$ GeV, in central Au+Au collisions at $\sqrt{s_{\text{NN}}} = 200$ GeV. Grey points: current analysis with 10 nb^{-1} [10]; red points: projection for 110 nb^{-1} . Error bars are statistical only.

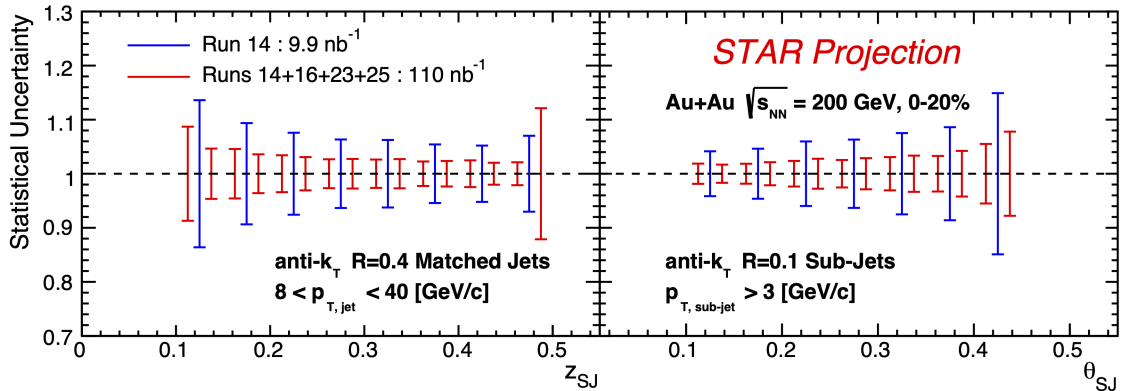


Figure 62: Two-panel figure showing statistical uncertainty for the two-subject observables in 0-20% central Au+Au collisions for 10 nb^{-1} in blue and projection for 110 nb^{-1} in red.

2143 Jet evolution produces a unique pattern of radiation in both angle and momentum, and
 2144 jet substructure observables are a broad class of measurements of combinations of the jet
 2145 constituents' angle and/or momentum via algorithms or correlations. As the jet undergoes
 2146 interactions with the medium, jet substructure modification for a given jet energy (e.g. com-
 2147 paring the heavy-ion results to those in p+p collisions) has been used as a way to access the
 2148 microscopic properties of the medium. By selecting on jets based on their substructure, we
 2149 can study how a particular class of jets interacts with the medium to determine the effects
 2150 of e.g. color coherence, dead cone, etc. on parton-medium interactions. In other words, the
 2151 STAR jet program for Runs 23–25 will focus on jet substructure as a jet-tagger.

2152 Recent theory calculations have shown significant differences between energy loss signa-
 2153 tures for jets that are perceived by the medium as a single or multiple color charges [227].
 2154 Algorithms such as SoftDrop and sub-jets [236, 237] provide observables that correspond to
 2155 the splitting within jets via momentum fractions and an inherent angular scale which then
 2156 serve as a proxy for the resolution scale in the medium. This is often referred to as coherent
 2157 vs. de-coherent energy loss where the coherent length of the medium is related to its temper-
 2158 ature and \hat{q} [?]. By isolating population of jets based on their substructure, one can directly
 2159 probe energy loss for varying resolution scales. The integrated luminosity from Runs '23/'25
 2160 datasets will not only provide a substantial increase in statistics in the current measure-
 2161 ments of jet substructure, they also make the phase space available for rare processes such
 2162 as wide angle emissions from high- p_T jets. This enables STAR to extend our current mea-
 2163 surements of differential energy loss [2] to fine binning in the opening angles and momentum
 2164 fractions, as shown in Fig. 62. The current resolution of $\delta\theta_{\text{SJ}} = 0.1$ [2] is predominantly due
 2165 to statistical limitations in our older dataset sample. The significant increase in integrated
 2166 luminosity for Runs 23-25, coupled with excellent tracking resolution of the STAR TPC will
 2167 reduce the opening angles resolution to 0.025 and have significant population of jets where
 2168 we can further identify and select jet topologies in both z and θ and study energy loss in
 2169 a three-dimensional fashion. By extending to high energy splittings within jets, at varied

2170 opening angles, we can probe earlier formation times whereby vacuum-like emissions and
2171 medium induced radiations are expected to occur.

2172 STAR is uniquely situated to perform high impact differential measurements of parton
2173 energy loss starting from the unbiased, semi-inclusive jet population, to a topologically spe-
2174 cial population of jets, selected via jet substructure observables.

2175 Given the unique nature of jet-medium interactions at RHIC, with the jet and sub-jet
2176 scales sufficiently closer to the medium scale, the aforementioned measurements bolster the
2177 importance of the STAR jet program with the goal of extracting the microscopic properties
2178 of the QGP as outlined in the 2015 LRP.

2179 3 Cold QCD Physics with $p^\uparrow p^\uparrow$ and $p^\uparrow + A$ Collisions at 2180 510 and 200 GeV

2181 The exploration of the fundamental structure of strongly interacting matter has always
2182 thrived on the complementarity of lepton scattering and purely hadronic probes. As the
2183 community eagerly anticipates the future Electron Ion Collider, an outstanding scientific
2184 opportunity remains to complete “must-do” measurements in p+p and $p+A$ physics during
2185 the final years of RHIC. These measurements will be essential if we are fully to realize the
2186 scientific promise of the EIC, by providing a comprehensive set of measurements in hadronic
2187 collisions that, when combined with future data from the EIC, will establish the validity
2188 and limits of factorization and universality. The Run-22 and Run-24 program outlined here
2189 will, on the one hand, lay the groundwork for the EIC, both scientifically and in terms of
2190 refining the experimental requirements of the physics program at the EIC, and thus are
2191 the natural next steps on the path to the EIC. On the other hand, while much of the
2192 physics in this program is unique to proton-proton and proton-nucleus collisions and offers
2193 discovery potential on its own, when combined with data from the EIC it will provide a
2194 broad foundation to a deeper understanding of fundamental QCD.

2195 The separation between the intrinsic properties of hadrons and interaction-dependent
2196 dynamics, formalized by the concept of factorization, is a cornerstone of QCD and largely
2197 responsible for the predictive power of the theory in many contexts. While this concept
2198 and the associated notion of universality of the quantities that describe hadron structure
2199 have been successfully tested for unpolarized and, to a lesser extent, longitudinally polarized
2200 parton densities, its experimental validation remains an unfinished task for much of what the
2201 EIC is designed to study – the three-dimensional structure of the proton and the physics of
2202 dense partonic systems in heavy nuclei. To establish the validity and limits of factorization
2203 and universality, it is essential to have data from *both* lepton-ion and proton-ion collisions,
2204 with experimental accuracy that makes quantitative comparisons meaningful.

2205 Beginning in Run-22, STAR will be in a unique position to provide this essential p+p and
2206 $p + A$ data. A full suite of forward detectors will be installed, providing excellent charged-
2207 particle tracking at high pseudorapidity ($2.5 < \eta < 4$) for the first time, coupled with both
2208 electromagnetic and hadronic calorimetry. This will enable STAR to explore the interesting
2209 regimes of high- x (largely valence quark) and low- x (primarily gluon) partonic physics with
2210 unparalleled precision. In addition, mid-rapidity detector upgrades motivated primarily by
2211 the BES-II program – the iTPC, eTOF, and EPD systems – will substantially extend STAR’s
2212 already excellent kinematic reach and particle identification capabilities beyond those that
2213 existed during previous p+p and $p + A$ runs.

2214 For the case of p+p spin physics, it’s important to recognize the complementary roles
2215 that will be played by Run-22 at 510 GeV and Run-24 at 200 GeV. The combination of
2216 510 GeV p+p collisions and the STAR Forward Upgrade will provide access to forward jet
2217 physics at perturbative scales, thereby enabling measurements at the highest and lowest x
2218 values. In parallel, mid-rapidity measurements at 510 and, especially, 200 GeV will interpo-
2219 late between the high and low x values, with significant overlaps to probe evolution effects

2220 and provide cross-checks. Together, the two runs will allow STAR to measure fundamental
2221 proton properties, such as the Sivers and transversity distributions, over nearly the entire
2222 range $0.005 < x < 0.5$.

2223 Run-24 will also provide outstanding opportunities to probe fundamental questions re-
2224 garding QCD in cold nuclear matter. The STAR Forward Upgrade will enable an extensive
2225 suite of measurements probing the quark-gluon structure of heavy nuclei and the regime of
2226 low- x non-linear gluon dynamics, as predicted by saturation models. STAR will also ex-
2227 plore how a nucleus, serving as a color filter, modifies the propagation, attenuation, and
2228 hadronization of colored quarks and gluons.

2229 For these reasons, STAR requests at least 16 weeks of polarized $p+p$ data-taking at $\sqrt{s} =$
2230 510 GeV in Run-22. All data taking will involve proton beams polarized transversely relative
2231 to their momentum direction in order to focus on those observables where factorization,
2232 universality, and/or evolution remain open questions, with spins aligned either vertically or
2233 radially at the STAR IR (still to be determined through more detailed simulation studies).
2234 Based on the latest guidance from CAD, and mindful of ‘lessons learned’ in previous $p + p$
2235 runs at full energy, we will ask for luminosity-leveling of the collision rate to maximize the
2236 efficiency of our main tracking detectors. Assuming we will have running conditions similar to
2237 those achieved in Run-17, we expect to sample at least 400 pb^{-1} for our rare / non-prescaled
2238 triggers.

2239 STAR also requests at least 11 weeks of polarized $p + p$ data-taking at $\sqrt{s} = 200 \text{ GeV}$
2240 and 11 weeks of polarized $p+\text{Au}$ data-taking at $\sqrt{s_{NN}} = 200 \text{ GeV}$ during Run-24. Similar
2241 to Run-22, all of the running will involve transversely polarized protons, with the choice
2242 between vertical or radial polarization to be determined during the coming year. Based on
2243 recent (08-21-20) C-AD guidance, we expect to sample at least 235 pb^{-1} of $p+p$ collisions
2244 and 1.3 pb^{-1} of $p+\text{Au}$ collisions. These totals represent 4.5 times the luminosity that STAR
2245 sampled during transversely polarized $p + p$ collisions in Run-15, and 3 times the luminosity
2246 that STAR sampled during transversely polarized $p+\text{Au}$ collisions in Run-15.

2247 **3.1 Run-22 Request for $p^\uparrow p^\uparrow$ Collisions at 510 GeV**

2248 **3.1.1 Inclusive transverse spin asymmetries at forward rapidities**

2249 The experimental study of spin phenomena in nuclear and particle physics has a long history
2250 of producing important, and often surprising, results. Attempts to understand such data
2251 have pushed the field forward, forcing the development of both new theoretical frameworks
2252 and new experimental techniques. Recent and ongoing detector upgrades at STAR, at mid-
2253 and forward-rapidity, coupled with the versatility of RHIC, will allow us to gain new insights
2254 into long-standing puzzles, and to probe more deeply the complexities of emergent behavior
2255 in QCD.

2256 Results from PHENIX and STAR have shown that large transverse single-spin asym-
2257 metries (TSSA) for inclusive hadron production, first seen in $p+p$ collisions at fixed-target
2258 energies and modest p_T , extend to the highest RHIC center-of-mass energies, $\sqrt{s} = 500 \text{ GeV}$,
2259 and surprisingly large p_T . Figure 63 summarizes the world data for the inclusive pion asym-

2260 metries A_N as a function of Feynman- x . The asymmetries are seen to be nearly independent
 2261 of \sqrt{s} over the very wide range of roughly 5 to 500 GeV.

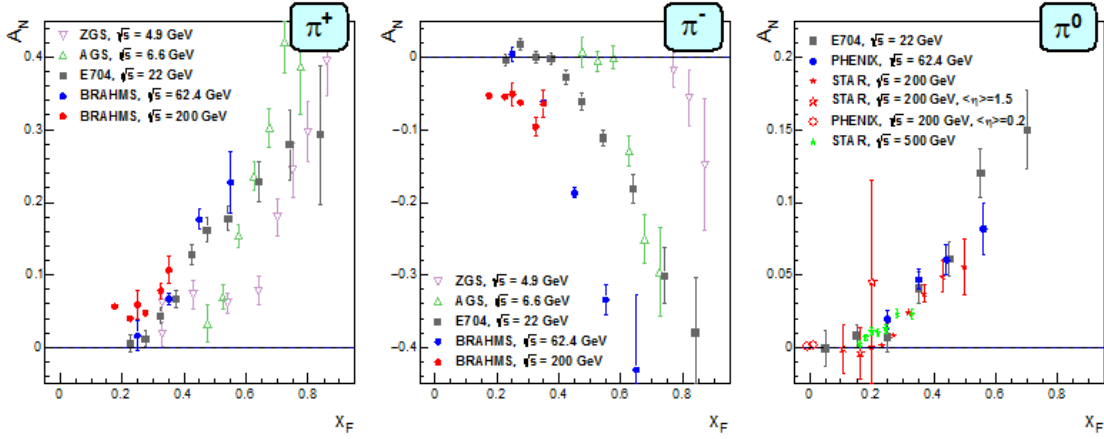


Figure 63: Transverse single-spin asymmetry A_N measurements for charged and neutral pions at different center-of-mass energies as a function of Feynman- x .

2262 To understand the observed TSSAs, one needs to go beyond the conventional leading-
 2263 twist (twist-2) collinear parton picture for the hard-scattering processes. Two theoretical
 2264 formalisms have been developed to try to explain these sizable asymmetries in the QCD
 2265 framework: transverse-momentum-dependent (TMD) parton distribution and fragmentation
 2266 functions, such as the Sivers and Collins functions; and transverse-momentum-integrated
 2267 (collinear) quark-gluon-quark correlations, which are twist-3 distributions in the initial state
 2268 proton or in the fragmentation process. For many of the experimentally accessible spin
 2269 asymmetries, several of these functions can contribute, and need to be disentangled in order
 2270 to understand the experimental data in detail, in particular the observed p_T dependence.
 2271 These functions manifest their spin dependence either in the initial state—for example, the
 2272 Sivers distribution and its twist-3 analog, the Efremov-Teryaev-Qui-Sterman (ETQS) func-
 2273 tion [238]—or in the final state via the fragmentation of polarized quarks, such as in the
 2274 Collins function and related twist three function $\hat{H}_{FU}(z, z_z)$.

2275 Incorporating the fragmentation term within the collinear twist-3 approach demonstrated
 2276 the ability of this formalism to describe the large values of A_N for π^0 production observed at
 2277 RHIC [239]. In this work, the relevant (non-pole) 3-parton collinear fragmentation function
 2278 $\hat{H}_{FU}(z, z_z)$ was fit to the RHIC data. The so-called soft-gluon pole term, involving the ETQS
 2279 function $T_{q,F}(x_1, x_2)$, was also included by fixing $T_{q,F}$ through its well-known relation to the
 2280 TMD Sivers function f_{1T}^\perp . The authors obtained a very good description of the data due to
 2281 the inclusion of the non-pole fragmentation function and based on this work they were able
 2282 to make predictions for π^+ and π^- production asymmetries A_N at the forward rapidities
 2283 covered by the STAR upgrades, $2.5 < \eta < 4$. The results are shown in Fig. 64 for $\sqrt{s} = 200$
 2284 and 500 GeV for two rapidity ranges, $2 < \eta < 3$ and $3 < \eta < 4$.

2285 In Run-22, with the full suite of forward tracking detectors and calorimetry installed,
 2286 STAR will for the first time be able to map out inclusive charged-hadron asymmetries up to

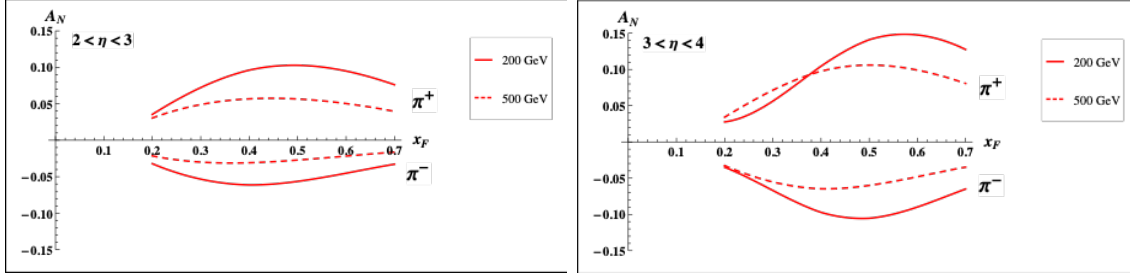


Figure 64: Predictions for A_N for π^+ and π^- production over the ranges $2 < \eta < 3$ (left) and $3 < \eta < 4$ (right) at $\sqrt{s} = 200$ GeV (solid lines) and 500 GeV (dashed lines). The $\sqrt{s} = 200$ GeV BRAHMS A_N data for charged pions cover up to x_F of 0.3.

2287 the highest energies achievable at RHIC and at these forward rapidities. It would be very
 2288 interesting to confirm that these asymmetries are indeed largely independent of center-of-
 2289 mass energy. The measurements of A_N for charged hadrons, together with analogous data
 2290 (from Run-22 as well as previous STAR runs) on A_N for direct photons and neutral pions,
 2291 should provide the best data set in the world to constrain the evolution and flavor dependence
 2292 of the twist-3 ETQS distributions and to determine if the 3-parton collinear fragmentation
 2293 function H_{FU} is the main driver of the large forward inclusive asymmetries.

2294 3.1.2 Sivers and Efremov-Teryaev-Qui-Sterman Function

2295 There is great theoretical interest in testing the relation between the ETQS correlation
 2296 functions and the Sivers function. As discussed above, both the Sivers and the ETQS
 2297 functions encapsulate partonic spin correlations within the proton, but they are formally
 2298 defined in different frameworks. While the Sivers function is a TMD quantity that depends
 2299 explicitly on spin-dependent transverse partonic motion k_T , the ETQS function is a twist-3
 2300 collinear distribution, in which SSAs are generated through soft collinear gluon radiation

2301 Measurements of forward jet production from the ANDY collaboration [240] indicated
 2302 rather small asymmetries. This was argued to be consistent with the idea that the twist-3
 2303 parton correlation functions for up and down valence quarks should cancel, because their
 2304 behavior reflects the Sivers functions extracted from fits to the SIDIS data that demonstrate
 2305 opposite sign, but equal magnitude, up and down quark Sivers functions. Preliminary STAR
 2306 results on charge-tagged dijets at mid-rapidity (see Fig. 29) support this interpretation, with
 2307 the caveat that the measured observable (a spin-dependent $\langle k_T \rangle$) is defined in the TMD, and
 2308 not the twist-3, framework.

2309 To better test quantitatively the relation between the two regimes, one can measure spin
 2310 asymmetries for jets which are *intentionally* biased towards up or down quark jets via detec-
 2311 tion of a high- z charged hadron within the jet. Higher-twist calculations of jet asymmetries
 2312 based on the Sivers function predict sizeable effects for these flavor-enhanced jets. With the
 2313 suite of new forward detectors installed at STAR, full jet reconstruction, along with iden-
 2314 tification of a high- z hadron of known charge sign, will be possible at high pseudorapidity.
 2315 Using realistic jet smearing in a forward calorimeter and tracking system, and requiring a

2316 charged hadron with $z > 0.5$, the asymmetries can be separated and compared to the pre-
 2317 dictions for the Sivers function based on current SIDIS data. The expected uncertainties,
 2318 plotted at the predicted values, can be seen in Figure 65. Dilutions by underlying event and
 2319 beam remnants were taken into account. The simulations have assumed only an integrated
 2320 luminosity of 100 pb^{-1} at $\sqrt{s} = 200 \text{ GeV}$, which is significantly lower than what is currently
 2321 expected for a 200 GeV polarized p+p run in 2024. The same measurement is possible at
 2322 500 GeV.

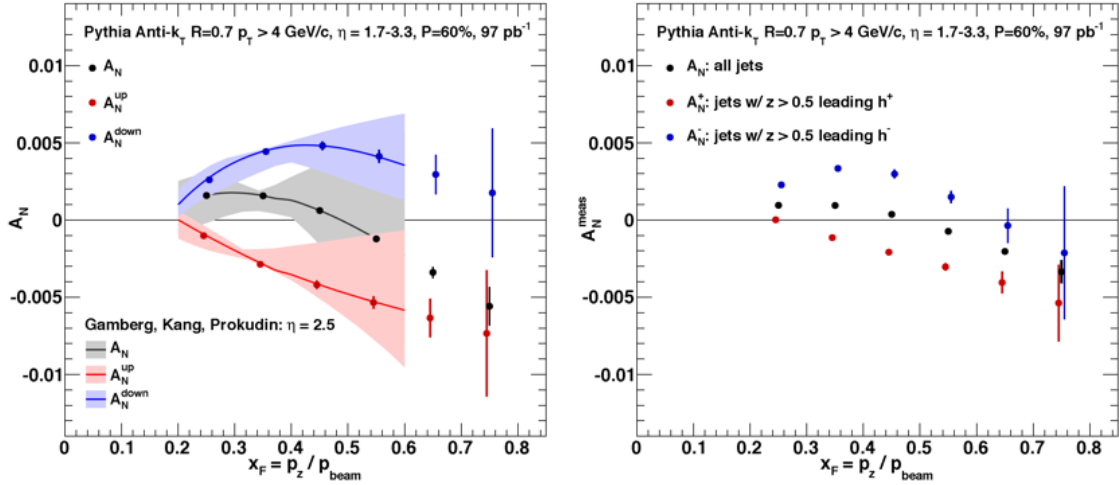


Figure 65: Left: up quark (red points), down quark (blue points) and all jet (black points) single spin asymmetries as a function of x_F as calculated by the ETQS based on the SIDIS Sivers functions. Right: Expected experimental sensitivities for jet asymmetries tagging in addition a positive hadron with z above 0.5 (red points), a negative hadron with z above 0.5 (blue points) or all jets (black) as a function of x_F . Note: these figures are for 200 GeV center-of-mass energy proton collisions – the 500 GeV results are expected to be qualitatively similar.

2323 In a TMD framework, the Sivers effect manifests itself as a correlation (a triple product)
 2324 between the transverse momentum of a parton (\vec{k}_T) with momentum fraction x , and the
 2325 transverse spin (\vec{S}) of a polarized proton moving in the longitudinal (\vec{p}) direction. Thus,
 2326 for transversely polarized protons, the Sivers effect probes whether the k_T of the constituent
 2327 quarks is preferentially oriented in a direction perpendicular to both the proton momentum
 2328 and its spin. Momentum conservation then implies that the two jets in the final state will
 2329 not emerge back-to-back on average, but instead will ‘tilt’ in the direction of the summed
 2330 k_T of the initial state partons. Moreover, the (average) tilt of interest will reverse direction
 2331 under a ‘flip’ of the proton spin; a spin-dependent $\langle k_T \rangle$ can then be extracted by associating
 2332 the azimuthal opening angle of the jet pair with this tilt.

2333 STAR carried out an earlier measurement of this transverse single-spin asymmetry using
 2334 a dijet dataset with $\sim 1 \text{ pb}^{-1}$ of integrated luminosity [112], and found it to be consistent
 2335 with zero within 2σ . An ongoing, and much improved analysis, has been described in the
 2336 Highlights section of this document, Sect. 1.3. Perhaps most significantly, the jets were
 2337 sorted according to their net charge Q , calculated by summing the signed momentum of all

2338 particle tracks with $p > 0.8$ GeV, to minimize underlying event contributions, yielding jet
 2339 samples with enhanced contributions from u quarks (positive Q) and d quarks (negative Q),
 2340 with a large set near $Q = 0$ dominated by gluons. Simple kinematics allow for conversion
 2341 from the spin-dependent ‘tilt’ of the dijet pair to a value of k_T on an event-by-event basis;
 2342 these are then sorted by the Q of the jet and binned by the summed pseudorapidities of
 2343 the outgoing jets, $\eta^{\text{total}} \equiv \eta_3 + \eta_4$. Because the contributions of different partons (u , d , all
 2344 else) to $\langle k_T \rangle$ vary with both Q and also η^{total} , in a way that can be estimated robustly using
 2345 simulation, the data can be inverted to yield values of $\langle k_T \rangle$ for the individual partons, though
 2346 with coarser binning in η^{total} .

2347 With the new forward detectors in place, along with the enhanced reach in η afforded by
 2348 the iTPC, this technique can be expanded in Run-22 to cover pseudorapidities at STAR from
 2349 roughly -1 to +4, though with a gap at $1.5 < \eta < 2.5$. Despite this gap, values of $\langle k_T \rangle$ can be
 2350 extracted for u and d quarks for η^{total} ranging from ~ -1.5 to as high as 7 with reasonable
 2351 statistics. This latter regime will probe $2 \rightarrow 2$ hard scattering events in which $x_1 \gg x_2$,
 2352 *i.e.*, a sample enriched in valence quarks interacting with low- x gluons. Such measurements,
 2353 exploiting the full kinematic reach of STAR, will not only allow precise determinations of
 2354 the average transverse partonic motion, $\langle k_T \rangle$, exhibited by individual partonic species in
 2355 the initial state, but will provide important information on the x dependence of the proton
 2356 Sivers functions.

2357 Collisions at $\sqrt{s} = 510$ GeV will also allow STAR to continue our successful program
 2358 to study the evolution and sign change of the Sivers function. By focusing on interactions
 2359 in which the final state involves only weakly interacting particles, and hence the transverse
 2360 partonic motion (in a TMD framework) or the collinear gluon radiation (in twist-3) must be
 2361 in the initial state, one can test for the predicted sign change in A_N relative to interactions
 2362 in which these terms must appear in the final state, such as SIDIS measurements. The
 2363 improved tracking capabilities provided by the iTPC upgrade will allow us to push our mid-
 2364 rapidity W^\pm and Z^0 measurements to larger rapidity $y_{W/Z}$, a regime where the asymmetries
 2365 are expected to increase in magnitude and the anti-quark Sivers’ functions remain largely
 2366 unconstrained. Figure 66 demonstrates the expected precision of asymmetry measurements
 2367 after data from the 2017 run has been fully analyzed. In addition to the noted extension
 2368 of our kinematic reach, an additional 16 or more weeks of beam time at $\sqrt{s} = 510$ GeV in
 2369 Run-22 would increase our data set by more than a factor of 2. This experimental accuracy
 2370 would significantly enhance the quantitative reach of testing the limits of factorization and
 2371 universality in lepton-proton and proton-proton collisions.

2372 3.1.3 Transversity, Collins Function and Interference Fragmentation Function

2373 A complete picture of nucleon spin structure at leading twist must include contributions
 2374 from the unpolarized and helicity distributions, as well as those involving transverse polar-
 2375 ization, such as the transversity distribution [243–245]. The transversity distribution can
 2376 be interpreted as the net transverse polarization of quarks within a transversely polarized
 2377 proton. The difference between the helicity and transversity distributions for quarks and
 2378 antiquarks provides a direct, x -dependent connection to nonzero orbital angular momen-

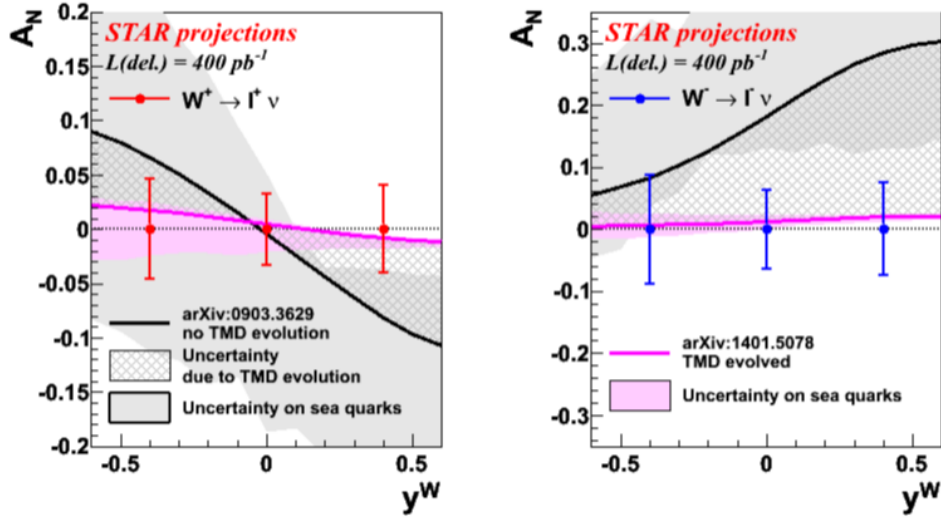


Figure 66: Projected uncertainties for transverse single-spin asymmetries of W^\pm and Z^0 bosons as functions of rapidity for a delivered integrated luminosity of 400 pb^{-1} and an average beam polarization of 55%. The solid light gray and pink bands represent the uncertainties on the KQ [241] and EIKV [242] known sea-quark Sivvers functions. The crosshatched dark grey region indicates the current uncertainty in the theoretical predictions due to TMD evolution.

2379 tum components in the wave function of the proton [246]. Recently, the first lattice QCD
 2380 calculation of the transversity distribution has been performed [247]. In addition, the mea-
 2381 surement of transversity has received substantial interest as a means to access the so-called
 2382 ‘tensor charge’ of the nucleon, defined as the integral over the valence quark transversity:
 2383 $\delta q^a = \int_0^1 [\delta q^a(x) - \delta \bar{q}^a(x)] dx$ [244, 248]. Measuring the tensor charge is very important for
 2384 several reasons. First, it is an essential and fundamental quantity to our understanding of
 2385 the spin structure of the nucleon. Also, the tensor charge can be calculated on the lattice
 2386 with comparatively high precision, due to the valence nature of transversity, and hence is
 2387 one of the few quantities that allow us to compare experimental results on the spin structure
 2388 of the nucleon directly to *ab initio* QCD calculations. Finally, the tensor charge describes
 2389 the sensitivity of observables in low-energy hadronic reactions to beyond the standard model
 2390 physics processes with tensor couplings to hadrons. Examples are experiments with ultra-
 2391 cold neutrons and nuclei.

2392 Transversity is difficult to access due to its chiral-odd nature, requiring the coupling of
 2393 this distribution to another chiral-odd distribution. Semi-inclusive deep-inelastic scattering
 2394 (SIDIS) experiments have successfully probed transversity through two channels: asym-
 2395 metric distributions of single pions, convoluting the TMD transversity distribution with
 2396 the TMD Collins fragmentation function, and azimuthally asymmetric distributions of di-
 2397 hadrons, coupling transversity to the so-called “interference fragmentation function” (IFF)
 2398 in the framework of collinear factorization. Yet in spite of a wealth of lepton-scattering

2399 data, the kinematic reach of existing SIDIS experiments limits the precision with which the
2400 proton's transversity can be extracted, as the range of Bjorken- x values that can be accessed
2401 does not extend above $x \sim 0.3$

2402 In hadronic collisions, the k_T integrated quark transversity distribution may be accessed
2403 via two channels. The first is the single spin asymmetry of the azimuthal distribution of
2404 hadrons in high energy jets [108]. In the jet+hadron channel, the collinear transversity
2405 distribution couples to the TMD Collins function [108, 109]. This makes p+p collisions a
2406 more direct probe of the Collins fragmentation function than Collins asymmetries in SIDIS
2407 [108], where a convolution with the TMD transversity distribution enters. This also makes
2408 the Collins asymmetry in p+p collisions an ideal tool to explore the fundamental QCD
2409 questions of TMD factorization, universality, and evolution. The second channel is the
2410 single spin asymmetry of pion pairs, where transversity couples to the collinear interference
2411 fragmentation function [249]. STAR mid-rapidity IFF data [250] have been included in the
2412 first extraction of transversity from SIDIS and proton-proton IFF asymmetries [251].

2413 The universality of TMD PDFs and fragmentation functions in p+p collisions has been an
2414 open question. General arguments [252, 253] have shown that factorization can be violated
2415 in hadron-hadron collisions for TMD PDFs like the Sivers function, though very recent
2416 calculations indicate the violations might be quite small [254, 255]. In contrast, while there
2417 is no general proof that the Collins effect in p+p collisions is universal to all orders, explicit
2418 calculations [108, 109, 256, 257] have shown that diagrams like those that violate factorization
2419 of the Sivers function make no contribution to the Collins effect at the one- or two-gluon
2420 exchange level, thereby preserving its universality at least to that level.

2421 Comparisons of the transversity distributions extracted from the Collins and IFF channels
2422 will allow STAR to study the size and nature of any factorization breaking effects for TMD
2423 observables in hadronic collisions. Likewise, comparisons with the transversity, Collins and
2424 IFF distributions extracted from SIDIS collisions will shed light on universality and constrain
2425 evolution effects. The measurement of evolution effects in TMD distributions is particularly
2426 important because, unlike the collinear case, TMD evolution contains a non-perturbative
2427 component that cannot be calculated directly. Measurements at \sqrt{s} of 200 and 510 GeV will
2428 provide additional experimental constraints on evolution effects and provide insights into the
2429 size and nature of TMD observables at the future Electron-Ion Collider.

2430 Extending measurements of di-hadron and Collins asymmetries to the forward direction
2431 during Run-22 will allow access to transversity in the region $x > 0.3$. This valence quark
2432 region is not currently probed by any experiments and is essential for the determination of
2433 the tensor charge, which receives 70% of its contributions from $0.1 < x < 1.0$. In addition,
2434 probing transversity in p+p collisions also provides better access to the d -quark transversity
2435 than is available in SIDIS, due to the fact that there is no charge weighting in the hard
2436 scattering QCD $2 \rightarrow 2$ process in p+p collisions. This is a fundamental advantage of p+p
2437 collisions, as any SIDIS measurement of the d -quark transversity has to be on a bound
2438 system, *e.g.* He-3, which ultimately requires nuclear corrections to extract distributions.
2439 The high scale we can reach in 500 GeV collisions at RHIC has allowed STAR [258] to
2440 demonstrate, for the first time, that previous SIDIS measurements at low scales are in fact

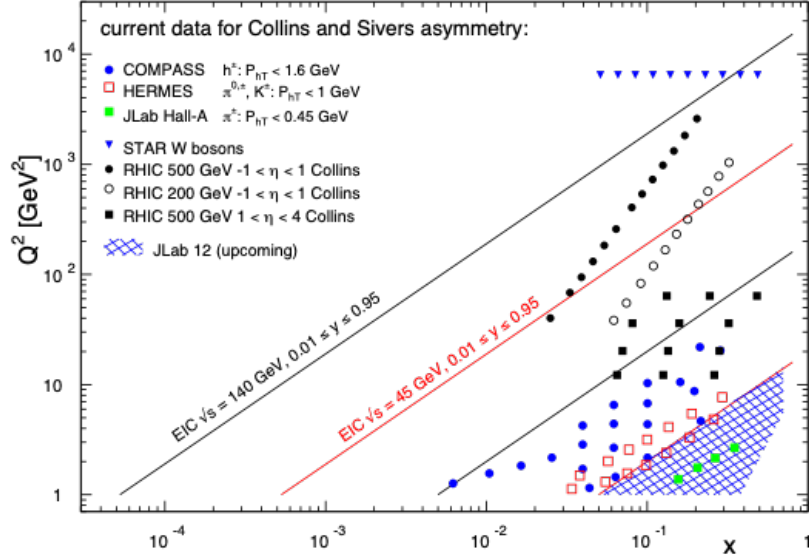


Figure 67: $x - Q^2$ coverage of RHC measurements compared to existing Collins and Sivers effect measurements in SIDIS and the future coverage of the EIC.

2441 accessing the nucleon at leading twist. Figure 67 shows the $x - Q^2$ coverage spanned by
 2442 the RHC measurements compared to the future EIC, JLab-12, and the current SIDIS world
 2443 data.

2444 Another fundamental advantage of p+p collisions is the ability to access gluons di-
 2445 rectly. While gluons cannot be transversely polarized in a transversely polarized spin 1/2
 2446 hadron, they can be linearly polarized. Similarly, there exists an equivalent of the Collins
 2447 fragmentation function for the fragmentation of linearly polarized gluons into unpolarized
 2448 hadrons [259]. The linear polarization of gluons is a largely unexplored phenomenon, but it
 2449 has been a focus of recent theoretical work, in particular due to the relevance of linearly po-
 2450 larized gluons in unpolarized hadrons for the p_T spectrum of the Higgs boson measured at the
 2451 LHC. Polarized proton collisions with $\sqrt{s} = 510$ GeV at RHC, in particular for asymmetric
 2452 parton scattering if jets are detected in the backward direction, are an ideal place to study
 2453 the linearly polarized gluon distribution in polarized protons. (Note that the distributions of
 2454 linearly polarized gluons inside an unpolarized and a polarized proton provide independent
 2455 information). A first measurement of the ‘‘Collins-like’’ effect for linearly polarized gluons
 2456 has been done by STAR with data from Run-11 [258], providing constraints on this function
 2457 for the first time.

2458 Figure 68 shows projected uncertainties for Collins asymmetries at 510 GeV with the
 2459 Forward Upgrade during Run-22. As indicated on the figure, jets with $3 < \eta < 4$ and
 2460 $3 < p_T < 9$ GeV/c will explore transversity in the important region $0.3 < x < 0.5$ that
 2461 has not yet been probed in SIDIS. A realistic momentum smearing of final state hadrons as
 2462 well as jets in this rapidity range was assumed and dilutions due to beam remnants (which
 2463 become substantial at rapidities close to the beam) and underlying event contributions have
 2464 been taken into account. As no dedicated particle identification at forward rapidities will be

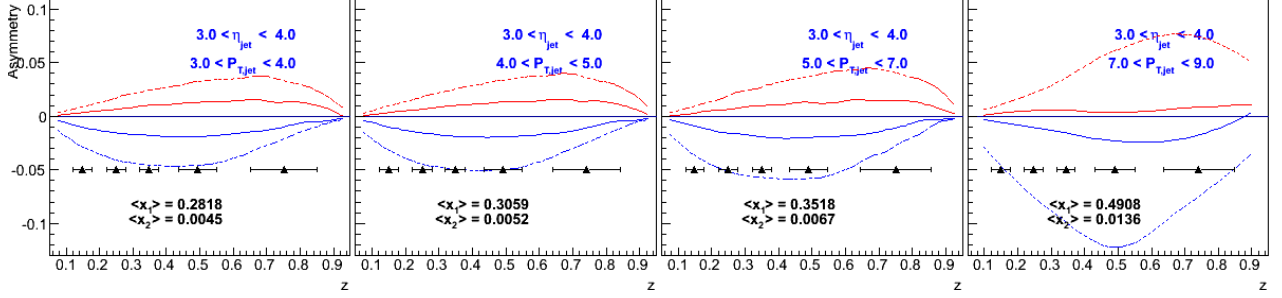


Figure 68: Expected h^- Collins asymmetry uncertainties at $3 < \eta < 4$ (black points) from a sampled luminosity of 268 pb^{-1} at $\sqrt{s} = 510 \text{ GeV}$, compared to positive (red) and negative (blue) pion asymmetries based on the Torino extraction (full lines) and the Soffer bound (dashed lines) as a function of hadron z for bins in jet p_T . Most uncertainties are smaller than the height of the triangles.

2465 available for these measurements, only charged hadrons were considered. This mostly reduces
 2466 the expected asymmetries due to dilution by protons (10-14%) and a moderate amount of
 2467 kaons (12-13%). As anti-protons are suppressed compared to protons in the beam remnants,
 2468 especially the negative hadrons can be considered a good proxy for negative pions ($\sim 78\%$
 2469 purity according to PYTHIA6). Given their sensitivity to the down quark transversity via
 2470 favored fragmentation, they are particularly important since SIDIS measurements, due to
 2471 their electromagnetic interaction, are naturally dominated by up-quarks. We have estimated
 2472 our statistical uncertainties based on an accumulated luminosity of 268 pb^{-1} , which leaves
 2473 nearly invisible uncertainties after smearing. These expected uncertainties are compared to
 2474 the asymmetries obtained from the transversity extractions based on SIDIS and Belle data
 2475 [260] as well as from using the Soffer positivity bound for the transversity PDF [261]. More
 2476 recent global fits have slightly different central up and down quark transversity distributions.
 2477 But due to the lack of any SIDIS data for $x > 0.3$, the upper uncertainties are compatible
 2478 with the Soffer bounds. This high- x coverage will give important insights into the tensor
 2479 charge, which is essential to understand the nucleon structure at leading twist.

2480 Although the studies presented here are for the Collins asymmetries, the resulting sta-
 2481 tistical uncertainties will be similar for other measurements using azimuthal correlations of
 2482 hadrons in jets. One important example is the measurement of ‘‘Collins-like’’ asymmetries to
 2483 access the distribution of linearly polarized gluons. As described earlier, the best kinematic
 2484 region to access this distribution is at backward angles with respect to the polarized proton
 2485 and at small jet p_T . Figure 68 shows that a high precision measurement of the distribution
 2486 of linearly polarized gluons down to $x \sim 0.005$ will be performed concurrently.

2487 It is also important to recognize that these hadron-in-jet measurements with the STAR
 2488 Forward Upgrade will provide very valuable experience detecting jets close to beam rapidity
 2489 that will inform the planning for future jet measurements in similar kinematics at the EIC.

2490 While the STAR Forward Upgrade will provide sensitivity to transversity to the highest
 2491 x , concurrent mid-rapidity measurements (see Fig. 67) will provide the most precise informa-
 2492 tion as a function of x , z , j_T , and Q^2 to probe questions of TMD factorization, universality,

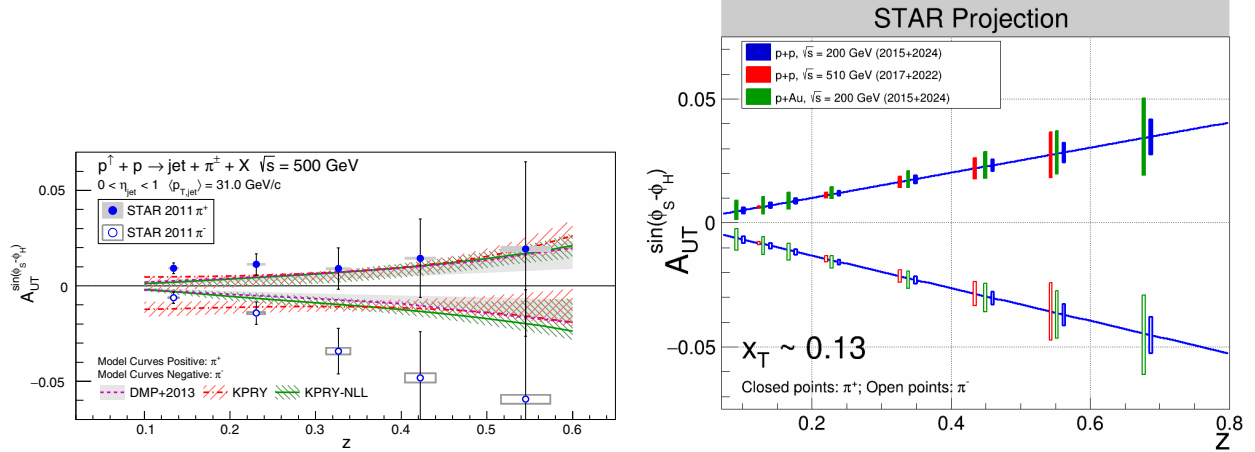


Figure 69: The left panel shows STAR measurements of the Collins asymmetry vs. pion z in 500 GeV p+p collisions from Run-11, compared to several model calculations. See [258] for details. The right panel shows projected statistical uncertainties for STAR Collins asymmetry measurements at $0 < \eta < 0.9$ in p+p at $\sqrt{s} = 200$ and 510 GeV and p+Au at $\sqrt{s_{NN}} = 200$ GeV. The points have arbitrarily been drawn on the solid lines, which represent simple linear fits to the STAR preliminary 200 GeV p+p Collins asymmetry measurements from 2015. (Note that only one bin is shown spanning $0.1 < z < 0.2$ for 510 GeV p+p, whereas three bins are shown covering the same z range for the 200 GeV measurements.)

2493 and evolution. The left panel of Fig. 69 shows published STAR measurements of the Collins
 2494 asymmetry vs. pion z in 500 GeV p+p collisions from 2011 [258]. The results, which repre-
 2495 sented the first ever observation of the Collins effect in p+p collisions, are consistent at the
 2496 2-sigma level with model predictions, with and without TMD evolution, derived from fits to
 2497 e^+e^- and SIDIS data [108,262]. However, greater precision is clearly necessary for a detailed
 2498 universality test, as well as to set the stage for the EIC.

2499 In 2017, STAR sampled about 14 times the luminosity that we recorded in 2011. In Run-
 2500 22, we propose to record another data set equivalent to 16 times the sampled luminosity from
 2501 2011. Furthermore, during Run-22 the iTPC will improve the dE/dx particle identification
 2502 compared to the previous years. Studies using the dE/dx distributions seen in our 200 GeV
 2503 p+p data from 2015 and the actual dE/dx resolution improvements that have been achieved
 2504 during BES-II indicate the iTPC will yield a 20 – 25% increase in the effective figure-of-
 2505 merit for pions with $|\eta| < 0.9$. The right-hand panel of Fig. 69 shows the projected STAR
 2506 statistical uncertainties for the Collins asymmetry at $0 < \eta < 0.9$ in 510 GeV p+p collisions
 2507 once the Run-17 and Run-22 data sets are fully analyzed. It’s also important to recognize
 2508 that the iTPC will also enable STAR to measure the Collins asymmetry over the region
 2509 $0.9 < \eta < 1.3$ during Run-22, in addition to the projections that are shown in Fig. 69.

2510 Statistical improvements from 2011 data [250] to 2017+’22 data comparable to those
 2511 shown for the Collins effect in Fig. 69 are also expected for mid-rapidity measurements of
 2512 transversity in 510 GeV p+p collisions using IFF asymmetries.

2513 3.1.4 Spatial Imaging of the Nucleon

2514 Diffractive and Ultra Peripheral processes at RHIC are an essential tool that can elucidate
2515 the origin of single-spin asymmetries in polarized p+p collisions and access the orbital motion
2516 of partons inside the proton. Also at an EIC diffractive processes have been identified as the
2517 golden tool to study several key physics programs

- 2518 • What is the spatial distribution of quarks and gluons inside the nucleon?
- 2519 • What is the role of orbital motion of sea quarks and gluons in building the nucleon
2520 spin?
- 2521 • saturation in nuclei.

2522 **Diffractive:** The essential characteristics of diffraction in QCD are summarized by two
2523 facts:

- 2524 • The event is still called diffractive if there is a rapidity gap. Due to the presence
2525 of a rapidity gap, the diffractive cross-section can be thought of as arising from an
2526 exchange of several partons with zero net color between the target and the projectile.
2527 In high-energy scattering, which is dominated by gluons, this color neutral exchange
2528 (at the lowest order) consists of at least two exchanged gluons. This color singlet
2529 exchange has historically been called the pomeron, which had a specific interpretation
2530 in Regge theory. A crucial question in diffraction is the nature of the color neutral
2531 exchange between the protons. This interaction probes, in a novel fashion, the nature
2532 of confining interactions within hadrons.
- 2533 • The proton/nuclear target is not always an opaque “black disk” obstacle of geometric
2534 optics. A projectile, which interacts more weakly due to color-screening and asymptotic
2535 freedom, is likely to produce a different diffractive pattern from a larger, more strongly
2536 interacting, projectile.

2537 HERA discovered that 15% of the total ep cross-section is given by diffractive events
2538 (for details see [263] and references therein), basically independent of kinematics. At RHIC
2539 center-of-mass energies diffractive scattering events constitute $\sim 25\%$ of the total inelastic
2540 p+p cross-section [264]. As described above diffraction is defined as an interaction that is
2541 mediated by the exchange of the quantum numbers of the vacuum, as shown in Figure 70.
2542 Experimentally these events can be characterized by the detection of a very forward scattered
2543 proton and jet (singly diffractive) or two jets (doubly diffractive) separated by a large rapidity
2544 gap. Central diffraction, where two protons, separated by a rapidity gap, are reconstructed
2545 along with a jet at mid-rapidity, are also present, but suppressed compared to singly and
2546 doubly diffractive events. To date, there have been no data in p+p collisions studying spin
2547 effects in diffractive events at high \sqrt{s} apart from measuring single spin asymmetries in
2548 elastic p+p scattering [265–268].

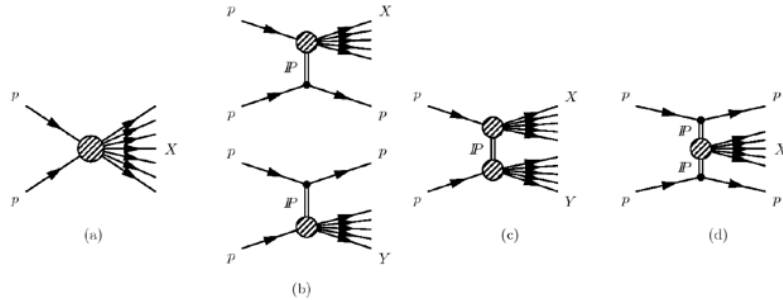


Figure 70: Schematic diagrams of (a) nondiffractive, $pp \rightarrow X$, (b) singly diffractive, $pp \rightarrow Xp$ or $pp \rightarrow pY$, (c) doubly diffractive, $pp \rightarrow XY$, and (d) centrally diffracted, $pp \rightarrow pXp$, events.

2549 A discovery of large transverse single spin asymmetries in diffractive processes would
 2550 open a new avenue to study the properties and understand the nature of the diffractive
 2551 exchange in p+p collisions. One of the primary observables of STAR to access transverse
 2552 spin phenomena has been forward neutral pion production in transversely polarized p+p
 2553 collisions at both $\sqrt{s} = 200$ and 500 GeV. Figure 31 shows the isolated and non-isolated
 2554 transverse single spin asymmetries A_N for π^0 detected in the STAR FMS at $2.5 < \eta < 4.0$
 2555 as a function of x_F .

2556 All these observations might indicate that the underlying subprocess causing a significant
 2557 fraction of the large transverse single spin asymmetries in the forward direction are not of
 2558 $2 \rightarrow 2$ parton scattering processes but of diffractive nature. Figure ?? shows the inclusive π^0
 2559 cross-section for p+p collisions at $\sqrt{s}=200$ GeV versus the leading π^0 energy for three fixed
 2560 forward pseudorapidities [269]. PYTHIA-8 [14] was used to evaluate the fraction of hard
 2561 diffractive events [270] contributing to the inclusive π^0 cross-section at forward rapidities.
 2562 Figure 71 shows the hard diffractive cross-section for π^0 production at $\sqrt{s}=200$ GeV and 500
 2563 GeV for a rapidity range of $2.5 < \eta < 4.0$ with and without applying several experimental
 2564 cuts, i.e. the proton in the STAR Roman Pot acceptance. The prediction from this PYTHIA-
 2565 8 simulation is that 20% of the total inclusive cross-section at forward rapidities is of
 2566 diffractive nature. This result is in agreement with measurements done over a wide range of
 2567 \sqrt{s} (see Figure 12 in Ref. [263]).

2568 In 2015 STAR collected data in $\sqrt{s} = 200$ GeV transversely polarized p+p collisions,
 2569 where an isolated π^0 is detected in the forward pseudorapidity range along with the forward-
 2570 going proton p, which scatters with a near-beam forward pseudorapidity into Roman Pot
 2571 detectors. The sum of the π^0 and the scattered proton energies is consistent with the incident
 2572 proton energy of 100 GeV, indicating that no further particles are produced in this direction.
 2573 Correlations between the π^0 and scattered proton have been presented [271], along with
 2574 single-spin asymmetries which depend on the azimuthal angles of both the pion and the
 2575 proton. This is the first time that spin asymmetries have been explored for this process, and
 2576 a model to explain their azimuthal dependence is needed.

2577 The STAR Forward Upgrade will be a game changer for diffractive measurements at
 2578 RHIC. It will allow the reconstruction of full jets both at $\sqrt{s}=200$ GeV and 500 GeV. As
 2579 at HERA we will be able to reconstruct jets produced with the scattered proton tagged
 2580 in Roman Pots and/or requiring rapidity gaps. Measuring spin asymmetries for diffractive
 2581 events as function of \sqrt{s} might reveal surprises, which will inspire new physics opportunities

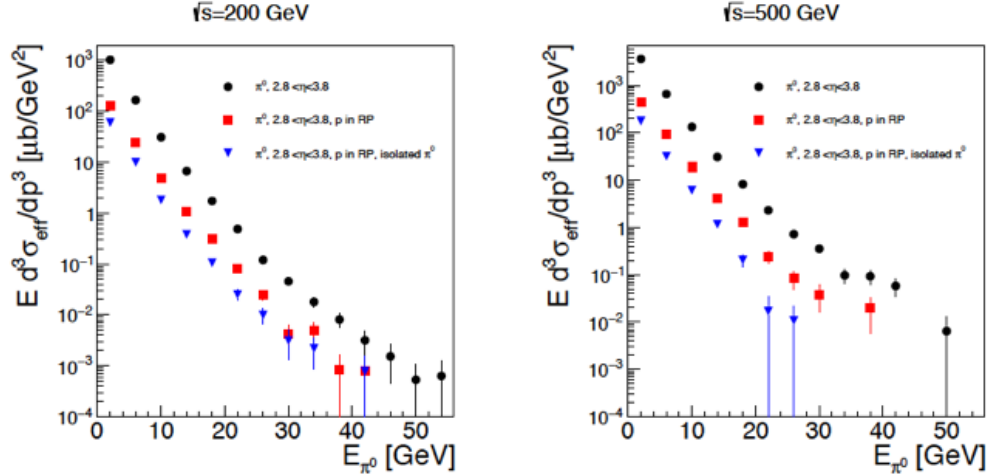


Figure 71: Estimate of the cross-section for hard diffractive processes at $\sqrt{s}=200$ GeV and 500 GeV using Pythia 8. The different points reflect different analysis cuts applied: π^0 in rapidity $2.8 < \eta < 3.8$ (black), one proton is required to be detected in the STAR Roman Pot acceptance (red) and an isolation cut of 35 mrad around the π^0 (blue).

2582 for EIC, i.e SSA in polarized eA collisions.

2583 3.2 Run-24 Request for Polarized p+p and p+A Collisions at 200 2584 GeV

2585 Run-24, with polarized p+p and p+Au collisions at $\sqrt{s_{NN}} = 200$ GeV, will likely be the last
2586 RHIC spin/cold QCD run. This run will provide STAR with the unique opportunity to in-
2587 vestigate these 200 GeV collision systems with the Forward Upgrade providing full tracking
2588 and calorimetry coverage over the region $2.5 < \eta < 4$ and the iTPC providing enhanced
2589 particle identification and expanded pseudorapidity coverage at mid-rapidity. These power-
2590 ful detection capabilities, when combined with substantially increased sampled luminosity
2591 compared to Run-15, will enable critical measurements to probe universality and factoriza-
2592 tion in transverse spin phenomena and nuclear PDFs and fragmentation functions, as well as
2593 low- x non-linear gluon dynamics characteristic of the onset of saturation. This will provide
2594 unique insights into fundamental QCD questions in the near term, and essential baseline
2595 information for precision universality tests when combined with measurements from the EIC
2596 in the future.

2597 We therefore request at least 11 weeks of polarized p+p data-taking at $\sqrt{s} = 200$ GeV
2598 and 11 weeks of polarized p+Au data-taking at $\sqrt{s_{NN}} = 200$ GeV during Run-24. All of
2599 the running will involve transversely polarized protons, with the choice between vertical
2600 or radial polarization to be determined during the coming year. Based on recent (08-21-
2601 20) C-AD guidance, we expect to sample at least 235 pb^{-1} of p+p collisions and 1.3 pb^{-1}
2602 of p+Au collisions. These totals represent 4.5 times the luminosity that STAR sampled
2603 during transversely polarized p+p collisions in Run-15 and 3 times the luminosity that

2604 STAR sampled during transversely polarized p +Au collisions in Run-15.

2605 3.2.1 Spin Physics with Polarized p + p and p +Au Collisions at 200 GeV

2606 Section 1.3 described several very mature STAR analyses that are based on the transversely
2607 polarized p + p and p +Au data sets that we recorded during 2015. Run-24 will enable STAR
2608 to probe these questions with a far more capable detector and much larger data sets than were
2609 available during Run-15, thereby allowing us to set the stage for related future measurements
2610 at the EIC. Here we give brief descriptions of several of the opportunities presented by Run-
2611 24.

2612 **Forward transverse spin asymmetries:** Section 1.3 presents a small subset of the re-
2613 sults that STAR will publish very soon in a pair of papers discussing forward transverse spin
2614 asymmetries in p + p , p +Al, and p +Au collisions measured with the Forward Meson Spec-
2615 trometer (FMS). One paper focuses on the dynamics that underlie the large asymmetries
2616 that have been seen to date. Figure 31 shows that A_N for forward π^0 production in p + p
2617 collisions at 200 and 500 GeV is substantially larger when the π^0 is isolated than when it
2618 is accompanied by additional nearby photons. The same analysis also shows that A_N for
2619 inclusive electromagnetic jets (EM-jets) in 200 and 500 GeV collisions is substantially larger
2620 than that for EM-jets that contain three or more photons and that the Collins asymmetry
2621 for π^0 in EM-jets is very small. The other paper focuses on the nuclear dependence of A_N
2622 for π^0 in $\sqrt{s_{NN}} = 200$ GeV collisions. It presents a detailed mapping of A_N as functions of
2623 x_F and p_T for all three collision systems. Figure 30 shows the observed nuclear dependence
2624 is very weak. The same analysis shows that isolated *vs.* non-isolated π^0 behave similarly in
2625 p +Al and p +Au collisions as they do in p + p collisions.

2626 These two papers will provide a wealth of new data to inform the ongoing discussion
2627 regarding the origin of the large inclusive hadron transverse spin asymmetries that have
2628 been seen in p + p collisions at forward rapidity over a very broad range of collision energies.
2629 Nonetheless, the STAR Forward Upgrade will be a game changer for such investigations. It
2630 will enable measurements of A_N for $h^{+/-}$, in addition to π^0 . It will enable isolation criteria to
2631 be applied to the $h^{+/-}$ and π^0 that account for nearby charged, as well as neutral, fragments.
2632 It will enable full jet asymmetry and Collins effect measurements, again for $h^{+/-}$ in addition
2633 to π^0 , rather than just EM-jet measurements. It will permit all of these measurements to
2634 be performed at both 510 GeV, as discussed in Sects. 3.1.1 and 3.1.2, and at 200 GeV. And
2635 all of these observables can be tagged by forward protons detected in the STAR Roman
2636 pots to identify the diffractive component of the observed transverse spin asymmetries. For
2637 p + p there will be considerable overlap between the kinematics at the two energies, but the
2638 510 GeV measurements will access higher p_T , while the 200 GeV measurements will access
2639 higher x_F . Meanwhile, at 200 GeV we will also perform the full suite of measurements in
2640 p +Au to identify any nuclear effects. Figure 64 shows one set of predictions for the inclusive
2641 $\pi^{+/-}$ A_N in 200 and 500 GeV p + p collisions, while Fig. 65 shows the estimated sensitivity
2642 of one hadron-in-jet measurement that will help to isolate the Sivers effect contribution at
2643 200 GeV.

2644 **Sivers effect:** Sections 1.3 and 3.1.2 describe the first ever observation of the Sivers effect in
 2645 dijet production. Such measurements are crucial to explore questions regarding factorization
 2646 of the Sivers function in dijet hadroproduction [252–255]. Those results were derived from
 2647 200 GeV transverse spin data that STAR recorded in 2012 and 2015 (total sampled luminosity
 2648 $\sim 75 \text{ pb}^{-1}$ for the two years combined). Nonetheless, the uncertainties remain large, as can be
 2649 seen in Fig. 29. Run-24 data will reduce the uncertainties for $|\eta_3 + \eta_4| < 1$ by a factor of two.
 2650 The increased acceptance from the iTPC will reduce the uncertainties at $|\eta_3 + \eta_4| \approx 2.5$ by a
 2651 much larger factor, while the Forward Upgrade will enable the measurements to be extended
 2652 to even larger values of $|\eta_3 + \eta_4|$. When combined with the 510 GeV data from 2017 and
 2653 2022 (see Sect. 3.1.2), the results will provide a detailed mapping *vs. x* for comparison to
 2654 results to Sivers functions extracted from SIDIS, Drell-Yan, and vector boson production.

2655 **Transversity and related quantities:** As described in Sect. 3.1.3, measurements of the
 2656 Collins asymmetry and IFF in p+p collisions at RHIC probe fundamental questions regarding
 2657 TMD factorization, universality, and evolution. Data from 200 GeV p+p collisions will play
 2658 an essential role toward answering these questions. Figure 67 shows that 200 GeV p+p
 2659 collisions interpolate between the coverage that we will achieve during Run-22 at high-*x*
 2660 with the Forward Upgrade and at low-*x* with the STAR mid-rapidity detectors. They will
 2661 also provide a significant overlapping region of *x* coverage, but at Q^2 values that differ by a
 2662 factor of 6. This will provide valuable information about evolution effects, as well as cross-
 2663 checks between the two measurements. Furthermore, for most of the overlapping *x* region,
 2664 200 GeV p+p collisions will also provide the greatest statistical precision (see for example
 2665 Fig. 69), thereby establishing the most precise benchmark for future comparisons to *ep* data
 2666 from the EIC.

2667 The high statistical precision of the Run-24 data will enable detailed multi-dimensional
 2668 binning for the Collins asymmetry results. This is particularly valuable because, as empha-
 2669 sized in [108, 109], hadron-in-jet measurements in p+p collisions provide a direct probe of
 2670 the Collins fragmentation function because they combine it with the *collinear* transversity
 2671 distribution. In general, the observed asymmetries are functions of jet (p_T, η) , hadron (z, j_T) ,
 2672 and Q^2 . However, the physics interpretations associated with these variables separate, with
 2673 p_T and η primarily coupling to the incident quark *x* and the polarization transfer in the
 2674 hard scattering, while z and j_T characterize the fragmentation kinematics. Thus, A_{UT} *vs.*
 2675 p_T , as shown in Fig. 28 for the preliminary 2015 analysis, provides information about the
 2676 transversity distribution. In parallel, the (z, j_T) dependence, integrated over a wide range of
 2677 jet p_T , as shown in Fig. 72 for the preliminary 2015 results, provides a detailed look at the
 2678 Collins fragmentation function. Note that STAR finds the maximum value of A_{UT} shifts to
 2679 higher j_T as z increases. The statistical uncertainties in Figs. 28 and 72 will be reduced by
 2680 a factor of 2.5 when Run-15 and Run-24 data are combined together.

2681 The 2015 Collins analysis has also, for the first time, measured the Collins effect for
 2682 charged kaons in p+p collisions, as shown in Fig. 73. The asymmetries for K^+ , which
 2683 like π^+ have a contribution from favored fragmentation of *u* quarks, are about 1.5-sigma
 2684 larger than those for π^+ in Fig. 28, while those for K^- , which can only come from unfavored

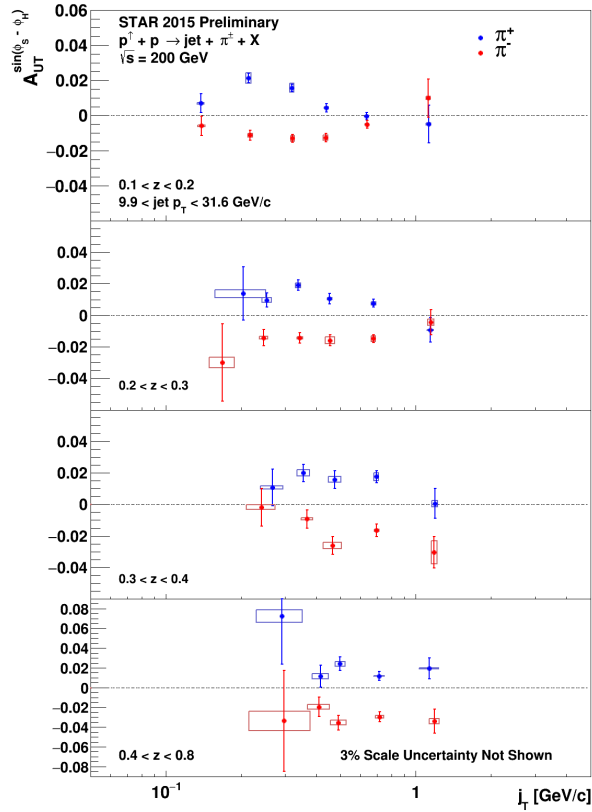


Figure 72: Preliminary 2015 results for the Collins asymmetry for charged pions in 200 GeV p+p collisions as a function of z and j_T , integrated over $9.9 < p_T < 31.6$ GeV/ c and $0 < \eta < 0.9$.

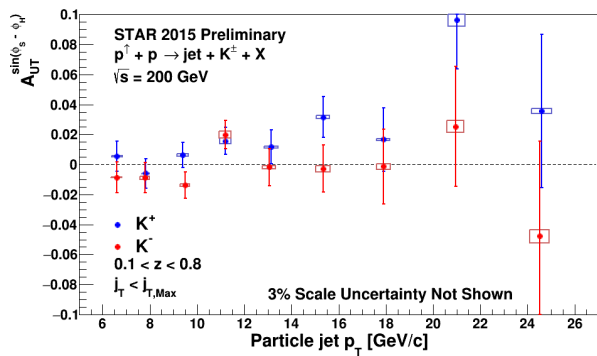


Figure 73: Preliminary 2015 results for the $K^{+/-}$ Collins asymmetries *vs.* jet p_T for $0 < \eta < 0.9$ in 200 GeV p+p collisions.

2685 fragmentation, are consistent with zero at the 1-sigma level. These trends are similar to those
 2686 found in SIDIS by HERMES [272] and COMPASS [273], and provide additional insight into
 2687 the Collins fragmentation function. This same analysis with Run-24 data will yield statistical
 2688 uncertainties a factor of 3 smaller than those in Fig. 73. This is a much greater improvement
 2689 than would be expected from the increase in sampled luminosity thanks to the improved
 2690 dE/dx resolution provided by the iTPC. In addition, the iTPC will enable the measurements
 2691 in Figs. 28, 72, and 73 to be extended to an additional higher η bin ($0.9 < \eta < 1.3$).

2692 RHIC has the unique opportunity to extend the Collins effect measurements to nuclei.
 2693 This will provide an alternative look at the universality of the Collins effect in hadroproduc-
 2694 tion by dramatically increasing the color flow options of the sort that have been predicted
 2695 to break factorization for TMD PDFs like the Sivers effect [252, 253]. This will also explore
 2696 the spin dependence of the hadronization process in cold nuclear matter. STAR collected a
 2697 proof-of-principle data set during the 2015 p +Au run that is currently under analysis. Those
 2698 data will provide a first estimate of medium-induced effects. However, the small nuclear ef-
 2699 fects seen by STAR for forward inclusive $\pi^0 A_N$ (see Fig. 30) indicate that greater precision
 2700 will likely be needed. Figure 69 shows the projected 2015+'24 statistical uncertainties for
 2701 the p +Au Collins asymmetry measurement at $\sqrt{s_{NN}} = 200$ GeV, compared to those for the
 2702 p + p at the same energy.

2703 **Ultra-peripheral collisions:** The formalism of generalized parton distributions (GPDs)
 2704 provides a theoretical framework which addresses some of the above questions [274–277].
 2705 Constraints on GPDs have mainly been provided by exclusive reactions in DIS, e.g. deeply
 2706 virtual Compton scattering. RHIC, with its unique capability to collide transversely polar-
 2707 ized protons at high energies, has the opportunity to measure A_N for exclusive J/Ψ produc-
 2708 tion in ultra-peripheral collisions (UPCs) [278]. In such a UPC process, a photon emitted by
 2709 the opposing beam particle (p or A) collides with the polarized proton. The measurement is
 2710 at a fixed $Q^2 \sim M_{J/\psi}^2 \approx 10$ GeV² and $10^{-4} < x < 10^{-1}$. A nonzero asymmetry would be the
 2711 first signature of a nonzero GPD E_g for gluons, which is sensitive to spin-orbit correlations
 2712 and is intimately connected with the orbital angular momentum carried by partons in the
 2713 nucleon and thus with the proton spin puzzle.

2714 The 2015 p^\uparrow +Au data allowed a proof-of-principle of such a measurement. A trigger
 2715 requiring back-to-back energy deposits in the Barrel Electromagnetic Calorimeter selected
 2716 J/Ψ candidates. The e^+e^- mass distribution after selection cuts is shown in the left of
 2717 Fig. 74, and the pair p_T distribution of the J/ψ mass peak is shown on the right of the
 2718 figure. The data are well described by the STARlight model [279] (colored histograms in
 2719 the figure), including the dominant $\gamma+p^\uparrow \rightarrow J/\psi$ signal process and the $\gamma+Au \rightarrow J/\psi$ and
 2720 $\gamma+\gamma \rightarrow e^+e^-$ background processes. The left of Fig. 75 shows the transverse asymmetry A_N^γ
 2721 for the signal J/ψ , which have a mean photon-proton center-of-mass energy $W_{\gamma p} \approx 24$ GeV.
 2722 The result is consistent with zero. Also shown is a prediction based on a parameterization
 2723 of E_g [280]; the present data provide no discrimination of this prediction.

2724 This measurement can be greatly improved with a high statistics transversely polarized
 2725 p^\uparrow +Au run in 2024. The integrated luminosity for the 2015 measurement was 140 nb^{-1} ; the

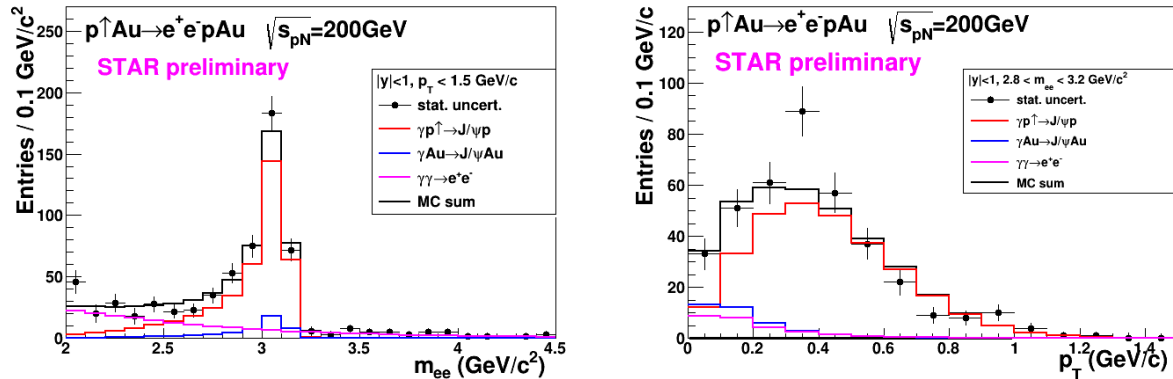


Figure 74: Mass distribution of selected e^+e^- pairs (left), and p_T distribution of the J/ψ mass peak (right). The colored histograms are the indicated processes modelled by STARlight and the sum fit to the data.

2726 2024 run will provide 1.3 pb^{-1} , allowing a sizeable reduction of statistical uncertainty in the
 2727 same $W_{\gamma p}$ range. However, the Forward Upgrade and iTPC will also provide a significant
 2728 extension of the $W_{\gamma p}$ range of the measurement. The right of Fig. 75 shows the accepted
 2729 cross section for $\gamma + p^\dagger \rightarrow J/\psi$ for various detector pseudorapidity ranges. With the full
 2730 detector, the sensitive cross section is a factor of five times the central barrel alone and the
 2731 expected asymmetry is substantially larger. The statistical uncertainty on A_N^γ as shown in
 2732 the left of Fig. 75 will be ≈ 0.02 , offering a powerful test of a non-vanishing E_g . Also, the
 2733 accepted region has a lower mean $W_{\gamma p} \approx 14 \text{ GeV}$. Predictions based on E_g parameterizations
 2734 such as shown in the figure have a larger asymmetry at lower $W_{\gamma p}$, with increased possibility
 2735 of a nonzero result. Alternatively, the increased statistics will allow a measurement of A_N^γ
 2736 in bins of $W_{\gamma p}$.

2737 Similar measurements are also possible with future $p^\dagger + p^\dagger$ runs at $\sqrt{s} = 200$ and 510
 2738 GeV. However, the UPC cross section scales with Z^2 of the the nucleus emitting the photon;
 2739 for protons this is $1/79^2$ relative to Au nuclei, which makes analogous measurements in p+p
 2740 collisions extremely luminosity hungry.

2741 3.2.2 Physics Opportunities with Unpolarized proton-Nucleus Collisions

2742 Our quest to understand QCD processes in Cold Nuclear Matter (CNM) centers on the
 2743 following fundamental questions:

- 2744 • Can we experimentally find evidence of a novel universal regime of non-linear QCD
 2745 dynamics in nuclei?
- 2746 • What is the role of saturated strong gluon fields, and what are the degrees of freedom
 2747 in this high gluon density regime?
- 2748 • What is the fundamental quark-gluon structure of light and heavy nuclei?

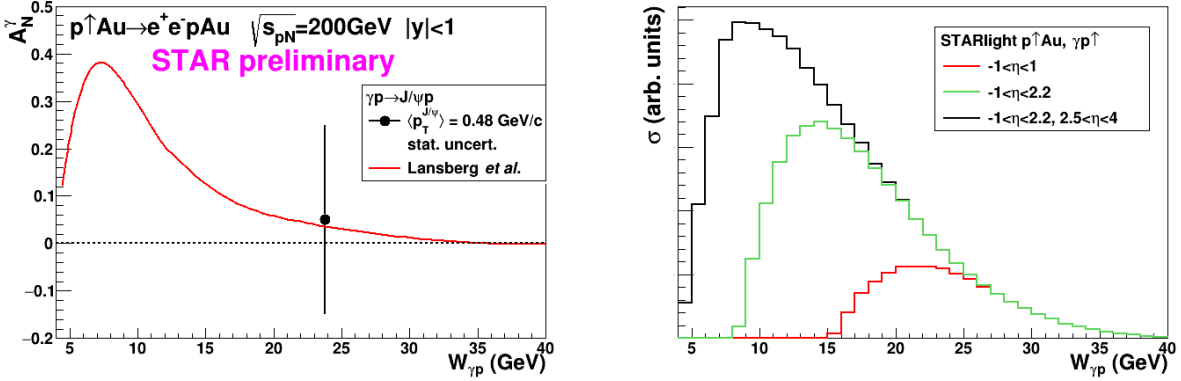


Figure 75: Left: The measured J/ψ transverse asymmetry A_N^γ and a prediction based on a parameterization of E_g . Right: The accepted cross section for $\gamma+p\uparrow\rightarrow J/\psi$ for various detector pseudorapidity η ranges; the black curve shows the result for the full STAR detector with the Forward Upgrade and the iTPC.

- 2749 • Can a nucleus, serving as a color filter, provide novel insight into the propagation,
 2750 attenuation and hadronization of colored quarks and gluons?

2751 Various aspects of these questions have been addressed by numerous experiments and
 2752 facilities around the world, most of them at significantly lower center-of-mass energies and
 2753 kinematic reach than RHIC. Deep inelastic scattering on nuclei addresses some of these
 2754 questions with results from, for instance, HERMES at DESY [281–283], CLAS at JLab [284],
 2755 and in the future at the JLab 12 GeV. This program is complemented by hadron-nucleus
 2756 reactions in fixed target p+A at Fermilab (E772, E886, and E906) [285] and at the CERN-
 2757 SPS.

2758 In the following we propose a measurement program unique to RHIC to constrain the
 2759 initial state effects in strong interactions in the nuclear environment. We also highlight the
 2760 complementarity to the LHC p+Pb program and stress why RHIC data are essential and
 2761 unique in the quest to further our understanding of nuclei. The uniqueness of the RHIC
 2762 program is based on the flexibility of the RHIC accelerator to run collisions of different
 2763 particle species at very different center-of-mass energies. This in combination with the
 2764 enhanced STAR detector capabilities in Run-24 allows to disentangle nuclear effects in the
 2765 initial and final state as well as leading twist shadowing from saturation effects in a kinematic
 2766 regime where all these effects are predicted to be large. Most of the discussed measurements
 2767 critically rely on the Forward Upgrade.

2768 **The initial state of nuclear collisions:**

2769 **Nuclear parton distribution functions:** A main emphasis of the 2015 and later
 2770 p+A runs is to determine the initial conditions of the heavy ion nucleus before the collision
 2771 to support the theoretical understanding of the A+A program both at RHIC and the LHC.

2772 In the following, the current status of nPDFs will be discussed, including where the unique
 2773 contribution of RHIC lie, in comparison to the LHC and the future EIC.

2774 Our current understanding of nuclear parton distribution functions (nPDFs) is still very
 2775 limited, in particular, when compared with the rather precise knowledge of PDFs for free
 2776 protons collected over the past 30 years. Figure 76 shows an extraction of nPDFs from
 2777 available data, along with estimates of uncertainties. All results are shown in terms of
 2778 the nuclear modification ratios, i.e., scaled by the respective PDF of the free proton. The
 2779 yellow bands indicate regions in x where the fits are not constrained by data [286] and
 2780 merely reflect the freedom in the functional form *assumed* in the different fits. Clearly, high
 2781 precision data at small x and for various different values of Q^2 are urgently needed better to
 2782 constrain the magnitude of suppression in the x region where non-linear effects in the scale
 2783 evolution are expected. In addition, such data are needed for several different nuclei, as
 2784 the A-dependence of nPDFs cannot be predicted from first principles in pQCD and, again,
 2785 currently relies on assumptions. Note that the difference between DSSZ [287] and EPS09
 2786 for the gluon modification arise from the different treatment of the PHENIX midrapidity
 2787 $\pi^0 R_{dAu}$ data [288], which in the EPS09 [289] fit are included with an extra weight of 20. The
 2788 $\pi^0 R_{dAu}$ data are the only data, which can probe the gluon in the nucleus directly, but these
 2789 data also suffer from unknown nuclear effects in the final state (see [290]). Therefore, it is
 2790 absolutely critical to have high precision data only sensitive to nuclear modification in the
 2791 initial state over a wide range in x and intermediate values of Q^2 (away from the saturation
 2792 regime) to establish the nuclear modification of gluons in this kinematic range.

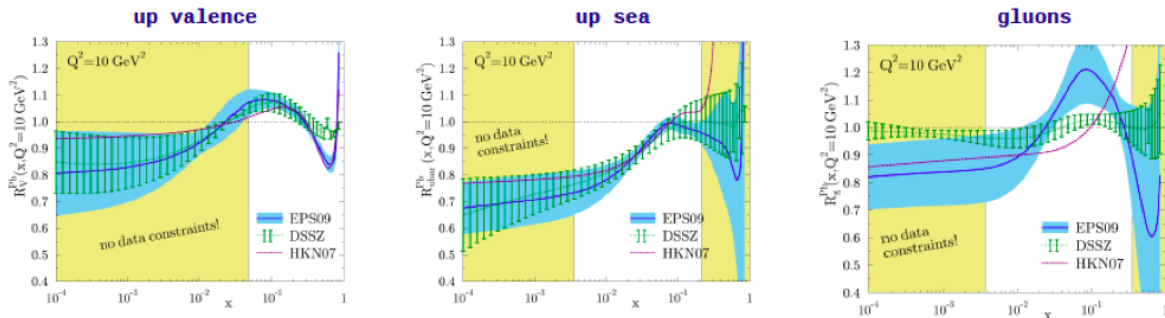


Figure 76: Summary of the most recent sets of nPDFs. The central values and their uncertainty estimates are given for the up valence quark, up sea quark, and the gluon. The yellow bands indicate regions in x where the fits are not constrained by any data (taken from Ref. [286]).

2793 It is important to realize that the measurements from RHIC are compelling and essential
 2794 even when compared to what can be achieved in p+Pb collisions at the LHC. Due to the
 2795 higher center-of-mass system energy most of the LHC data have very high Q^2 , where the
 2796 nuclear effects are already reduced significantly by evolution and are therefore very difficult to
 2797 constrain. Two recent articles [291,292] assessed the impact of the available LHC Run-I p+Pb
 2798 data on determinations of nPDFs. The rather moderate impact of these data is illustrated
 2799 in Figure 77. Note that the extra weight factor of 20 for the PHENIX midrapidity $\pi^0 R_{dAu}$
 2800 data [288] in the original EPS09 [289] fit was removed in all of the new fits, leading to a

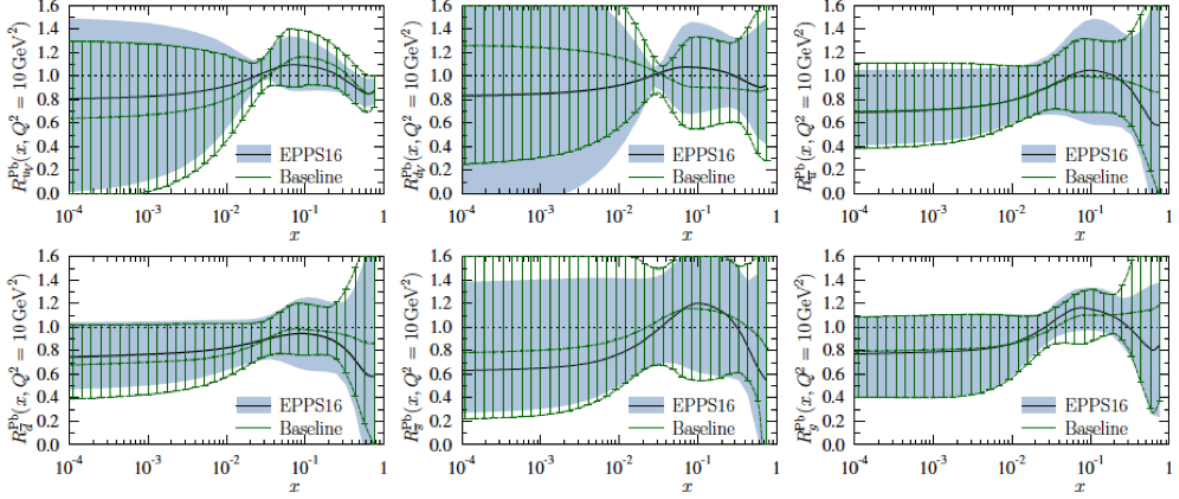


Figure 77: The nuclear modifications at $Q^2=10\text{GeV}^2$ from the EPPS-16 fit (black central line and light-blue bands) compared with the Baseline fit (green curves with hatching) which uses only the data included in the EPS09 fit.

2801 much smaller nuclear modification factor for gluons, especially at medium to high x .

2802 RHIC has the *unique* capability to provide data in a kinematic regime (moderate Q^2 and
 2803 medium-to-low x) where the nuclear modification of the sea quark and the gluon is expected
 2804 to be sizable and currently completely unconstrained. In addition, and unlike the LHC,
 2805 RHIC has the potential to vary the nucleus in p+A collisions and as such also constrain the
 2806 A -dependence of nPDFs.

2807 Extraction of this information is less ambiguous if one uses processes in which strong
 2808 (QCD) final-state interactions can be neglected or reduced. Such golden channels would
 2809 include: a measurement of R_{pA} for Drell-Yan production at forward pseudo-rapidities with
 2810 respect to the proton direction ($2.5 < \eta < 4$.) to constrain the nuclear modifications of sea-
 2811 quarks; and of R_{pA} for direct photon production in the same kinematic regime to constrain
 2812 the nuclear gluon distribution. The first measurement of R_{pA} for direct photon production
 2813 has been done already during the p+Au and p+Al runs in 2015, with recorded luminosities
 2814 by STAR of $L_{pAu} = 0.45 \text{ pb}^{-1}$ and $L_{pAl} = 1 \text{ pb}^{-1}$, respectively. The anticipated statistical
 2815 precision for pA runs in 2015 and projections for the run in 2024 are shown in Fig. 78. The
 2816 Forward Upgrade with its tracking at forward rapidities will also provide the possibility to
 2817 measure R_{pA} for positive and negatively charged hadrons.

2818 Figure 79(left) shows the significant impact of the Run-2015 and 2024 R_{pA} for direct
 2819 photon production on the corresponding theoretical expectations and their uncertainties
 2820 obtained with the EPPS-16 set of nPDFs. The uncertainty bands are obtained through a
 2821 reweighting procedure [293] by using the projected data shown in Fig. 78 and randomizing
 2822 them according to their expected statistical uncertainties around the central values obtained
 2823 with the current set of EPPS-16 nPDFs. Figure 79(right) shows how these measurements
 2824 will help significantly in further constraining the nuclear gluon distribution in a broad range

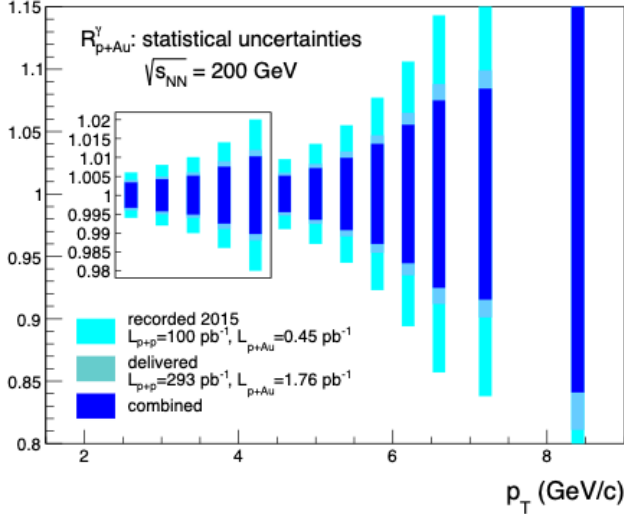


Figure 78: Projected statistical uncertainties for R_{pA} for direct photons in Run-2015 (light blue) and a run in 2024 (blue) and the sum of both (dark blue). The recorded luminosity for Run-2015 was $L_{pAu} = 450 \text{ nb}^{-1}$ and $L_{pp} = 100 \text{ pb}^{-1}$. The delivered luminosity for Run-2024 is assumed to be $L_{pAu} = 1.8 \text{ pb}^{-1}$ and $L_{pp} = 300 \text{ pb}^{-1}$.

2825 of x that is roughly correlated with accessible transverse momenta of the photon, i.e., few
 2826 times $10^{-3} < x < \text{few times } 10^{-2}$. The relevant scale Q^2 is set be $\sim p_T^2$ and ranges from 6
 2827 GeV^2 to about 40 GeV^2 . Like all other inclusive probes in p+p and pA collisions, e.g., jets,
 2828 no access to the exact parton kinematics can be provided event-by-event but global QCD
 2829 analyses easily account for that. After the p+Au run in 2024, the statistical precision of the
 2830 prompt photon data will be sufficient to contribute to a stringent test of the universality
 2831 of nuclear PDFs when combined with the expected data from the EIC (see Figure 2.22 and
 2832 2.23 in Ref [294]).

2833 Figure 80 shows the kinematic coverage in x - Q^2 of past, present, and future experiments
 2834 capable of constraining nuclear parton distribution functions. The experiments shown pro-
 2835 vide measurements that access the initial state parton kinematics on an event-by-event basis
 2836 (in a leading order approximation) while remaining insensitive to any nuclear effects in the
 2837 final state. Some of the LHC experiments cover the same x -range as DY at forward pseudo-
 2838 rapidities at RHIC but at a much higher scale Q^2 , where nuclear modifications are already
 2839 significantly reduced [292, 295, 296]. At intermediate Q^2 , DY at RHIC will extend the low- x
 2840 reach by nearly one decade compared to EIC.

2841 The biggest challenge of a DY measurement is to suppress the overwhelming hadronic
 2842 background: the total DY cross-section is about 10^{-5} to 10^{-6} smaller than the corresponding
 2843 hadron production cross-sections. Therefore, the probability of misidentifying a hadron
 2844 track as a lepton has to be suppressed to the order of 0.1% while maintaining reasonable
 2845 electron detection efficiencies. To that end, we have studied the combined electron/hadron
 2846 discriminating power of the Forward Upgrade. It was found that by applying multivariate
 2847 analysis techniques to the features of EM/hadronic shower development and momentum
 2848 measurements we can achieve hadron rejection powers of 200 to 2000 for hadrons of 15 GeV
 2849 to 50 GeV with 80% electron detection efficiency.

2850 The same procedure as for the direct photon R_{pA} was used to study the potential impact
 2851 of the DY R_{pA} data for the EPPS-19 sets of nPDFs. We expect again a significant impact

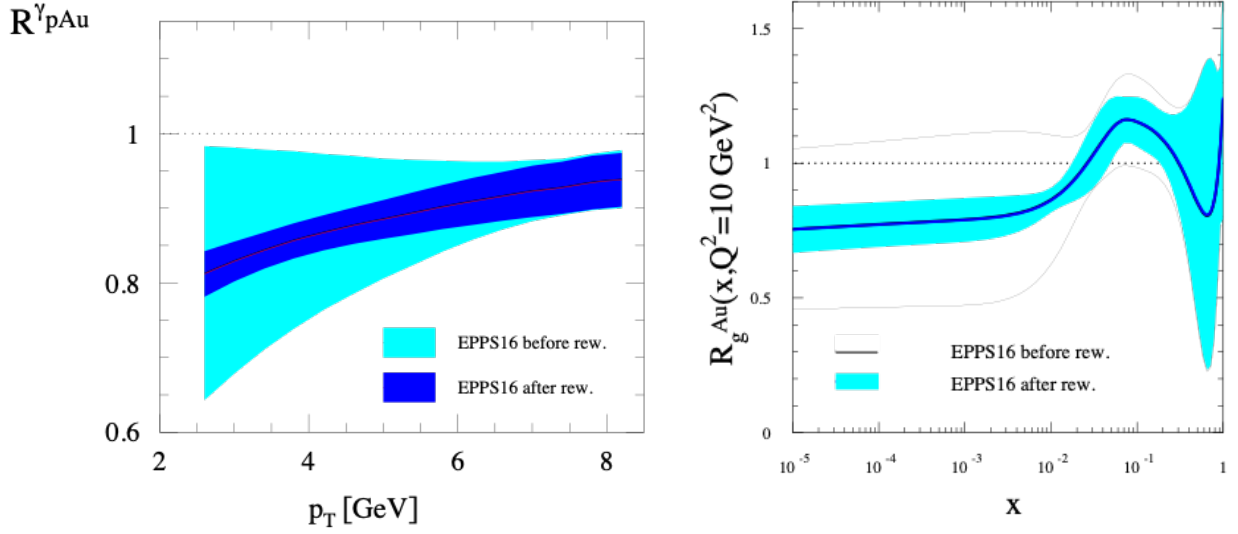


Figure 79: (left) The impact of the direct photon R_{pA} data measured in Run-2015 (blue band) and for the anticipated statistics for the future $p+Au$ run in 2024 (dark blue band) compared with the current uncertainties (cyan band) from EPPS-16. (right) The impact of the direct photon R_{pA} data measured in Run-2015 and for the anticipated statistics for the future $p+Au$ run in 2024 on EPPS-16. The impact is shown on the nuclear suppression factor R_g of nPDF to the proton PDF, the grey bands represent the uncertainties before including the RHIC pseudo data.

2852 on the uncertainties of R_{pA} DY upon including the projected and properly randomized data.
 2853 Clearly, the DY data from RHIC will be instrumental in reducing present uncertainties in
 2854 nuclear modifications of sea quarks. Again, these data will prove to be essential in testing the
 2855 fundamental universality property of nPDFs in the future when EIC data become available.

2856 STAR's unique detector capabilities will provide the first data on J/Ψ -production in
 2857 ultra-peripheral collisions. This measurement provides access to the spatial gluon distri-
 2858 bution by measuring the t -dependence of $d\sigma/dt$. As follows from the optical analogy, the
 2859 Fourier-transform of the square root of this distribution yields the source distribution of the
 2860 object probed. To study the gluon distribution in the gold nucleus, events need to be tagged
 2861 where the photon is emitted from the proton. For both observables a measurement with
 2862 different nuclei is required to pin down the A -dependence of nPDFs. The J/Ψ -production
 2863 in ultra-peripheral collisions requires significantly more statistics than accumulated to date.

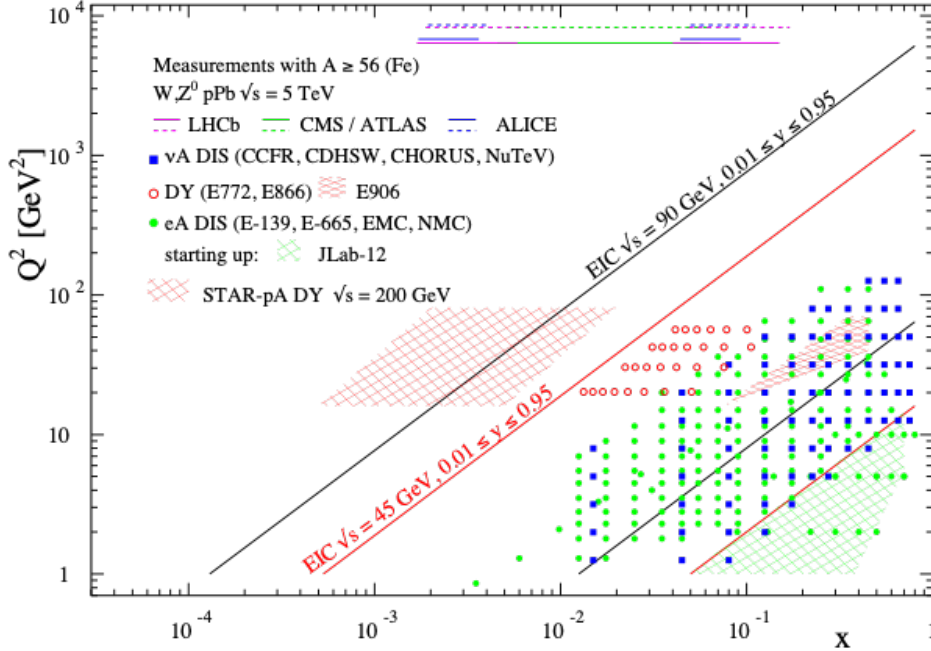


Figure 80: The kinematic coverage in $x \sqrt{Q^2}$ of past, present and future experiments constraining nPDFs with access to the exact parton kinematics event-by-event and no fragmentation in the final state.

2864 **Gluon saturation:** Our understanding of the proton structure and of the nuclear
 2865 interactions at high energy would be advanced significantly with the definitive discovery
 2866 of the saturation regime [297–303]. Saturation physics would provide an infrared cutoff for
 2867 perturbative calculations, the saturation scale Q_s , which grows with the atomic number of the
 2868 nucleus A and with decreasing value of x . If Q_s is large it makes the strong coupling constant
 2869 small, $\alpha_s(Q_s^2) \ll 1$ allowing for perturbative QCD calculations to be under theoretical
 2870 control.

2871 It is well known that PDFs grow at small- x . If one imagines how such a high number of
 2872 small- x partons would fit in the (almost) unchanged proton radius, one arrives at the picture
 2873 presented in Figure 81: the gluons and quarks are packed very tightly in the transverse plane.
 2874 The typical distance between the partons decreases as the number of partons increases, and
 2875 can get small at low- x (or for a large nucleus instead of the proton). One can define the
 2876 saturation scale as the inverse of this typical transverse inter-parton distance. Hence Q_s
 2877 indeed grows with A and decreasing x .

2878 The actual calculations in saturation physics start with the classical gluon fields (as gluons
 2879 dominate quarks at small- x) [304–310], which are then evolved using the nonlinear small- x
 2880 BK/JIMWLK evolution equations [311, 312, 312–320]. The saturation region can be well-
 2881 approximated by the following formula: $Q_s^2 \sim (A/x)^{1/3}$. Note again that at small enough
 2882 x the saturation scale provides an IR cutoff, justifying the use of perturbative calculations.
 2883 This is important beyond saturation physics, and may help us better understand small- x
 2884 evolution of the TMDs.

2885 While the evidence in favor of saturation physics has been gleaned from the data col-
 2886 lected at HERA, RHIC and the LHC, the case for saturation is not sealed and alternative
 2887 explanations of these data exist. The EIC is slated to provide more definitive evidence for

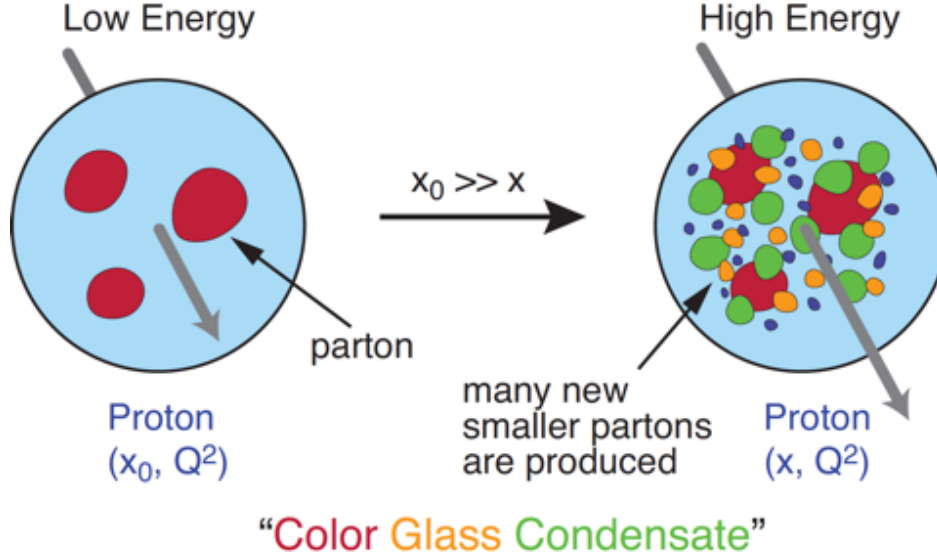


Figure 81: Proton wave function evolution towards small- x

2888 saturation physics [321]. To help the EIC complete the case for saturation, it is mandatory to
 2889 generate higher-precision measurements in p+A collisions at RHIC. These higher-precision
 2890 measurements would significantly enhance the discovery potential of the EIC as they would
 2891 enable a stringent test of universality of the CGC. We stress again that a lot of theoretical
 2892 predictions and results in the earlier Sections of this document would greatly benefit from
 2893 saturation physics: the small- x evolution of TMDs in a longitudinally or transversely polar-
 2894 ized proton, or in an unpolarized proton, can all be derived in the saturation framework [322]
 2895 in a theoretically better-controlled way due to the presence of Q_s . Hence saturation physics
 2896 may help us understand both the quark and gluon helicity PDFs as well as the Siverts and
 2897 Boer-Mulders functions.

2898 The saturation momentum is predicted to grow approximately like a power of energy,
 2899 $Q_s^2 \sim E^{\lambda/2}$ with $\lambda \sim 0.2-0.3$, as phase space for small- x (quantum) evolution opens up. The
 2900 saturation scale is also expected to grow in proportion to the valence charge density at the
 2901 onset of small- x quantum evolution. Hence, the saturation scale of a large nucleus should
 2902 exceed that of a nucleon by a factor of $A^{1/3} \sim 5$ (on average over impact parameters). RHIC
 2903 is capable of running p+A collisions for different nuclei to check this dependence on the mass
 2904 number. This avoids potential issues with dividing say p+Pb collisions in N_{part} classes [323].
 2905 Figure 82 shows the kinematic coverage in the x - Q^2 plane for p+A collisions at RHIC, along
 2906 with previous e+A measurements and the kinematic reach of an EIC. The saturation scale for
 2907 a Au nucleus and the proton is also shown. To access at RHIC a kinematic regime sensitive
 2908 to saturation with $Q^2 > 1 \text{ GeV}^2$ requires measurements at forward rapidities. For these
 2909 kinematics the saturation scale is moderate, on the order of a few GeV^2 , so measurements
 2910 sensitive to the saturation scale are by necessity limited to semi-hard processes.

2911 Until today the golden channel at RHIC to observe strong hints of saturation has been
 2912 the angular dependence of two-particle correlations, because it is an essential tool for testing

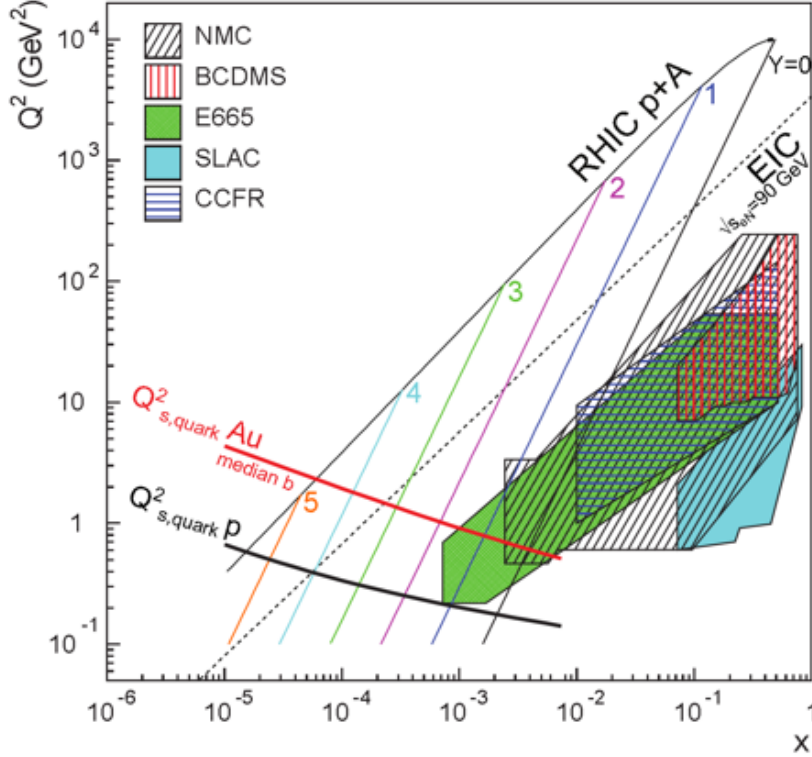


Figure 82: Kinematic coverage in the $x - Q^2$ plane for p+A collisions at RHIC, along with previous e+A measurements, the kinematic reach of an electron-ion collider, and estimates for the saturation scale Q_s in Au nuclei and protons. Lines are illustrative of the range in x and Q^2 covered with hadrons at various rapidities.

2913 the underlying QCD dynamics [323]. In forward-forward correlations facing the $p(d)$ beam
 2914 direction one selects a large- x parton in the $p(d)$ interacting with a low- x parton in the
 2915 nucleus. For $x < 0.01$ the low- x parton will be back-scattered in the direction of the large-
 2916 x parton. Due to the abundance of gluons at small x , the backwards-scattered partons
 2917 are dominantly gluons, while the large- x partons from the $p(d)$ are dominantly quarks.
 2918 The measurements of di-hadron correlations by STAR and PHENIX [324, 325], have been
 2919 compared with theoretical expectations using the CGC framework based on a fixed saturation
 2920 scale Q_s and considering valence quarks in the deuteron scattering off low- x gluons in the
 2921 nucleus with impact parameter $b = 0$ [326, 327]. Alternative calculations [328] based on both
 2922 initial and final state multiple scattering, which determine the strength of this transverse
 2923 momentum imbalance, in which the suppression of the cross-section in d+Au collisions arises
 2924 from cold nuclear matter energy loss and coherent power corrections have also been very
 2925 successful to describe the data.

2926 The 2015 $p+Au$ run at RHIC has provided unique opportunities to study this channel in
 2927 more detail at STAR. The high delivered integrated luminosities allow one to vary the trigger
 2928 and associated particle p_T from low to high values and thus crossing the saturation boundary
 2929 as shown in Figure 82 and reinstate the correlations for central p+A collisions for forward-
 2930 forward π^0 's. Studying di-hadron correlations in p+A collisions instead of d+A collisions has
 2931 a further advantage. In reference [329], the authors point out that the contributions from
 2932 double-parton interactions to the cross-sections for $dA \rightarrow \pi^0\pi^0 X$ are not negligible. They
 2933 find that such contributions become important at large forward rapidities, and especially in

2934 the case of d+A scattering. Figure 83 shows the results for the di-hadron correlations for π^0
 2935 from the 2015 p+p and p+A run. Shown is the ratio of the area, the width and the level of
 2936 pedestal of the backward peak for p+Au and p+p as function of the p_T of the trigger and
 the associated π^0 and the activity in the collision as measured by the BBC. The results show

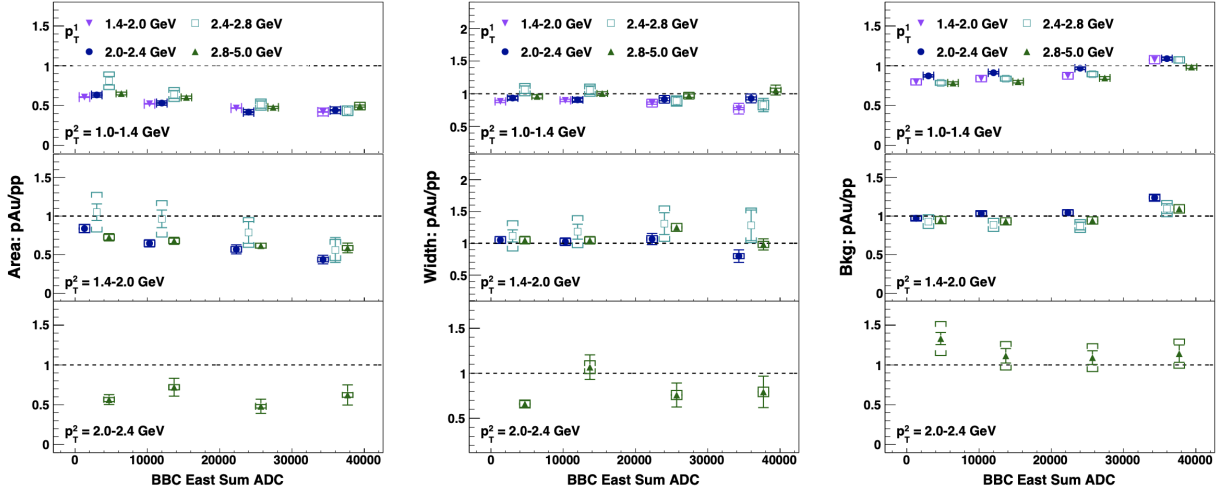


Figure 83: The results for the di-hadron correlations for π^0 from the 2015 p+p and p+A run.

2937 basically no change in the width of the backward peak and the background/pedestal the peak
 2938 is sitting on shows only up to a 20% increase in pA+u to p+p. But the area of the of the
 2939 backward peak shows a large suppression with increasing activity in the collision. For fixed
 2940 activity the biggest suppression is observed for the smallest trigger p_T in combination with
 2941 the smallest p_T for the associated π^0 . This behaviour is consistent with different calculations
 2942 based on the CGC formalism. This result is the first clean observable, which cannot yet
 2943 be explained in a different framework than CGC and as such a clear hint for non-linear
 2944 effects. With the Forward Upgrade several other channels, i.e charged di-hadron and di-jets
 2945 correlations, will also be available, which will allow a rigorous test of the calculation in the
 2946 CGC formalism. It is noted that these results are crucial for the equivalent measurements at
 2947 an EIC, which are planned at close to identical kinematics, because only if non-linear effects
 2948 are seen with different complementary probes, i.e., ep and pA, one can claim a discovery of
 2949 saturation effects and their universality.

2950 It is important to note that for the measurements to date in p(d)+A collisions both
 2951 initial and final states interact strongly, leading to severe complications in the theoretical
 2952 treatment (see [331, 332], and references therein). As described in detail in the Section
 2953 above in p+A collisions, these complications can be ameliorated by removing the strong
 2954 interaction from the final state, by using photons and Drell-Yan electrons. The Run-2015
 2955 p+A run will for the first time provide data on R_{pA} for direct photons and therefore allow
 2956 one to test CGC based predictions on this observable as depicted in Figure 84 (taken from
 2957 Ref. [330]). The higher delivered integrated luminosity for the upcoming p+Au run in
 2958 2024 together with the Forward Upgrade will enable one to study more luminosity hungry
 2959

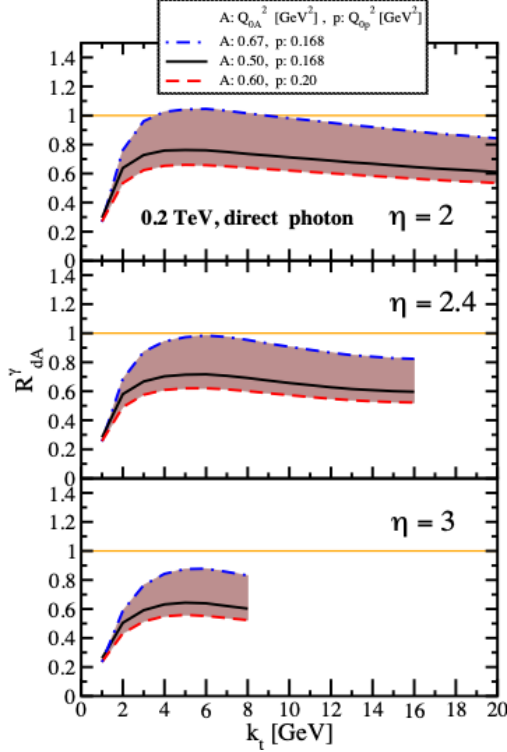


Figure 84: Nuclear modification factor for direct photon production in p(d)A collisions at various rapidities at RHIC $\sqrt{s} = 0.2$ TeV. The curves are the results obtained from Eq. (12) in Ref. [330] and the solution to rcBK equation using different initial saturation scales for a proton Q_{op} and a nucleus Q_{oA} . The band shows our theoretical uncertainties arising from allowing a variation of the initial saturation scale of the nucleus in a range consistent with previous studies of DIS structure functions as well as particle production in minimum-bias p+p, p+A and A+A collisions in the CGC formalism, see Ref. [330] for details.

2960 processes and/or complementary probes to the di- π^0 correlations, i.e. di-hadron correlations
 2961 for charged hadrons, photon-jet, photon-hadron and di-jet correlations.

2962 We use direct photon plus jet (direct γ +jet) events as an example channel to indicate
 2963 what can be done in 2024. These events are dominantly produced through the gluon Compton
 2964 scattering process, $g+q \rightarrow \gamma+q$, and are sensitive to the gluon densities of the nucleon
 2965 and nuclei in p+p and p+A collisions. Through measurements of the azimuthal correlations
 2966 in p+A collisions for direct γ +jet production, one can study gluon saturation phenomena
 2967 at small-x. Unlike di-jet production that is governed by both the Weizsäcker-Williams and
 2968 dipole gluon densities, direct γ +jet production only accesses the dipole gluon density, which
 2969 is better understood theoretically [330, 333]. On the other hand, direct γ +jet production
 2970 is experimentally more challenging due to its small cross-section and large background con-
 2971 tribution from di-jet events in which photons from fragmentation or hadron decay could be
 2972 misidentified as direct photons. The feasibility to perform direct γ +jet measurements with
 2973 the Forward Upgrade in unpolarized p+p and p+Au collisions at $\sqrt{s_{NN}} = 200$ GeV has been
 2974 studied. PYTHIA-8.189 [334] was used to produce direct γ +jet and di-jet events. In order
 2975 to suppress the di-jet background, the leading photon and jet are required to be balanced
 2976 in transverse momentum, $|\phi^\gamma - \phi^{jet}| > 2\pi/3$ and $0.5 < \frac{p_T^\gamma}{p_T^{jet}} < 2$. Both the photon and jet
 2977 have to be in the forward acceptance $1.3 < \eta < 4.0$ with $p_T > 3.2$ GeV/c in 200 GeV p+p
 2978 collisions. The photon needs to be isolated from other particle activities by requiring the
 2979 fraction of electromagnetic energy deposition in the cone of $\Delta R=0.1$ around the photon is
 2980 more than 95% of that in the cone of $\Delta R=0.5$. Jets are reconstructed by an anti- k_T algo-

2981 rithm with $\Delta R=0.5$. After applying these selection cuts, the signal-to-background ratio is
 2982 around 3:1 [335]. The expected number of selected direct γ +jet events is around 1.0M/0.9M
 2983 at $\sqrt{s_{NN}} = 200$ GeV in p+Au collisions for the proposed run in 2024. We conclude that a
 2984 measurement of direct photon-hadron correlation from p+A collisions is feasible, which is
 2985 sensitive to the gluon density in $0.001 < x < 0.005$ in the Au nucleus where parton saturation
 2986 is expected.

2987 **The final state:**

2988 **Nuclear fragmentation functions:** In spite of the remarkable phenomenological suc-
 2989 cesses of QCD, a quantitative understanding of the hadronization process is still one of the
 2990 great challenges for the theory. Hadronization describes the transition of a quark or gluon
 2991 into a final state hadron. It is a poorly understood process even in elementary collisions.
 2992 RHIC’s unique versatility will make it possible to study hadronization in vacuum and in the
 2993 nuclear medium, and additionally with polarized beams (see Sect. 3.2.1 for the latter).

2994 It has long been recognized that the hadron distributions within jets produced in p+p
 2995 collisions are closely related to the fragmentation functions that have typically been measured
 2996 in e^+e^- collisions and SIDIS. The key feature of this type of observable is the possibility to
 2997 determine the relevant momentum fraction z experimentally as the ratio of the hadron to the
 2998 jet transverse momentum. Recently [336] a quantitative relationship has been derived in a
 2999 form that enables measurements of identified hadrons in jets in p+p collisions to be included
 3000 in fragmentation function fits on an equal footing with e^+e^- and SIDIS data. Furthermore,
 3001 hadrons in p+p jets provide unique access to the gluon fragmentation function, which is
 3002 poorly determined in current fits [337], in part due to some tension found in the inclusive
 3003 high p_T pion yields measured by the PHENIX and ALICE collaborations. Here, the proposed
 3004 measurements can provide valuable new insight into the nature of this discrepancy.

3005 This development motivated STAR to initiate a program of identified particle fragmen-
 3006 tation function measurements using p+p jet data at 200 and 500 GeV from 2011, 2012, and
 3007 2015. Figure 85 shows the precision that is anticipated for identified π^+ and π^- in 200 GeV
 3008 p+p collisions for three representative jet p_T bins after the existing data from 2012 and 2015
 3009 are combined with future 200 GeV p+p data from 2024. Identified kaon and (anti)proton
 3010 yields will also be obtained, with somewhat less precision, over a more limited range of hadron
 3011 z . Once the 2017 data are fully analyzed, the uncertainties for 510 GeV p+p collisions will
 3012 be comparable to that shown in Fig. 85 at high jet p_T , and a factor of ~ 2 larger than shown
 3013 in Fig. 85 at low jet p_T . Identified hadron yields will also be measured multi-dimensionally
 3014 vs. j_T , z , and jet p_T , which will provide important input for unpolarized TMD fits.

3015 Data from the HERMES experiment [281, 283, 338] have shown that production rates of
 3016 identified hadrons in semi-inclusive deep inelastic $e+A$ scattering differ from those in $e+p$
 3017 scattering. These differences cannot be explained by nuclear PDFs, as nuclear effects of
 3018 strong interactions in the initial state should cancel in this observable. Only the inclusion of
 3019 nuclear effects in the hadronization process allows theory to reproduce all of the dependencies
 3020 (z , x , and Q^2) of R_{eA} seen in SIDIS, as shown in Fig. 86.

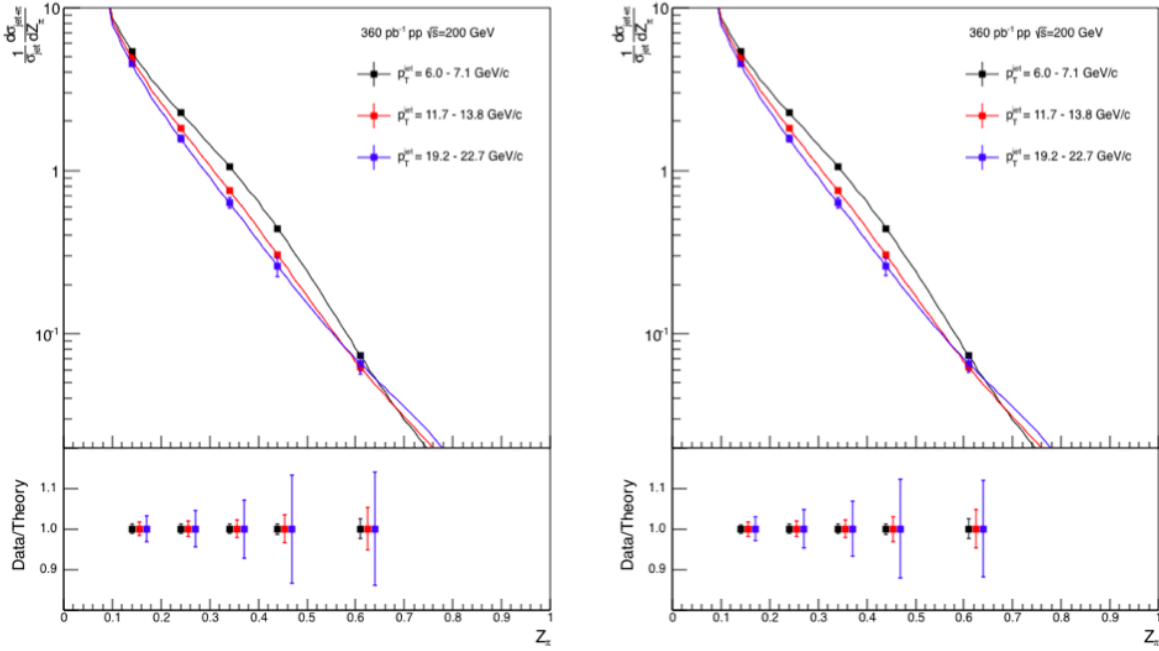


Figure 85: Anticipated precision for identified π^+ (left) and π^- (right) within jets at $|\eta| < 0.4$ in 200 GeV p+p collisions for three representative jet p_T bins. The data points are plotted on theoretical predictions based on the DSS14 pion fragmentation functions [336, 337]. Kaons and (anti)protons will also be measured, over the range from $z < 0.5$ at low jet p_T to $z < 0.2$ at high jet p_T , with uncertainties a factor of ~ 3 larger than those for pions.

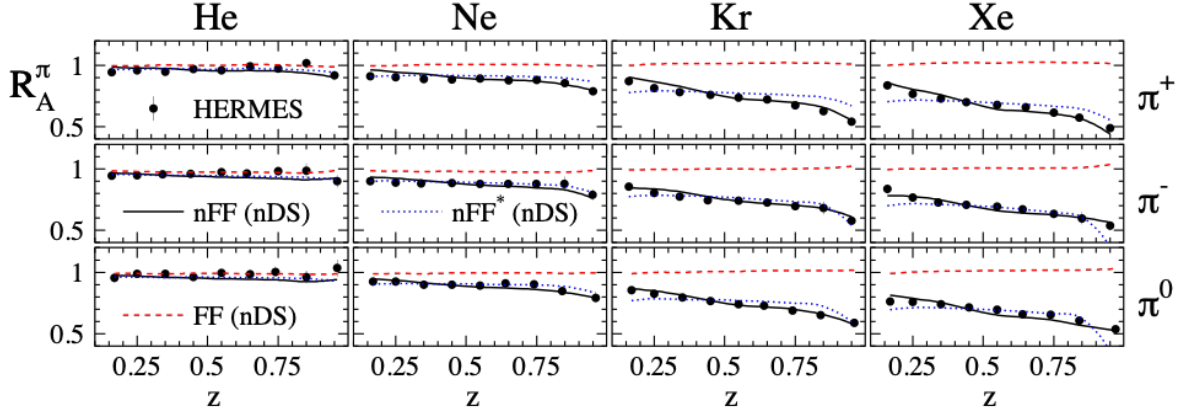


Figure 86: R_{eA} in SIDIS for different nuclei in bins of z as measured by HERMES [281, 283, 338]. The solid lines correspond to the results using effective nuclear FF [290] and the nDS medium modified parton densities [339]. The red dashed lines are estimates assuming the nDS medium modified PDFs but standard DSS vacuum FFs [340, 341] and indicate that nPDFs are insufficient to explain the data

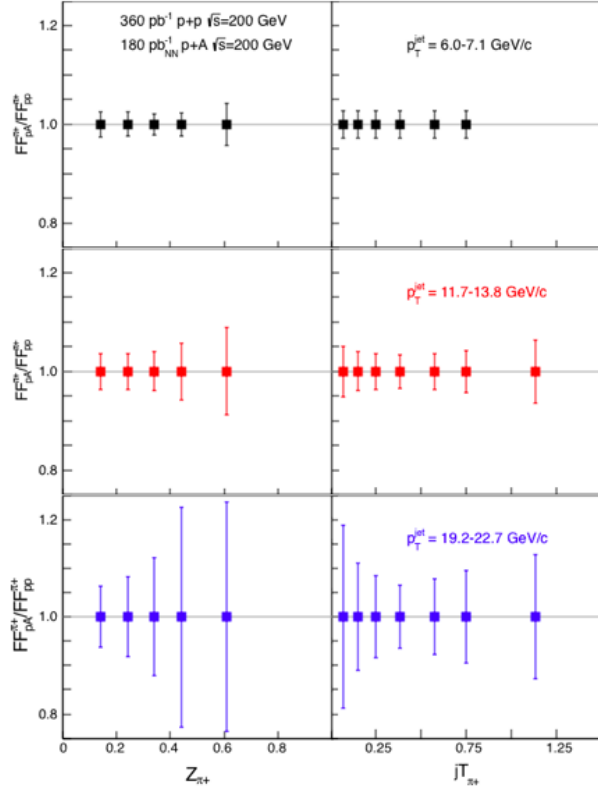


Figure 87: Anticipated precision for measurements of π^+ fragmentation functions in p+A/p+p at $|\eta| < 0.4$ vs. z and j_T in 2024 for three representative jet p_T bins. Uncertainties for π^- will be similar to those shown here for π^+ , while those for kaons and (anti)protons will be a factor of ~ 3 larger. Note that, to be species independent, the nucleon-nucleon equivalent luminosity is specified for p+A.

3021 It is critical to see if these hadronization effects in cold nuclear matter persist at the
 3022 higher \sqrt{s} and Q^2 accessed at RHIC and EIC – both to probe the underlying mechanism,
 3023 which is not understood currently, and to explore its possible universality. The combination
 3024 of p+p jet data from RHIC and future SIDIS data from EIC will also provide a much clearer
 3025 picture of modified gluon hadronization than will be possible with EIC data alone. Using
 3026 the 200 GeV p+Au data collected in 2015, STAR will be able to make a first opportunistic
 3027 measurement of these hadron-jet fragmentation functions in nuclei, but the precision will
 3028 be limited. Additional data will be needed in 2024 in order to provide a sensitive test for
 3029 universality, as shown in Figure 87.

3030 4 Detector Updates, Operations, and Opportunities

3031 In this section we discuss the performance of the endcap Time of Flight (eTOF) in Run-20
3032 and progress of the construction of the Forward upgrades. The iTPC and EPD were fully
3033 integrated for Run-19.

3034 4.1 Status and Performance of the eTOF

3035 The full eTOF hardware installation was completed in Nov. 2018 followed by the first
3036 data taking started in Feb. 2019 by recording about 580 M Au+Au events at $\sqrt{s_{NN}} =$
3037 19.6 GeV with an eTOF participation of 85%. However, due to several beam loss events
3038 causing instantaneous high currents on the readout strips all eTOF preamplifier boards got
3039 damaged and no further useful operation was possible during that year. It was decided to
3040 replace all preamplifier boards with an improved version using ESD protections diodes on
3041 the input. Beside minor issues eTOF showed an excellence performance during Run-20. A
3042 reliable start-up procedure and control interface was implemented that allows the full system
3043 to be controlled via only 2 commands issued by the shift crew. For Run-20 an improved
3044 clock distribution method was installed offering a system synchronization in the order of 35
3045 ps over the full wheel. Figure 88 shows the width of the time distribution (red corresponds
3046 to the Gaussian sigma and blue to the RMS) obtained by measuring the arrival time of
3047 injected pulser signals on every TDC board. The stability of the system is demonstrated on
3048 the right plot of Fig. 88. Here the mean of time distribution width from all pulser channels
3049 is plotted vs. the run number. The range of 130 runs reflects a time period of several days.

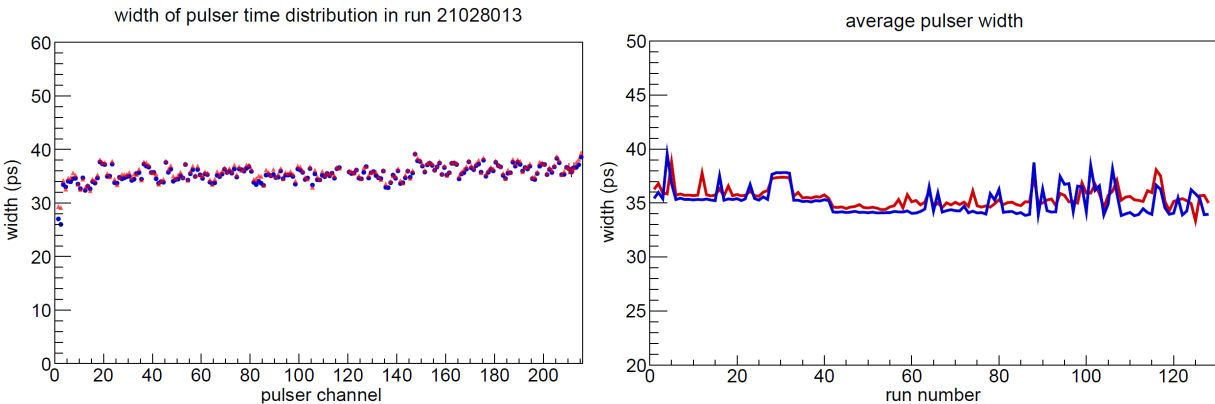


Figure 88: Left: Width of the time distribution obtained by measuring the arrival time of injected pulser signals on every GET4 board. Right: Mean of the time distribution width vs. the run number.

3050 All fixed target runs in 2020 were successfully completed and about 100 M events with
3051 eTOF data were collected for each energy. For the $\sqrt{s_{NN}} = 11.5$ GeV collider run 235 M
3052 events with eTOF data were recorded.

3053 In order to demonstrate the eTOF performance fixed target data at $\sqrt{s_{NN}} = 7.7$ GeV were
3054 calibrated and the matching efficiency with the TPC has been deduced as function of the

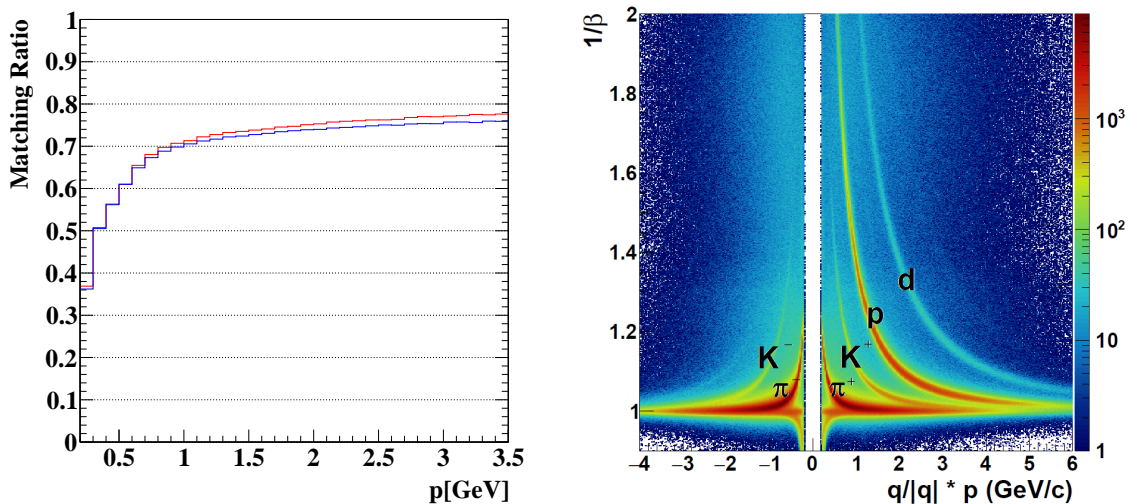


Figure 89: Left: Matching efficiency of MRPC hits in respect to the extrapolated TPC tracks as function of the particle momentum. Right: $1/\beta$ as function of particle momentum. The separation of kaons from pions up to a momenta of 2.5 GeV/c demonstrates the PID capability of eTOF.

3055 particle momentum (see left Fig.89). At a momentum of 1 GeV/c a matching efficiency of
 3056 70% is obtained for both MRPC types (red and blue are different MRPC types with different
 3057 electrode materials). Beyond 1 GeV/c the curve levels off at 75%. The time resolution (not
 3058 shown here) was determined to be in the order of 80 ps. The good time resolution is
 3059 reflected in the $1/\beta$ versus the particle momenta plot shown in the right Fig. 89. The narrow
 3060 particle bands allow for a kaon to pion separation of up to a momenta of 2.5 GeV/c which
 3061 demonstrates the excellent PID capability of eTOF.

3062 For the upcoming period no major hardware changes for eTOF are foreseen. During Run-
 3063 20 one MRPC counter developed a high dark current and noise and will be replaced at the
 3064 next shutdown. Due to COVID-19 travel restrictions it is planed to ship a fully assembled
 3065 module (3 MRPC counters) to BNL as a replacement for the module housing the broken
 3066 counter. On a different module it is planned to replace one GBTx readout card, which is
 3067 currently not working. A substantial eTOF upgrade will be performed on the firmware side
 3068 of the readout FPGAs. This implies also small adaptations in the control software. With
 3069 this upgrade an improved startup reliability and a more stable operation is intended.

3070 4.2 Forward Upgrade

3071 STAR is constructing a forward detector system, realized by combining tracking with elec-
 3072 tromagnetic and hadronic calorimeters for the years beyond 2021. It will have superior
 3073 detection capability for neutral pions, photons, electrons, jets and leading hadrons covering
 3074 a region of $2.5 < \eta < 4$. The design of the Forward Calorimeter System (FCS) is driven by
 3075 consideration of detector performance, integration into STAR and cost optimization. For
 3076 the electromagnetic calorimeter the refurbished PHENIX sampling EMCal is used, and the

3077 hadronic calorimeter will be newly constructed as a sandwich iron scintillator plate sampling
 3078 type, based on the extensive STAR Forward Upgrade and EIC Calorimeter Consortium R&D.
 3079 The existing EPD will be used as a trigger detector especially for a 2 electron trigger. Both
 3080 calorimeters share the same cost-effective readout electronics, with SiPMs as photo-sensors.
 3081 This FCS system will have very good ($\sim 10\%/\sqrt{E}$) electromagnetic and ($\sim 50\%/\sqrt{E} + 10\%$)
 3082 hadronic energy resolutions. In addition, a Forward Tracking System (FTS) is being con-
 3083 structed. The FTS will be capable of discriminating hadron charge sign for transverse
 3084 asymmetry and Drell-Yan measurements in $p+A$. In heavy ion collisions, measurements of
 3085 charged particle transverse momenta of $0.2 < p_T < 2$ GeV/ c with 20-30% momentum res-
 3086 olution are required. To keep multiple scattering and photon conversion background under
 3087 control, the material budget of the FTS must be small. Hence, the FTS design is based on
 3088 three Silicon mini-strip detectors that consist of disks with a wedge-shaped design to cover
 3089 the full azimuth and $2.5 < \eta < 4.0$; they are read out radially from the outside to minimize
 3090 the material. The Si-disks are combined with four small-strip Thin Gap Chamber (sTGC)
 3091 wheels following the ATLAS design [342, 343]. The Si mini-strip disks will be placed in the
 3092 region $z = 146.6 - 173.7$ cm. The 4 sTGC wheels would be placed 30 cm apart starting from
 3093 $z = 273$ cm. The Si-Disks readout is based on APV chips, which will reuse the readout chain
 3094 of the IST, which was part of the STAR HFT. For the sTGC the readout will be based on
 3095 the ATLAS VMM3 chip [344].

3096 4.2.1 Status

3097 Following the successful directors review in November 2018, the project submitted a proposal
 3098 for a NSF MRI for construction of EMCAL and HCAL and the associated electronics. The
 3099 NSF MRI was approved in Summer 2019 and work has been ongoing on all aspects of the
 3100 upgrade.

3101 4.2.2 Forward Calorimetry System

3102 The platform that supports the HCAL and EMCAL was installed in 2019, followed by the
 3103 installation of the refurbished PHENIX EMCAL blocks. The installed EMCAL blocks are
 3104 depicted in Fig. 90. The HCAL absorber blocks are under production at Chapman Lake
 3105 Instrumentation and Getto Industrial Plating. The first sets of blocks have arrived at BNL.
 3106 The scintillating tiles have been produced and all 18200 are in hand. About 10,000 of these
 3107 have been polished at ACU and Valpo and are ready for installation. Other parts are being
 3108 fabricated at Rutgers, Temple and Ohio State. Front-end electronics cards with the SiPM
 3109 and readout for both EMCAL and HCAL are in production and testing is underway. The
 3110 front end cards will be readout by 78 DEP/ADC boards and 3 trigger processor boards
 3111 housed in 5 crates. About half have been delivered and are undergoing testing at BNL.
 3112 All DAQ PCs and receiver cards have been installed. The installation will commence once
 3113 the Run-20b is completed by mid September, and key personnel has come to BNL. The
 3114 commissioning of the FCS will be continue during Run-21, and will be ready for Run-22.

3115 **4.2.3 Forward Silicon Tracking**

3116 The procedures for the Si-detector module fabrication has been developed and documented.
3117 Several prototype mechanical structures with hybrids mounted have been produced and
3118 two wedges were assembled with Si-sensors. Performance of two fully assembled prototype
3119 wedges have been evaluated with cosmic ray data and show that all channels can be read out,
3120 the signal-to-noise meets requirements, and the efficiency is higher than 90%. The design
3121 of the support structures and the interface to the detector modules is nearly complete. If
3122 time allows in the upcoming shutdown a test installation of the support frame into STAR
3123 is planned. The cooling system, which was used previously for the HFT IST sub-system,
3124 has been revived and verification of its performance is on-going. An internal production
3125 readiness review with external reviewers was held on August 3, and the initial steps of the
3126 mass production have started. The review recommendations, which were useful, will be
3127 implemented. Currently there is only limited schedule float for the installation in August
3128 2021.

3129 **4.2.4 sTGC Tracking**

3130 A full prototype module of the sTGC was designed and produced at Shandong University
3131 and tested. This module is now at BNL to undergo testing with the n-pentane gas system
3132 that is being built at BNL. Due to space constraints around the beam pipe the final detector
3133 will have pentagon shaped modules. The design is complete and production has started
3134 of the final pre-production module; mass production is expected to start in October. The
3135 read-out electronics are based on the ATLAS VMM3 chip [344] developed for the same kind
3136 of detector. The strips of each sTGC layer can be handled by 24 Front-End Boards (FEB).
3137 In total 96 FEBs are needed for 4 sTGC layers. The FEBs are vertically inserted in the
3138 sTGC chamber. The signals are send to Readout Boards (ROD) placed in standard VME
3139 crates and interfaced to the STAR DAQ. The electronics design and fabrication is done at
3140 USTC, Hefei. The FEB prototype boards have been tested, the prototype RDO is under
3141 construction, and the VMM3 chips are being procured. The design of the installation and
3142 mounting frames need to be finished. The n-pentane gas system and Interlocks have been
3143 designed and have been approved.

3144 **4.2.5 Software**

3145 The trigger algorithms for the FCS have been well defined and simulated, the FPGA codes
3146 are currently under development.

3147 The forward tracking utilizing hits from the 4 sTGC planes and 3 Si-layers has been de-
3148 veloped and good performance has been demonstrated. As forward tracking is very different
3149 than mid-rapidity tracking new tools had to be developed. The tracking algorithm is based
3150 on modern techniques and depends on GENFIT, a general purpose tracking toolkit and in
3151 addition the iLCSoft KiTrack a Cellular Automata library are used to seed track finding.



Figure 90: A view of the installed forward ECal detector halves, left and right from the beam pipe.

References

- 3152
- 3153 [1] STAR, L. Adamczyk *et al.*, Phys. Rev. C **96**, 024905 (2017), 1702.01108.
- 3154 [2] STAR, R. Kunnawalkam Elayavalli, PoS **HardProbes2018**, 090 (2018), 1903.12115.
- 3155 [3] STAR, J. Adam *et al.*, Phys. Rev. D **101**, 052004 (2020), 1912.08187.
- 3156 [4] STAR, J. Adam *et al.*, Phys. Rev. C **102**, 014905 (2020), 1911.12168.
- 3157 [5] STAR, J. Adam *et al.*, (2019), 1906.09363.
- 3158 [6] STAR, J. Adam *et al.*, (2020), 2003.02114.
- 3159 [7] STAR, J. Adam *et al.*, (2020), 2006.00582.
- 3160 [8] STAR, J. Adam *et al.*, Phys. Rev. D **100**, 052005 (2019), 1906.02740.
- 3161 [9] STAR, R. Licenik, Measurement of inclusive jet production in Au+Au collisions at
3162 $\sqrt{s_{NN}} = 200$ GeV by the STAR experiment, 2020, 2008.10006.
- 3163 [10] STAR, N. R. Sahoo, Measurement of γ +jet and π^0 +jet in central Au+Au collisions
3164 at $\sqrt{s_{NN}} = 200$ GeV with the STAR experiment, 2020, 2008.08789.
- 3165 [11] ATLAS, M. Aaboud *et al.*, Phys. Rev. Lett. **123**, 042001 (2019), 1902.10007.
- 3166 [12] CMS, S. Chatrchyan *et al.*, Phys. Rev. C **90**, 024908 (2014), 1406.0932.
- 3167 [13] STAR, S. Oh, Measurement of semi-inclusive jet fragmentation functions in Au+Au
3168 collisions at $\sqrt{s_{NN}} = 200$ GeV in STAR, 2020, 2008.08631.
- 3169 [14] T. Sjöstrand *et al.*, Comput. Phys. Commun. **191**, 159 (2015), 1410.3012.
- 3170 [15] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **119**, 062301 (2017), 1609.03878.
- 3171 [16] SN0619 : A Proposal for STAR Inner TPC Sector Upgrade (iTPC) [https://drupal.](https://drupal.star.bnl.gov/STAR/starnotes/public/sn0619)
3172 [star.bnl.gov/STAR/starnotes/public/sn0619](https://drupal.star.bnl.gov/STAR/starnotes/public/sn0619).
- 3173 [17] STAR, CBM eTOF Group, The CBM Collaboration eTOF Group, (2016), 1609.05102.
- 3174 [18] J. Adams *et al.*, Nucl. Instrum. Meth. A **968**, 163970 (2020), 1912.05243.
- 3175 [19] STAR, J. Adam *et al.*, Phys. Rev. C **98**, 014910 (2018), 1805.04400.
- 3176 [20] STAR, L. Adamczyk *et al.*, Nature **548**, 62 (2017), 1701.06657.
- 3177 [21] W.-T. Deng and X.-G. Huang, Phys. Rev. C **93**, 064907 (2016), 1603.06117.

- 3178 [22] STAR, J. Adams, Differential measurements of Λ polarization in Au+Au
3179 collisions and a search for the magnetic field by STAR, in *the XXVII-*
3180 *Ith International Conference on Ultra-relativistic Nucleus-Nucleus Collisions*, 2019,
3181 <https://indico.cern.ch/event/792436/contributions/3535689/>.
- 3182 [23] I. Karpenko and F. Becattini, *Eur. Phys. J. C* **77**, 213 (2017), 1610.04717.
- 3183 [24] Y. Jiang, Z.-W. Lin, and J. Liao, *Phys. Rev. C* **94**, 044910 (2016), 1602.06580, [Erra-
3184 tum: *Phys.Rev.C* 95, 049904 (2017)].
- 3185 [25] K. Schilling, P. Seyboth, and G. E. Wolf, *Nucl. Phys. B* **15**, 397 (1970), [Erratum:
3186 *Nucl.Phys.B* 18, 332 (1970)].
- 3187 [26] STAR, S. Singha, Measurement of global spin alignment of k^*0 , $k^*\pm$
3188 and ϕ vector mesons using the STAR detector at RHIC, in *the XXVII-*
3189 *Ith International Conference on Ultra-relativistic Nucleus-Nucleus Collisions*, 2019,
3190 <https://indico.cern.ch/event/792436/contributions/3535685/>.
- 3191 [27] R. Bryan and B. Scott, *Phys. Rev.* **177**, 1435 (1969).
- 3192 [28] M. Nagels, T. Rijken, and J. de Swart, *Phys. Rev. D* **17**, 768 (1978).
- 3193 [29] STAR, L. Adamczyk *et al.*, *Phys. Rev. Lett.* **115**, 222301 (2015), 1505.07812.
- 3194 [30] STAR, S. Huang, Explore the nuclei deformation with a expanding
3195 QGP fireball, in *The 36th Winter Workshop on Nuclear Dynamics*, 2020,
3196 <https://indico.cern.ch/event/841247/contributions/3740391/>.
- 3197 [31] G. Giacalone, *Phys. Rev. Lett.* **124**, 202301 (2020), 1910.04673.
- 3198 [32] U. Heinz and R. Snellings, *Ann. Rev. Nucl. Part. Sci.* **63**, 123 (2013), 1301.2826.
- 3199 [33] STAR, J. Adam *et al.*, (2020), 2006.13537.
- 3200 [34] ALICE, S. Acharya *et al.*, *Phys. Lett.* **B773**, 68 (2017), 1705.04377.
- 3201 [35] P. Alba *et al.*, (2017), 1711.05207.
- 3202 [36] B. Schenke, C. Shen, and P. Tribedy, *Phys. Rev.* **C99**, 044908 (2019), 1901.04378.
- 3203 [37] H. Song, S. A. Bass, U. Heinz, T. Hirano, and C. Shen, *Phys. Rev. Lett.* **106**, 192301
3204 (2011), 1011.2783, [Erratum: *Phys.Rev.Lett.* 109, 139904 (2012)].
- 3205 [38] D. Teaney and L. Yan, *Phys. Rev.* **C83**, 064904 (2011), 1010.1876.
- 3206 [39] Z. Qiu and U. W. Heinz, *Phys.Rev.* **C84**, 024911 (2011), 1104.0650.
- 3207 [40] G. Denicol, H. Niemi, E. Molnar, and D. Rischke, *Phys. Rev. D* **85**, 114047 (2012),
3208 1202.4551, [Erratum: *Phys.Rev.D* 91, 039902 (2015)].

- 3209 [41] W. Florkowski, R. Ryblewski, and M. Spaliński, Phys. Rev. D **94**, 114025 (2016),
3210 1608.07558.
- 3211 [42] K. Dusling and R. Venugopalan, Phys. Rev. Lett. **108**, 262001 (2012), 1201.2658.
- 3212 [43] PHENIX, C. Aidala *et al.*, Nature Phys. **15**, 214 (2019), 1805.02973.
- 3213 [44] CMS, S. Chatrchyan *et al.*, Phys. Lett. B **724**, 213 (2013), 1305.0609.
- 3214 [45] STAR, R. A. Lacey, Long-range collectivity in small collision-systems with two- and
3215 four-particle correlations @ STAR, in *28th International Conference on Ultrarelativistic*
3216 *Nucleus-Nucleus Collisions*, 2020, 2002.11889.
- 3217 [46] L. D. McLerran and T. Toimela, Phys. Rev. D **31**, 545 (1985).
- 3218 [47] H. Weldon, Phys. Rev. D **42**, 2384 (1990).
- 3219 [48] R. Rapp and J. Wambach, Eur. Phys. J. A **6**, 415 (1999), hep-ph/9907502.
- 3220 [49] P. M. Hohler and R. Rapp, Phys. Lett. B **731**, 103 (2014), 1311.2921.
- 3221 [50] C. Jung and L. von Smekal, Phys. Rev. D **100**, 116009 (2019), 1909.13712.
- 3222 [51] H.-T. Ding *et al.*, Phys. Rev. D **83**, 034504 (2011), 1012.4963.
- 3223 [52] J. Atchison and R. Rapp, J. Phys. Conf. Ser. **832**, 012057 (2017).
- 3224 [53] R. Rapp and H. van Hees, Phys. Lett. B **753**, 586 (2016), 1411.4612.
- 3225 [54] R. Chatterjee, D. K. Srivastava, U. W. Heinz, and C. Gale, Phys. Rev. C **75**, 054909
3226 (2007), nucl-th/0702039.
- 3227 [55] J. Deng, Q. Wang, N. Xu, and P. Zhuang, Phys. Lett. B **701**, 581 (2011), 1009.3091.
- 3228 [56] P. Mohanty *et al.*, Phys. Rev. C **85**, 031903 (2012), 1111.2159.
- 3229 [57] E. Speranza, A. Jaiswal, and B. Friman, Phys. Lett. B **782**, 395 (2018), 1802.02479.
- 3230 [58] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **113**, 022301 (2014), 1312.7397, [Adden-
3231 dum: Phys.Rev.Lett. 113, 049903 (2014)].
- 3232 [59] STAR, L. Adamczyk *et al.*, Phys. Rev. C **92**, 024912 (2015), 1504.01317.
- 3233 [60] STAR, L. Adamczyk *et al.*, Phys. Lett. B **750**, 64 (2015), 1501.05341.
- 3234 [61] STAR, J. Adam *et al.*, (2018), 1810.10159.
- 3235 [62] B. B. Brandt, A. Francis, H. B. Meyer, and H. Wittig, JHEP **03**, 100 (2013), 1212.4200.
- 3236 [63] STAR, J. Adam *et al.*, (2019), 1910.12400.

- 3237 [64] C. Li, J. Zhou, and Y.-J. Zhou, Phys. Rev. D **101**, 034015 (2020), 1911.00237.
- 3238 [65] S. Klein, A. Mueller, B.-W. Xiao, and F. Yuan, (2020), 2003.02947.
- 3239 [66] STAR Collaboration, J. Adam *et al.*, Phys. Rev. Lett. **124**, 172301 (2020).
- 3240 [67] J. Adam *et al.*, Physics Letters B **797**, 134917 (2019).
- 3241 [68] STAR Collaboration, J. Adam *et al.*, Phys. Rev. Lett. **123**, 162301 (2019).
- 3242 [69] STAR Collaboration, J. Adam *et al.*, Phys. Rev. C **99**, 034908 (2019).
- 3243 [70] M. Djordjevic, Phys. Rev. Lett. **112**, 042302 (2014).
- 3244 [71] H. van Hees, M. Mannarelli, V. Greco, and R. Rapp, Phys. Rev. Lett. **100**, 192301
3245 (2008).
- 3246 [72] S. Cao, G.-Y. Qin, and S. A. Bass, Phys. Rev. C **92**, 024907 (2015).
- 3247 [73] P. B. Gossiaux and J. Aichelin, Phys. Rev. C **78**, 014904 (2008).
- 3248 [74] M. Djordjevic and M. Djordjevic, Phys. Rev. C **90**, 034910 (2014).
- 3249 [75] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. Lett. **118**, 212301 (2017).
- 3250 [76] STAR Collaboration, L. Adamczyk *et al.*, Phys. Rev. C **95**, 034907 (2017).
- 3251 [77] D. Kharzeev and R. D. Pisarski, Phys.Rev. **D61**, 111901 (2000), hep-ph/9906401.
- 3252 [78] D. Kharzeev, Phys.Lett. **B633**, 260 (2006), hep-ph/0406125.
- 3253 [79] S. A. Voloshin, Phys. Rev. Lett. **105**, 172301 (2010), 1006.1020.
- 3254 [80] W.-T. Deng, X.-G. Huang, G.-L. Ma, and G. Wang, Phys. Rev. C **97**, 044901 (2018),
3255 1802.02292.
- 3256 [81] H.-J. Xu *et al.*, Phys. Rev. Lett. **121**, 022301 (2018), 1710.03086.
- 3257 [82] J. Hammelmann, A. Soto-Ontoso, M. Alvioli, H. Elfner, and M. Strikman, (2019),
3258 1908.10231.
- 3259 [83] STAR 2017 BUR [https://drupal.star.bnl.gov/STAR/system/files/STAR_BUR_](https://drupal.star.bnl.gov/STAR/system/files/STAR_BUR_Run1718_v22_0.pdf)
3260 [Run1718_v22_0.pdf](https://drupal.star.bnl.gov/STAR/system/files/STAR_BUR_Run1718_v22_0.pdf).
- 3261 [84] STAR, J. Adam *et al.*, (2019), 1911.00596.
- 3262 [85] J. Zhao, H. Li, and F. Wang, Eur. Phys. J. C **79**, 168 (2019), 1705.05410.
- 3263 [86] STAR, J. Adam *et al.*, (2020), arXiv:2006.05035.

- 3264 [87] STAR, P. Tribedy, Nucl. Phys. A **967**, 740 (2017), 1704.03845.
- 3265 [88] STAR, J. Zhao, (2020), 2002.09410.
- 3266 [89] CMS, A. M. Sirunyan *et al.*, Phys. Rev. **C97**, 044912 (2018), 1708.01602.
- 3267 [90] H.-j. Xu *et al.*, Chin. Phys. **C42**, 084103 (2018), 1710.07265.
- 3268 [91] S. A. Voloshin, Phys. Rev. **C98**, 054911 (2018), 1805.05300.
- 3269 [92] STAR, J. Adam *et al.*, (2020), arXiv:2006.04251.
- 3270 [93] STAR, Y. Lin, (2020), 2002.11446.
- 3271 [94] N. Magdy, S. Shi, J. Liao, N. Ajitanand, and R. A. Lacey, Phys. Rev. C **97**, 061901
3272 (2018), 1710.01717.
- 3273 [95] A. Tang, Chin. Phys. C **44**, 054101 (2020), 1903.04622.
- 3274 [96] L. Finch and S. Murray, Phys. Rev. C **96**, 044911 (2017), 1801.06476.
- 3275 [97] STAR, M. Zurek, Longitudinal double-spin asymmetries for inclusive jets produced
3276 in $\sqrt{s} = 200$ GeV proton-proton collisions at STAR, in *Nuclear Physics Seminar,*
3277 *Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8632/>.
- 3278 [98] STAR, N. Lukow, Longitudinal double-spin asymmetries for dijets pro-
3279 duced in $\sqrt{s} = 200$ GeV proton-proton collisions at STAR, 2020,
3280 <https://drupal.star.bnl.gov/STAR/blog/nlukow/dijet-all-2015-preliminary>.
- 3281 [99] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **115**, 092002 (2015), 1405.5134.
- 3282 [100] STAR, L. Adamczyk *et al.*, Phys. Rev. **D95**, 071103 (2017), 1610.06616.
- 3283 [101] D. de Florian, R. Sassot, M. Stratmann, and W. Vogelsang, Phys. Rev. Lett. **113**,
3284 012001 (2014), 1404.4293.
- 3285 [102] NNPDF, E. R. Nocera, R. D. Ball, S. Forte, G. Ridolfi, and J. Rojo, Nucl. Phys.
3286 **B887**, 276 (2014), 1406.5539.
- 3287 [103] J. Adam *et al.*, Physical Review D **100** (2019).
- 3288 [104] STAR, T. Lin, Azimuthal Transverse Single-Spin Asymmetries of Charged Pions
3289 Within Jets from Polarized pp Collisions at $\sqrt{s} = 200$ GeV, in *Nuclear Physics Semi-*
3290 *nar, Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8633/>.
- 3291 [105] STAR, H. Liu, Measurement of transverse single-spin asymmetries for dijet production
3292 in polarized p+p collisions at $\sqrt{s} = 200$ GeV at STAR, in *Nuclear Physics Seminar,*
3293 *Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8633/>.

- 3294 [106] STAR, S. Heppelmann, Preview from RHIC Run 15 pp and pAu Forward Neutral
3295 Pion Production from Transversely Polarized Protons, in *7th International Workshop*
3296 *on Multiple Partonic Interactions at the LHC*, pp. 228–231, 2016.
- 3297 [107] STAR, Z. Zhu, Measurement of transverse single spin asymmetries at forward rapidities
3298 by the STAR experiment at $\sqrt{s} = 200$ and 500 GeV at STAR, in *Nuclear Physics*
3299 *Seminar, Brookhaven National Laboratory*, 2020, <https://indico.bnl.gov/event/8632/>.
- 3300 [108] Z.-B. Kang, X. Liu, F. Ringer, and H. Xing, *JHEP* **11**, 068 (2017), 1705.08443.
- 3301 [109] Z.-B. Kang, A. Prokudin, F. Ringer, and F. Yuan, *Phys. Lett. B* **774**, 635 (2017),
3302 1707.00913.
- 3303 [110] U. D’Alesio, F. Murgia, and C. Pisano, *Phys. Rev. D* **83**, 034021 (2011), 1011.2692.
- 3304 [111] J. K. Adkins, *Studying Transverse Momentum Dependent Distributions in Polarized*
3305 *Proton Collisions via Azimuthal Single Spin Asymmetries of Charged Pions in Jets*,
3306 PhD thesis, University of Kentucky, 2015, 1907.11233.
- 3307 [112] STAR, B. Abelev *et al.*, *Phys. Rev. Lett.* **99**, 142003 (2007), 0705.4629.
- 3308 [113] J. Cammarota *et al.*, (2020), 2002.08384.
- 3309 [114] STAR, E. Braidot, Suppression of Forward Pion Correlations in d+Au Interactions
3310 at STAR, in *45th Rencontres de Moriond on QCD and High Energy Interactions*, pp.
3311 355–338, 2010, 1005.2378.
- 3312 [115] STAR, M. Posik, PoS **DIS2019**, 028 (2019), 1908.08490.
- 3313 [116] NuSea, R. S. Towell *et al.*, *Phys. Rev.* **D64**, 052002 (2001), hep-ex/0103030.
- 3314 [117] SeaQuest, K. Nagai, PoS **DIS2018**, 033 (2018).
- 3315 [118] STAR, S. Fazio, Measurements of weak boson production with the STAR detector,
3316 in *QCD Spin Physics: A Symposium to Honor Jacques Soffer, Brookhaven National*
3317 *Laboratory*, 2019, <https://indico.bnl.gov/event/6274/>.
- 3318 [119] V. Bertone, I. Scimemi, and A. Vladimirov, *JHEP* **06**, 028 (2019), 1902.08474.
- 3319 [120] M. A. Stephanov, K. Rajagopal, and E. V. Shuryak, *Phys. Rev. D* **60**, 114028 (1999),
3320 hep-ph/9903292.
- 3321 [121] M. Stephanov, *Phys. Rev. Lett.* **107**, 052301 (2011), 1104.1627.
- 3322 [122] STAR, J. Adam *et al.*, (2020), 2001.02852.
- 3323 [123] HADES, J. Adamczewski-Musch *et al.*, (2020), 2002.08701.
- 3324 [124] STAR, L. Adamczyk *et al.*, *Phys. Rev.* **C96**, 044904 (2017), 1701.07065.

- 3325 [125] F. Becattini, J. Manninen, and M. Gazdzicki, Phys. Rev. C **73**, 044905 (2006), hep-
3326 ph/0511092.
- 3327 [126] A. Andronic, P. Braun-Munzinger, and J. Stachel, Phys. Lett. B **673**, 142 (2009),
3328 0812.1186, [Erratum: Phys.Lett.B 678, 516 (2009)].
- 3329 [127] J. Cleymans, D. Elliott, A. Keranen, and E. Suhonen, Phys. Rev. C **57**, 3319 (1998),
3330 nucl-th/9711066.
- 3331 [128] J. Cleymans, A. Forster, H. Oeschler, K. Redlich, and F. Uhlig, Phys. Lett. B **603**,
3332 146 (2004), hep-ph/0406108.
- 3333 [129] I. Kraus, J. Cleymans, H. Oeschler, K. Redlich, and S. Wheaton, Phys. Rev. C **76**,
3334 064903 (2007), 0707.3879.
- 3335 [130] HADES, J. Adamczewski-Musch *et al.*, Phys. Lett. B **778**, 403 (2018), 1703.08418.
- 3336 [131] FOPI, K. Piasecki *et al.*, Phys. Rev. C **94**, 014901 (2016), 1602.04378.
- 3337 [132] A. Andronic, P. Braun-Munzinger, J. Stachel, and H. Stocker, Phys. Lett. B **697**, 203
3338 (2011), 1010.2995.
- 3339 [133] NA49, T. Anticic *et al.*, Phys. Rev. **C83**, 014901 (2011), 1009.1747.
- 3340 [134] BRAHMS, I. Arsene *et al.*, Phys. Rev. Lett. **93**, 242303 (2004), nucl-ex/0403005.
- 3341 [135] Yu. B. Ivanov, Phys. Rev. **C87**, 064905 (2013), 1304.1638.
- 3342 [136] A. Bzdak, V. Koch, D. Oliinychenko, and J. Steinheimer, Phys. Rev. **C98**, 054901
3343 (2018), 1804.04463.
- 3344 [137] M. Danysz *et al.*, Phys. Rev. Lett. **11**, 29 (1963).
- 3345 [138] D. Prowse, Phys. Rev. Lett. **17**, 782 (1966).
- 3346 [139] S. Aoki *et al.*, Prog. Theor. Phys. **85**, 1287 (1991).
- 3347 [140] H. Takahashi *et al.*, Nucl. Phys. A **721**, 951 (2003).
- 3348 [141] T. A. Rijken, M. Nagels, and Y. Yamamoto, Prog. Theor. Phys. Suppl. **185**, 14 (2010).
- 3349 [142] K. J. Sun, L.-W. Chen, C. M. Ko, J. Pu, and Z. Xu, Phys. Lett. **B781**, 499 (2018),
3350 1801.09382.
- 3351 [143] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **112**, 032302 (2014), 1309.5681.
- 3352 [144] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **113**, 092301 (2014), 1402.1558.
- 3353 [145] STAR, L. Adamczyk *et al.*, Phys. Lett. **B785**, 551 (2018), 1709.00773.

- 3354 [146] STAR, X. Luo, PoS **CPOD2014**, 019 (2015), 1503.02558.
- 3355 [147] X. Luo, Nucl. Phys. **A956**, 75 (2016), 1512.09215.
- 3356 [148] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **121**, 032301 (2018), 1707.01988.
- 3357 [149] STAR, J. Adam *et al.*, accepted by PRC **arXiv:1903.11778** (2019), 1903.11778.
- 3358 [150] STAR, D. Zhang, Light Nuclei (d , t) Production in Au + Au Collisions at $\sqrt{s_{NN}} =$
3359 7.7 - 200 GeV, in *28th International Conference on Ultrarelativistic Nucleus-Nucleus*
3360 *Collisions*, 2020, 2002.10677.
- 3361 [151] H. Liu *et al.*, Phys. Lett. B **805**, 135452 (2020), 1909.09304.
- 3362 [152] K. Dusling, W. Li, and B. Schenke, Int. J. Mod. Phys. **E25**, 1630002 (2016),
3363 1509.07939.
- 3364 [153] L. He *et al.*, Phys. Lett. **B753**, 506 (2016), 1502.05572.
- 3365 [154] A. Kurkela, U. A. Wiedemann, and B. Wu, Kinetic transport is needed to reliably ex-
3366 tract shear viscosity from pA and AA data, in Vol. arXiv:1805.04081, 2018, 1805.04081.
- 3367 [155] P. Romatschke, Eur. Phys. J. **C78**, 636 (2018), 1802.06804.
- 3368 [156] J. L. Nagle and W. A. Zajc, Ann. Rev. Nucl. Part. Sci. **68**, 211 (2018), 1801.03477.
- 3369 [157] B. Schenke, S. Schlichting, and R. Venugopalan, Phys. Lett. **B747**, 76 (2015),
3370 1502.01331.
- 3371 [158] J. L. Nagle *et al.*, Phys. Rev. Lett. **113**, 112301 (2014), 1312.4565.
- 3372 [159] PHENIX, A. Adare *et al.*, Phys. Rev. Lett. **111**, 212301 (2013), 1303.1794.
- 3373 [160] STAR, S. Huang, Nucl. Phys. A **982**, 475 (2019).
- 3374 [161] B. Schenke, C. Shen, and P. Tribedy, (2020), 2005.14682.
- 3375 [162] P. Romatschke, Eur. Phys. J. C **75**, 305 (2015), 1502.04745.
- 3376 [163] C. Shen, J.-F. Paquet, G. S. Denicol, S. Jeon, and C. Gale, Phys. Rev. C **95**, 014906
3377 (2017), 1609.02590.
- 3378 [164] Workshop on collectivity of small systems in high-energy collisions, 14-16 march 2019,
3379 <https://indico.cern.ch/event/771998/>, Accessed May 2019.
- 3380 [165] STAR, STAR Public Note SN0644 - Technical Design Report for the iTPC Upgrade,
3381 <https://drupal.star.bnl.gov/STAR/starnotes/public/sn0644>.

- 3382 [166] STAR, STAR Public Note SN0666 - An Event Plane Detector for STAR, <https://drupal.star.bnl.gov/STAR/starnotes/public/sn0666>.
3383
- 3384 [167] STAR, STAR Public Note SN0648 - The STAR Forward Calorimeter System and
3385 Forward Tracking System beyond BES-II, <https://drupal.star.bnl.gov/STAR/starnotes/public/sn0648>.
3386
- 3387 [168] Z. Citron *et al.* *Report from Working Group 5: Future physics opportunities for high-*
3388 *density QCD at the LHC with heavy-ion and proton beams* Vol. 7 (, 2019), pp. 1159–
3389 1410, 1812.06772.
- 3390 [169] ALICE, K. Aamodt *et al.*, Phys. Rev. Lett. **105**, 252302 (2010), 1011.3914.
- 3391 [170] PHENIX, S. S. Adler *et al.*, Phys. Rev. Lett. **94**, 232302 (2005), nucl-ex/0411040.
- 3392 [171] STAR, J. Adam *et al.*, Phys. Lett. **B784**, 26 (2018), 1804.08647.
- 3393 [172] ATLAS, M. Aaboud *et al.*, Eur. Phys. J. **C78**, 997 (2018), 1808.03951.
- 3394 [173] PHENIX, A. Adare *et al.*, Phys. Rev. Lett. **107**, 252301 (2011), 1105.3928.
- 3395 [174] ATLAS, G. Aad *et al.*, Phys. Rev. C **90**, 044906 (2014), 1409.1792.
- 3396 [175] STAR, B. I. Abelev *et al.*, Phys. Rev. Lett. **99**, 112301 (2007), nucl-ex/0703033.
- 3397 [176] CMS Collaboration, Phys. Lett. B **765**, 193 (2017), 1606.06198.
- 3398 [177] CMS Collaboration, Phys. Rev. Lett. **115**, 012301 (2015), 1502.05382.
- 3399 [178] ATLAS, M. Aaboud *et al.*, Phys. Lett. **B789**, 444 (2019), 1807.02012.
- 3400 [179] ALICE, S. Acharya *et al.*, Investigations of anisotropic flow using multi-particle
3401 azimuthal correlations in pp, p-Pb, Xe-Xe, and Pb-Pb collisions at the LHC,
3402 arXiv:1903.01790, 2019, 1903.01790.
- 3403 [180] PHENIX, C. Aidala *et al.*, Phys. Rev. Lett. **120**, 062302 (2018), 1707.06108.
- 3404 [181] A. Huss *et al.*, (2020), 2007.13758.
- 3405 [182] A. Huss *et al.*, (2020), 2007.13754.
- 3406 [183] I. Arsene *et al.*, Phys. Rev. C **75**, 034902 (2007), nucl-th/0609042.
- 3407 [184] J. Harris *et al.*, Phys. Lett. B **153**, 377 (1985).
- 3408 [185] R. Stock, Phys. Rept. **135**, 259 (1986).
- 3409 [186] J. Molitoris and H. Stoecker, Phys. Rev. C **32**, 346 (1985).

- 3410 [187] H. Kruse, B. Jacak, and H. Stoecker, Phys. Rev. Lett. **54**, 289 (1985).
- 3411 [188] P. Danielewicz, R. Lacey, and W. G. Lynch, Science **298**, 1592 (2002), nucl-th/0208016.
- 3412 [189] W. Reisdorf, Heavy ion collisions in the 1A GeV regime: how well do we join up to
3413 astrophysics?, in *International Workshop on Nuclear Symmetry Energy and Reaction*
3414 *Mechanisms*, 2013, 1307.4210.
- 3415 [190] J. Aichelin and C. M. Ko, Phys. Rev. Lett. **55**, 2661 (1985).
- 3416 [191] KAOS, C. T. Sturm *et al.*, Phys. Rev. Lett. **86**, 39 (2001), nucl-ex/0011001.
- 3417 [192] C. Hartnack and J. Aichelin, J. Phys. G **28**, 1649 (2002).
- 3418 [193] C. Fuchs *et al.*, J. Phys. G **28**, 1615 (2002), nucl-th/0103036.
- 3419 [194] T. Klahn *et al.*, Phys. Rev. C **74**, 035802 (2006), nucl-th/0602038.
- 3420 [195] T. Fischer *et al.*, Nature Astron. **2**, 980 (2018), 1712.08788.
- 3421 [196] A. Aprahamian *et al.*, Reaching for the horizon: The 2015 long range plan for
3422 nuclear science, 2015, The 2015 Long Range Plan for Nuclear Science “Reach-
3423 ing for the Horizon” [http://science.energy.gov/~media/np/nsac/pdf/2015LRP/](http://science.energy.gov/~media/np/nsac/pdf/2015LRP/2015_LRPNS_091815.pdf)
3424 [2015_LRPNS_091815.pdf](http://science.energy.gov/~media/np/nsac/pdf/2015LRP/2015_LRPNS_091815.pdf).
- 3425 [197] RHIC luminosity projection [https://www.rhichome.bnl.gov//RHIC/Runs/](https://www.rhichome.bnl.gov//RHIC/Runs/RhicProjections.pdf)
3426 [RhicProjections.pdf](https://www.rhichome.bnl.gov//RHIC/Runs/RhicProjections.pdf).
- 3427 [198] The STAR Forward Calorimeter System and Forward Tracking System beyond
3428 BES-II [https://drupal.star.bnl.gov/STAR/files/Proposal.ForwardUpgrade.](https://drupal.star.bnl.gov/STAR/files/Proposal.ForwardUpgrade.Nov_.2018.Review.pdf)
3429 [Nov_.2018.Review.pdf](https://drupal.star.bnl.gov/STAR/files/Proposal.ForwardUpgrade.Nov_.2018.Review.pdf).
- 3430 [199] G. Denicol, A. Monnai, and B. Schenke, Phys. Rev. Lett. **116**, 212301 (2016),
3431 1512.01538.
- 3432 [200] H. Niemi, G. S. Denicol, P. Huovinen, E. Molnar, and D. H. Rischke, Phys. Rev. **C86**,
3433 014909 (2012), 1203.2452.
- 3434 [201] B. Schenke and S. Schlichting, Phys. Rev. **C94**, 044907 (2016), 1605.07158.
- 3435 [202] W. Li, Nucl. Phys. A **967**, 59 (2017), 1704.03576.
- 3436 [203] P. Bozek, W. Broniowski, and J. Moreira, Phys. Rev. **C83**, 034911 (2011), 1011.3354.
- 3437 [204] J. Jia and P. Huo, Phys. Rev. **C90**, 034915 (2014), 1403.6077.
- 3438 [205] L.-G. Pang, H. Petersen, G.-Y. Qin, V. Roy, and X.-N. Wang, Eur. Phys. J. **A52**, 97
3439 (2016), 1511.04131.

- 3440 [206] CMS, V. Khachatryan *et al.*, Phys. Rev. C **92**, 034911 (2015), 1503.01692.
- 3441 [207] ATLAS, M. Aaboud *et al.*, Eur. Phys. J. C **78**, 142 (2018), 1709.02301.
- 3442 [208] STAR Collaboration, L. Adamczyk *et al.*, Manuscript under preparation (2016).
- 3443 [209] A. Behera, M. Nie, and J. Jia, Phys. Rev. Res. **2**, 023362 (2020), 2003.04340.
- 3444 [210] C. Shen and B. Schenke, Phys. Rev. **C97**, 024907 (2018), 1710.00881.
- 3445 [211] STAR, B. Abelev *et al.*, Phys. Rev. C **76**, 064904 (2007), 0706.0472.
- 3446 [212] Z.-T. Liang, J. Song, I. Upsal, Q. Wang, and Z.-B. Xu, (2019), 1912.10223.
- 3447 [213] Y. Xie, D. Wang, and L. P. Csernai, Eur. Phys. J. C **80**, 39 (2020), 1907.00773.
- 3448 [214] X.-L. Sheng, L. Oliva, and Q. Wang, Phys. Rev. D **101**, 096005 (2020), 1910.13684.
- 3449 [215] X.-L. Sheng, Q. Wang, and X.-N. Wang, (2020), 2007.05106.
- 3450 [216] A. Bazavov *et al.*, Phys. Rev. D **95**, 054504 (2017), 1701.04325.
- 3451 [217] S. Borsanyi *et al.*, JHEP **10**, 205 (2018), 1805.04445.
- 3452 [218] STAR, A. Pandav, (2020), 2003.12503.
- 3453 [219] B. Chen, M. Hu, H. Zhang, and J. Zhao, Phys. Lett. B **802**, 135271 (2020), 1910.08275.
- 3454 [220] X. Guo, S. Shi, N. Xu, Z. Xu, and P. Zhuang, Phys. Lett. B **751**, 215 (2015),
3455 1502.04407.
- 3456 [221] B. Chen, X. Du, and R. Rapp, Nucl. Part. Phys. Proc. **289-290**, 475 (2017),
3457 1612.02089.
- 3458 [222] ALICE, J. Adam *et al.*, JHEP **05**, 179 (2016), 1506.08804.
- 3459 [223] ATLAS, M. Aaboud *et al.*, Eur. Phys. J. C **78**, 762 (2018), 1805.04077.
- 3460 [224] CMS, A. M. Sirunyan *et al.*, Phys. Rev. Lett. **118**, 162301 (2017), 1611.01438.
- 3461 [225] JETSCAPE, S. Cao *et al.*, Phys. Rev. C **96**, 024909 (2017), 1705.00050.
- 3462 [226] JETSCAPE, A. Kumar *et al.*, Jet quenching in a multi-stage Monte Carlo approach,
3463 in *28th International Conference on Ultrarelativistic Nucleus-Nucleus Collisions*, 2020,
3464 2002.07124.
- 3465 [227] Y. Mehtar-Tani and K. Tywoniuk, Phys. Rev. D **98**, 051501 (2018), 1707.07361.
- 3466 [228] Y. Mehtar-Tani and K. Tywoniuk, Nucl. Phys. A **979**, 165 (2018), 1706.06047.

- 3467 [229] B. Zakharov, (2020), 2003.10182.
- 3468 [230] L. Chen, G.-Y. Qin, S.-Y. Wei, B.-W. Xiao, and H.-Z. Zhang, Phys. Lett. B **773**, 672
3469 (2017), 1607.01932.
- 3470 [231] D. A. Appel, Phys. Rev. D **33**, 717 (1986).
- 3471 [232] J. Blaizot and L. D. McLerran, Phys. Rev. D **34**, 2739 (1986).
- 3472 [233] ALICE, J. Adam *et al.*, JHEP **09**, 170 (2015), 1506.03984.
- 3473 [234] F. D’Eramo, K. Rajagopal, and Y. Yin, JHEP **01**, 172 (2019), 1808.03250.
- 3474 [235] ALICE, J. Norman.
- 3475 [236] A. J. Larkoski, S. Marzani, G. Soyez, and J. Thaler, JHEP **05**, 146 (2014), 1402.2657.
- 3476 [237] Y.-T. Chien and R. Kunnawalkam Elayavalli, (2018), 1803.03589.
- 3477 [238] J.-w. Qiu and G. F. Sterman, Phys. Rev. D **59**, 014004 (1999), hep-ph/9806356.
- 3478 [239] K. Kanazawa, Y. Koike, A. Metz, and D. Pitonyak, Phys. Rev. D **89**, 111501 (2014).
- 3479 [240] L. Bland *et al.*, Physics Letters B **750**, 660 (2015).
- 3480 [241] STAR, L. Adamczyk *et al.*, Phys. Rev. Lett. **116**, 132301 (2016), 1511.06003.
- 3481 [242] P. Sun, J. Isaacson, C. P. Yuan, and F. Yuan, Universal non-perturbative functions
3482 for sidis and drell-yan processes, 2014, 1406.3073.
- 3483 [243] J. P. Ralston and D. E. Soper, Nuclear Physics B **152**, 109 (1979).
- 3484 [244] R. Jaffe and X.-D. Ji, Nucl. Phys. B **375**, 527 (1992).
- 3485 [245] P. Mulders and R. Tangerman, Nucl. Phys. B **461**, 197 (1996), hep-ph/9510301,
3486 [Erratum: Nucl.Phys.B 484, 538–540 (1997)].
- 3487 [246] D. Sivers, Nuovo Cim. C **035N2**, 171 (2012), 1109.2521.
- 3488 [247] C. Alexandrou *et al.*, Phys. Rev. D **98**, 091503 (2018).
- 3489 [248] R. Jaffe and X.-D. Ji, Phys. Rev. Lett. **67**, 552 (1991).
- 3490 [249] J. C. Collins, S. F. Heppelmann, and G. A. Ladinsky, Nucl. Phys. B **420**, 565 (1994),
3491 hep-ph/9305309.
- 3492 [250] L. Adamczyk *et al.*, Phys. Lett. B **780**, 332–339 (2018).
- 3493 [251] M. Radici and A. Bacchetta, Phys. Rev. Lett. **120**, 192001 (2018), 1802.05212.

- 3494 [252] J. Collins and J.-W. Qiu, Phys. Rev. D **75**, 114014 (2007).
- 3495 [253] T. C. Rogers and P. J. Mulders, Phys. Rev. D **81**, 094006 (2010).
- 3496 [254] X. Liu, F. Ringer, W. Vogelsang, and F. Yuan, Factorization and its breaking in dijet
3497 single transverse spin asymmetries in pp collisions, 2020, 2008.03666.
- 3498 [255] Z.-B. Kang, K. Lee, D. Y. Shao, and J. Terry, The sivers asymmetry in hadronic dijet
3499 production, 2020, 2008.05470.
- 3500 [256] F. Yuan, Phys. Rev. Lett. **100**, 032003 (2008).
- 3501 [257] F. Yuan, Phys. Rev. D **77**, 074019 (2008).
- 3502 [258] STAR, L. Adamczyk *et al.*, Phys. Rev. D **97**, 032004 (2018), 1708.07080.
- 3503 [259] M. Anselmino *et al.*, Phys. Rev. D **73**, 014020 (2006), hep-ph/0509035.
- 3504 [260] M. Anselmino *et al.*, Phys. Rev. D **75**, 054032 (2007).
- 3505 [261] J. Soffer, Phys. Rev. Lett. **74**, 1292 (1995).
- 3506 [262] U. D'Alesio, F. Murgia, and C. Pisano, Phys. Lett. B **773**, 300–306 (2017).
- 3507 [263] H1, ZEUS, F. Aaron *et al.*, Eur. Phys. J. C **72**, 2175 (2012), 1207.4864.
- 3508 [264] CMS, V. Khachatryan *et al.*, Phys. Rev. D **92**, 012003 (2015), 1503.08689.
- 3509 [265] S. L. Bultmann *et al.*, Phys. Lett. B **579**, 245 (2004), nucl-ex/0305012.
- 3510 [266] S. Bultmann *et al.*, Phys. Lett. B **632**, 167 (2006), nucl-ex/0507030.
- 3511 [267] pp2pp, S. Bultmann *et al.*, Phys. Lett. B **647**, 98 (2007), nucl-ex/0610022.
- 3512 [268] STAR, L. Adamczyk *et al.*, Phys. Lett. B **719**, 62 (2013), 1206.1928.
- 3513 [269] SN0605: A polarized p+p and p+A program for the next years [https://drupal.](https://drupal.star.bnl.gov/STAR/starnotes/public/sn0605)
3514 [star.bnl.gov/STAR/starnotes/public/sn0605](https://drupal.star.bnl.gov/STAR/starnotes/public/sn0605).
- 3515 [270] I. Helenius, in *Talk at 7th International Work-*
3516 *shop on Multiple Partonic Interactions at the LHC*,
3517 <http://indico.ictp.it/event/a14280/session/266/contribution/1042/material/slides/0.pdf>.
- 3518 [271] STAR, C. Dilks, PoS **DIS2016**, 212 (2016), 1805.08875.
- 3519 [272] A. Airapetian *et al.*, Phys. Lett. B **693**, 11–16 (2010).
- 3520 [273] C. Adolph *et al.*, Phys. Lett. B **744**, 250–259 (2015).

- 3521 [274] D. Müller, D. Robaschik, B. Geyer, F.-M. Dittes, and J. Hořejší, Fortsch. Phys. **42**,
3522 101 (1994), hep-ph/9812448.
- 3523 [275] X.-D. Ji, Phys. Rev. Lett. **78**, 610 (1997), hep-ph/9603249.
- 3524 [276] A. Radyushkin, Phys. Lett. B **380**, 417 (1996), hep-ph/9604317.
- 3525 [277] M. Burkardt, Phys. Rev. D **62**, 071503 (2000), hep-ph/0005108, [Erratum: Phys.Rev.D
3526 66, 119903 (2002)].
- 3527 [278] S. Klein and J. Nystrand, Photoproduction of J/ψ and Upsilon in pp and anti-p p
3528 collisions, in *5th Workshop on Small x and Diffractive Physics*, 2003, hep-ph/0310223.
- 3529 [279] S. R. Klein, J. Nystrand, J. Seger, Y. Gorbunov, and J. Butterworth, Comput. Phys.
3530 Commun. **212**, 258 (2017), 1607.03838.
- 3531 [280] J. Lansberg, L. Massacrier, L. Szymanowski, and J. Wagner, Phys. Lett. B **793**, 33
3532 (2019), 1812.04553.
- 3533 [281] HERMES, A. Airapetian *et al.*, Phys. Lett. B **577**, 37 (2003), hep-ex/0307023.
- 3534 [282] HERMES, A. Airapetian *et al.*, Nucl. Phys. B **780**, 1 (2007), 0704.3270.
- 3535 [283] HERMES, A. Airapetian *et al.*, Phys. Lett. B **684**, 114 (2010), 0906.2478.
- 3536 [284] W. Brooks and H. Hakobyan, Nucl. Phys. A **830**, 361C (2009), 0907.4606.
- 3537 [285] NuSea, M. Vasilev *et al.*, Phys. Rev. Lett. **83**, 2304 (1999), hep-ex/9906010.
- 3538 [286] DIS-2014 [https://indico.cern.ch/event/258017/contributions/1588381/
3539 attachments/453934/629277/pA_DIS2014_Paukkunen.pdf](https://indico.cern.ch/event/258017/contributions/1588381/attachments/453934/629277/pA_DIS2014_Paukkunen.pdf).
- 3540 [287] D. de Florian, R. Sassot, P. Zurita, and M. Stratmann, Phys. Rev. D **85**, 074028
3541 (2012), 1112.6324.
- 3542 [288] PHENIX, S. Adler *et al.*, Phys. Rev. Lett. **98**, 172302 (2007), nucl-ex/0610036.
- 3543 [289] K. Eskola, H. Paukkunen, and C. Salgado, Nucl. Phys. A **855**, 150 (2011), 1011.6534.
- 3544 [290] R. Sassot, M. Stratmann, and P. Zurita, Phys. Rev. D **81**, 054001 (2010), 0912.1311.
- 3545 [291] K. J. Eskola, P. Paakkinen, H. Paukkunen, and C. A. Salgado, Eur. Phys. J. C **77**,
3546 163 (2017), 1612.05741.
- 3547 [292] N. Armesto, H. Paukkunen, J. M. Penín, C. A. Salgado, and P. Zurita, Eur. Phys. J.
3548 C **76**, 218 (2016), 1512.01528.
- 3549 [293] H. Paukkunen and P. Zurita, JHEP **12**, 100 (2014), 1402.6623.

- 3550 [294] E. Aschenauer *et al.*, (2014), 1409.1633.
- 3551 [295] H. Paukkunen, K. J. Eskola, and C. Salgado, Nucl. Phys. A **931**, 331 (2014), 1408.4563.
- 3552 [296] K. J. Eskola, H. Paukkunen, and C. A. Salgado, JHEP **10**, 213 (2013), 1308.6733.
- 3553 [297] L. Gribov, E. Levin, and M. Ryskin, Phys. Rept. **100**, 1 (1983).
- 3554 [298] E. Iancu and R. Venugopalan, The Color glass condensate and high-energy scattering
3555 in QCD, in *In *Hwa, R.C. (ed.) et al.: Quark gluon plasma* 249-3363*, 2003, hep-
3556 ph/0303204.
- 3557 [299] H. Weigert, Prog. Part. Nucl. Phys. **55**, 461 (2005), hep-ph/0501087.
- 3558 [300] J. Jalilian-Marian and Y. V. Kovchegov, Prog. Part. Nucl. Phys. **56**, 104 (2006),
3559 hep-ph/0505052.
- 3560 [301] F. Gelis, E. Iancu, J. Jalilian-Marian, and R. Venugopalan, Ann. Rev. Nucl. Part. Sci.
3561 **60**, 463 (2010), 1002.0333.
- 3562 [302] G. Giuliani, H. Zheng, and A. Bonasera, Prog. Part. Nucl. Phys. **76**, 116 (2014),
3563 1311.1811.
- 3564 [303] Y. V. Kovchegov and E. Levin *Quantum chromodynamics at high energy* Vol. 33 (Cam-
3565 bridge University Press, 2012).
- 3566 [304] A. H. Mueller and J.-w. Qiu, Nucl. Phys. B **268**, 427 (1986).
- 3567 [305] L. McLerran and R. Venugopalan, Phys. Rev. D **49**, 2233 (1994).
- 3568 [306] L. McLerran and R. Venugopalan, Phys. Rev. D **49**, 3352 (1994).
- 3569 [307] L. McLerran and R. Venugopalan, Phys. Rev. D **50**, 2225 (1994).
- 3570 [308] Y. V. Kovchegov, Phys. Rev. D **54**, 5463 (1996).
- 3571 [309] Y. V. Kovchegov, Phys. Rev. D **55**, 5445 (1997).
- 3572 [310] J. Jalilian-Marian, A. Kovner, L. McLerran, and H. Weigert, Phys. Rev. D **55**, 5414
3573 (1997).
- 3574 [311] A. H. Mueller, Nucl. Phys. **B415**, 373 (1994).
- 3575 [312] A. H. Mueller and B. Patel, Nucl. Phys. B **425**, 471 (1994), hep-ph/9403256.
- 3576 [313] I. Balitsky, Nucl. Phys. **B463**, 99 (1996), hep-ph/9509348.
- 3577 [314] I. Balitsky, Phys. Rev. D **60**, 014020 (1999), hep-ph/9812311.

- 3578 [315] Y. V. Kovchegov, Phys. Rev. **D60**, 034008 (1999), hep-ph/9901281.
- 3579 [316] Y. V. Kovchegov, Phys. Rev. D **61**, 074018 (2000), hep-ph/9905214.
- 3580 [317] J. Jalilian-Marian, A. Kovner, and H. Weigert, Phys. Rev. D **59**, 014015 (1998),
3581 hep-ph/9709432.
- 3582 [318] J. Jalilian-Marian, A. Kovner, A. Leonidov, and H. Weigert, Phys. Rev. D **59**, 014014
3583 (1998), hep-ph/9706377.
- 3584 [319] E. Iancu, A. Leonidov, and L. D. McLerran, Phys. Lett. B **510**, 133 (2001), hep-
3585 ph/0102009.
- 3586 [320] E. Iancu, A. Leonidov, and L. D. McLerran, Nucl. Phys. **A692**, 583 (2001), hep-
3587 ph/0011241.
- 3588 [321] A. Accardi *et al.*, Eur. Phys. J. A **52**, 268 (2016), 1212.1701.
- 3589 [322] Y. V. Kovchegov and M. D. Sievert, Nucl. Phys. B **903**, 164 (2016), 1505.01176.
- 3590 [323] CMS, S. Chatrchyan *et al.*, Eur. Phys. J. C **74**, 2951 (2014), 1401.4433.
- 3591 [324] STAR, E. Braidot, Nucl. Phys. A **854**, 168 (2011), 1008.3989.
- 3592 [325] PHENIX, A. Adare *et al.*, Phys. Rev. Lett. **107**, 172301 (2011), 1105.5112.
- 3593 [326] C. Marquet, Nucl. Phys. A **796**, 41 (2007), 0708.0231.
- 3594 [327] J. L. Albacete and C. Marquet, Phys. Rev. Lett. **105**, 162301 (2010), 1005.4065.
- 3595 [328] Z.-B. Kang, I. Vitev, and H. Xing, Phys. Rev. D **85**, 054024 (2012), 1112.6021.
- 3596 [329] M. Strikman and W. Vogelsang, Phys. Rev. D **83**, 034029 (2011), 1009.6123.
- 3597 [330] J. Jalilian-Marian and A. H. Rezaeian, Phys. Rev. D **86**, 034016 (2012), 1204.1319.
- 3598 [331] J. L. Albacete and C. Marquet, Nucl. Phys. A **854**, 154 (2011), 1009.3215.
- 3599 [332] K. J. Eskola, H. Paukkunen, and C. A. Salgado, JHEP **07**, 102 (2008), 0802.0139.
- 3600 [333] A. H. Rezaeian, Phys. Rev. D **86**, 094016 (2012), 1209.0478.
- 3601 [334] T. Sjostrand, S. Mrenna, and P. Z. Skands, Comput. Phys. Commun. **178**, 852 (2008),
3602 0710.3820.
- 3603 [335] Di-jet production from pythia8.189 is scaled down due to its overestimation of inclusive
3604 π_0 yields compared to those reported by BRAHMS in phys. rev. lett. 98 (2007) 252001
3605 and STAR in phys. rev. lett. 97 (2006) 152302.

- 3606 [336] T. Kaufmann, A. Mukherjee, and W. Vogelsang, Phys. Rev. D **92**, 054015 (2015),
3607 1506.01415, [Erratum: Phys.Rev.D 101, 079901 (2020)].
- 3608 [337] D. de Florian, R. Sassot, M. Epele, R. J. Hernández-Pinto, and M. Stratmann, Phys.
3609 Rev. D **91**, 014035 (2015), 1410.6027.
- 3610 [338] A. Khouaja *et al.*, Nucl. Phys. A **780**, 1 (2006).
- 3611 [339] D. de Florian and R. Sassot, Phys. Rev. D **69**, 074028 (2004), hep-ph/0311227.
- 3612 [340] D. de Florian, R. Sassot, and M. Stratmann, Phys. Rev. D **75**, 114010 (2007), hep-
3613 ph/0703242.
- 3614 [341] D. de Florian, R. Sassot, and M. Stratmann, Phys. Rev. D **76**, 074033 (2007),
3615 0707.1506.
- 3616 [342] V. Smakhtin *et al.*, Nucl. Instrum. Meth. **A598**, 196 (2009).
- 3617 [343] A. Abusleme *et al.*, Nucl. Instrum. Meth. **A817**, 85 (2016), 1509.06329.
- 3618 [344] Iakovidis, George, EPJ Web Conf. **174**, 07001 (2018).